

# Use of various seeding light sources for seeding FEL

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# Motivation for seeding

seeding for performing the energy exchange :

=> bunching on the undulator fundamental and harmonics

Low gain case : coherent harmonic generation

High gain case :

- quicker saturation => compactness and cost
- improved stability (intensity, wavelength, jitter) => pump-probe user applications
- «control» of temporal coherence properties : removal of spikes

# Required properties of a seeding source

«Ideally» :

- **wide spectral range** : tuneability, short wavelength and ultimately handling of «gap scan» (FFDW orbit correction, synchronised change of seed wavelength, undulator gap and monochromator setting)

- **Polarised** for proper energy exchange in the undulator / **adjustable polarisation**

- **high intensity** : power overcoming shot noise

*J. Wu et al, Appl. Phys. Lett. 90 (2007) 488*

*L. Giannessi et al, FEL04 (2004) 37*

cf strategies to reduce shot noise

*A. Gover et al. PRL 102, 154801 (2009), A. Nause et al. J. Appl. Phys. 107, 103101 (2010)*

*V. Linvinenko, Proceeding FEL 09, Liverpool, 2009*

*D. Ratner et al., PRL 14, 060710 (2011)*

- **spectral bandwidth** - adapted to the gain bandwidth?

- **adjustable pulse duration** / Chirp handling?

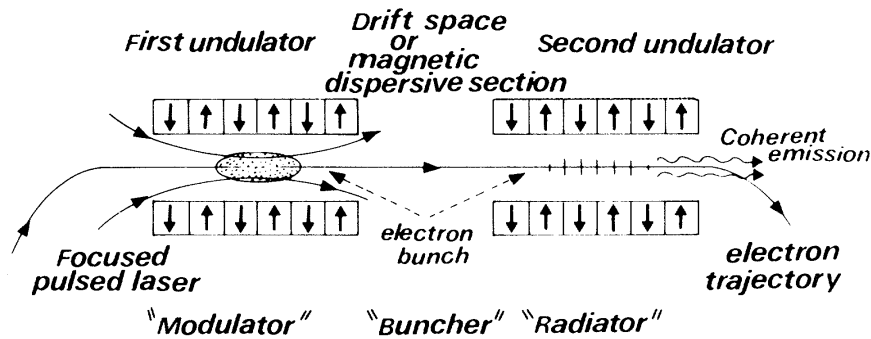
- **repetition rate** adapted to the the accelerator's one

# Required diagnostics of a seeding source

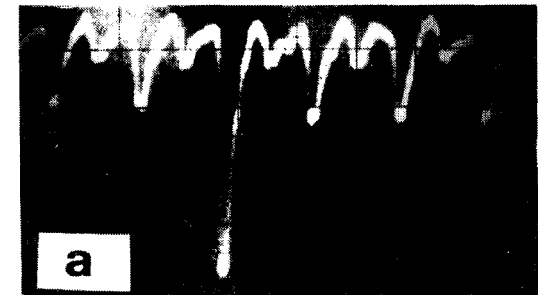
- power level monitoring
- spectral tuning between the seed and the undulator radiation :  
=> monochromator
- temporal synchronisation with the electron beam  
=> photodiode, streak-camera... with a radiation output of the electron beam (either undulator spontaneous emission, or transition radiation in the chicane/ dogleg of injection...)
- proper focus in the undulator  
=> telescope
- transverse overlap with the electron beam  
=> profile monitor (CCD, Yag screen, MCP...)

## First coherent harmonic generation results

ACO (Orsay, France), 166 MeV  
 Nd-Yag at 1.06  $\mu\text{m}$ , 20 Hz, 15 MW, 12 ns  
 => CHG at 352 nm



352 nm  
 $R_3 = 3$



352 nm  
 $R_3 > 100$



100 ns

Linewidth sharpening

B. Girard, Y. Lapierre, J. M. Ortéga, C. Bazin, M. Billardon, P. Elleaume, M. Bergher, M. Velghe, Y. Petroff, Optical frequency multiplication by an optical klystron, *Phys. Rev. Lett.* 53 (25) 2405-2408 (1984)

ACO (Orsay, France), 220 MeV  
 Nd-Yag at 1.06  $\mu\text{m}$  (36 MW) /2  
 = 532 nm, => CHG at 177 and 106.4 nm

Coherent Harmonic Generation in the Vacuum Ultra Violet Spectral Range on the Storage Ring ACO, R. Prazeres, J.M. Ortéga, M. Billardon, C. Bazin, M. Bergher, M.E. Couprie, H. Fang, M. Velghe Y. Petroff, *Nuclear Inst. and Methods in Physics Research A272* (1988), 68-72

Experimental results on ACO (1987)

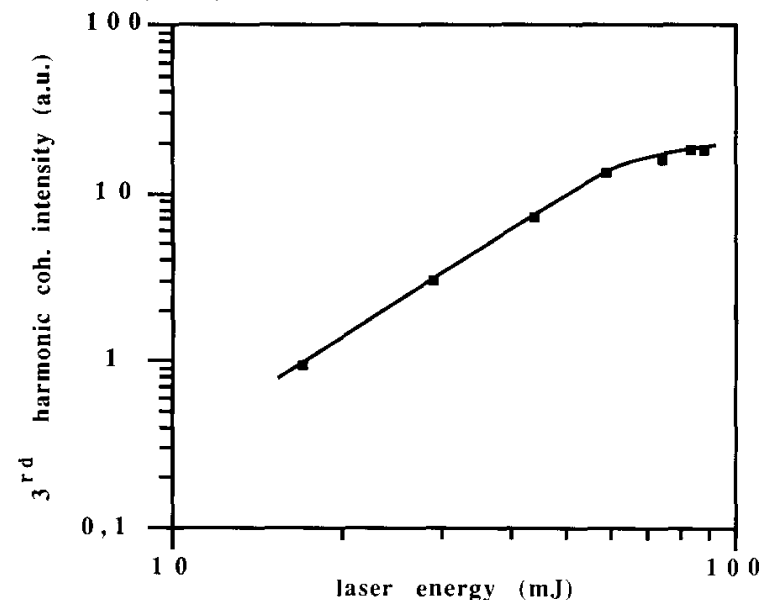
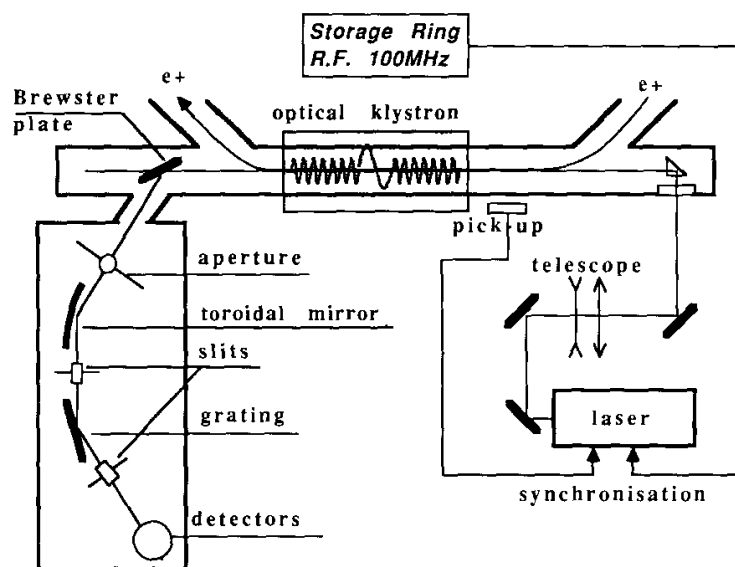
Observed harmonic	3	5
Corresponding wavelength [Å]	1773	1064
Integrated ratio $R_n^{\text{int}}$	350	3-4
monochromator bandwidth [Å]	2	2
monochromator angular aperture [mrad <sup>2</sup> ]	1.4	3
Spectral ratio $R_n(\lambda, \Omega)$	6000	100
Number of coherent photons/pulse	$1.5 \times 10^7$	$10^5$
in spectral width [Å]	0.1	0.07
in angular aperture [mrad]	0.2	0.1

# First coherent harmonic generation results

Super-ACO (Orsay, France), 700 MeV

Nd-Yag at  $1.06 \mu\text{m} / 2 : 100 \text{ mJ}, 300 \text{ ps}, 300 \text{ MW}, 10 \text{ Hz} \Rightarrow \text{CHG at } 177 \text{ nm and } 106 \text{ nm}$

Coherent Harmonic Generation in VUV with the optical klystron on the storage ring Super-ACO, R. Prazeres, P. Guyot-Sionnest, D. Jaroszynski, J.M. Ortéga, M. Billardon, M. E. Couprie, M. Velghe Nucl.Inst. Meth. A 304 (1991) 72-76



## Positron beam

Current per bunch, $i$ [mA]	6.7	6.7	22
Energy spread, $\sigma_\gamma/\gamma$	$5.6 \times 10^{-4}$	$5.6 \times 10^{-4}$	$9.5 \times 10^{-4}$
Bunch length, $\sigma_z$ [ps]	100	100	160

## Monochromator

Resolution, $\delta\lambda$ [Å]	12	2	4
Angular aperture, $\delta\theta$ [mrad]	0.3	0.3	1

## Integrated ratio

Measured, $R_n^{\text{int}}$	270	1400	25
Theoretical, $R_n^{\text{int}}$	3200	20000	260

## Spectral ratio, $R_n^{\text{spe}}(\lambda, \theta)$

	$5 \times 10^5$	$5 \times 10^5$	1200
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## Coherent photons, $N_{\text{coh}}$

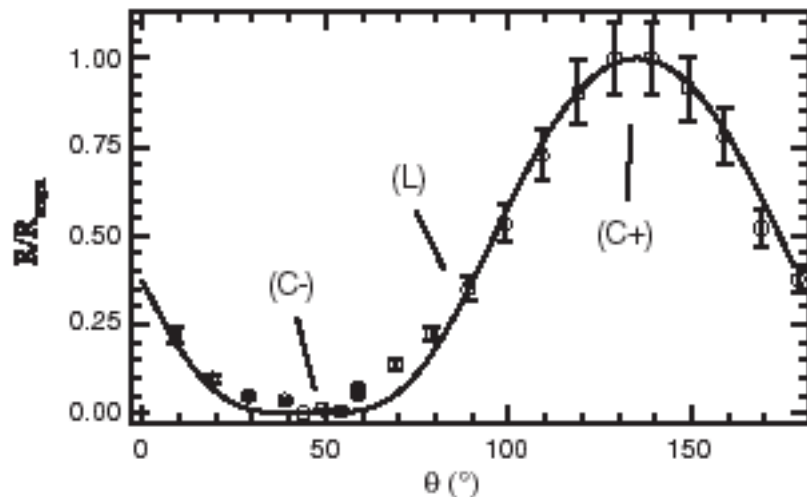
	$4 \times 10^8$	$4 \times 10^8$	$4 \times 10^7$
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ie, italy, 10-12 Dec. 2012

# Coherent harmonic generation : ex of use of circular polarisation

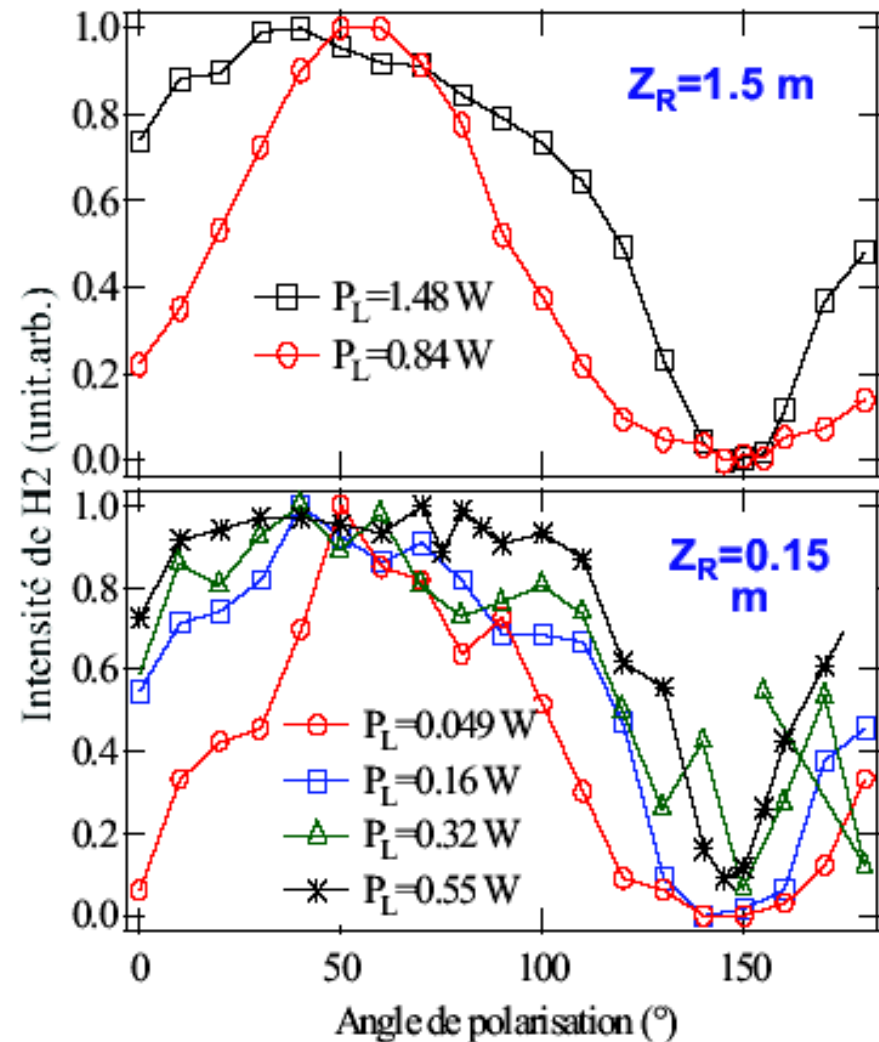
UVSOR-II, helical optical klystron, change of the laser polarisation

H2 intensity ratio



laser helicity

enhancement of laser power  
(focusing, intensity)  
=> broadening due to  
overmodulation



M. Labat, M. Hosaka, M. Shimada, M. Katoh, M.E. Couprie, "Optimization of a seeded Free-Electron Laser with helical undulators", *Phys. Rev. Lett.* 101, 164803 (2008).

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

## HGHG at BNL

L. H. Yu et al, Science 289, 2000, 932

L. H. Yu et al, PRL912003, 074801

• **Temporal Coherence** properties given by the seed (external laser) => High degree of temporal coherence

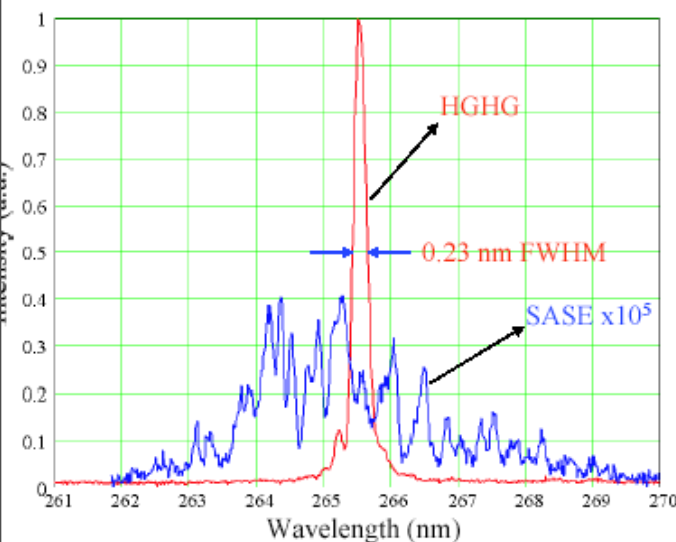
**Stability** (reduction of intensity fluctuations and jitter ) for user applications (pump probe)

**Quicker Saturation**

=>

Compactness and cost

**DUVFEL**  
(210 MeV, 30 mJ- 800 nm Ti:Sa laser)

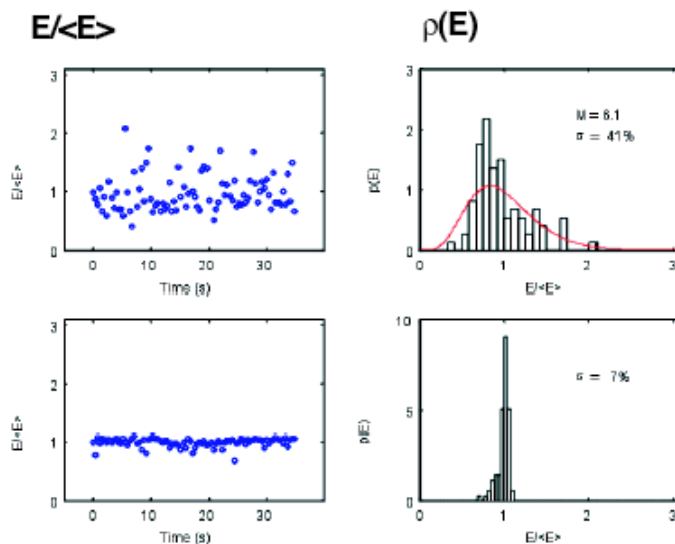


Spectral narrowing

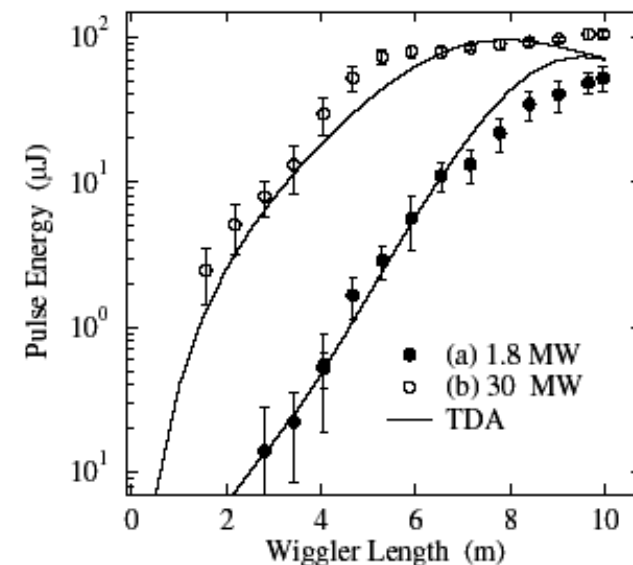
Non Linear Harmonics

Shot to Shot Intensity Fluctuation

Shows High Stability of HGHG output



Courtesy L. H. Yu

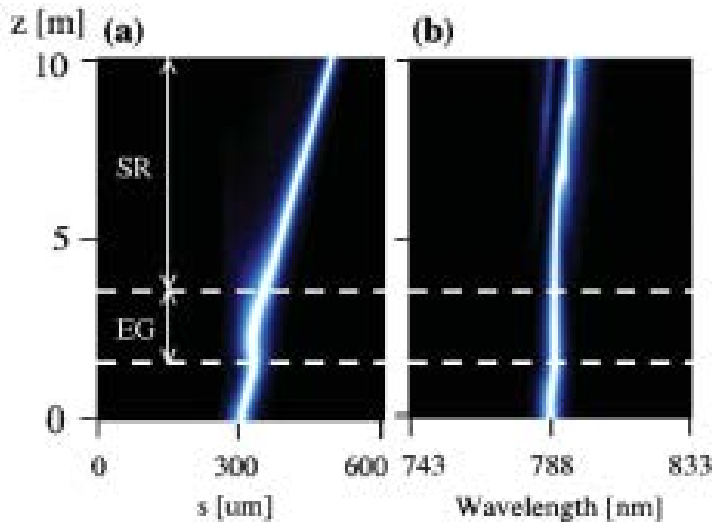


2. Pulse energy versus distance in the radiator for two values of the seed laser input power: (a) 1.8 MW and (b) 30 MW. The solid curves are simulation results by the TDA code.



## Regimes for seeding

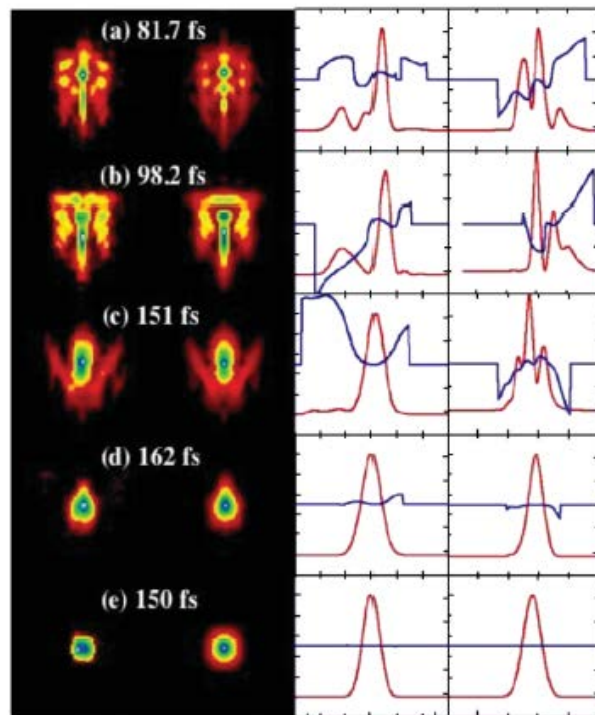
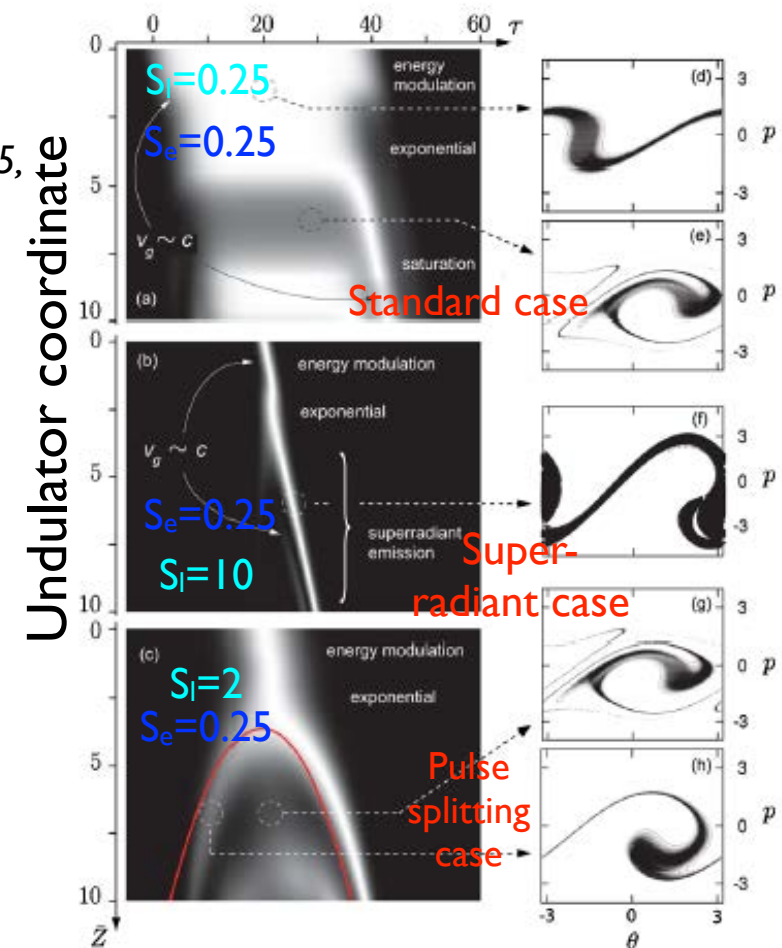
### Super-radiant mode



*L. Giannessi et al, Appl. Optics 2005, slippage length/potential depth*

### pulse splitting

### Undulator coordinate



*T. Watanabe et al. PRL 98, 034802 (2007)*

$$S_e = \frac{4\pi\rho N}{I_{..}}$$

$$S_{seed} = \frac{4\pi\rho N}{L_{seed}}$$

*M. Labat, N. Joly, S. Bielawski, C. Szwaj, C. Bruni, M.E. Couprie, Pulse splitting in short wavelength seeded FEL, Phys. Rev. Lett. 103, 264801 (2009)*

Seeding and self-seeding at new FEL sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

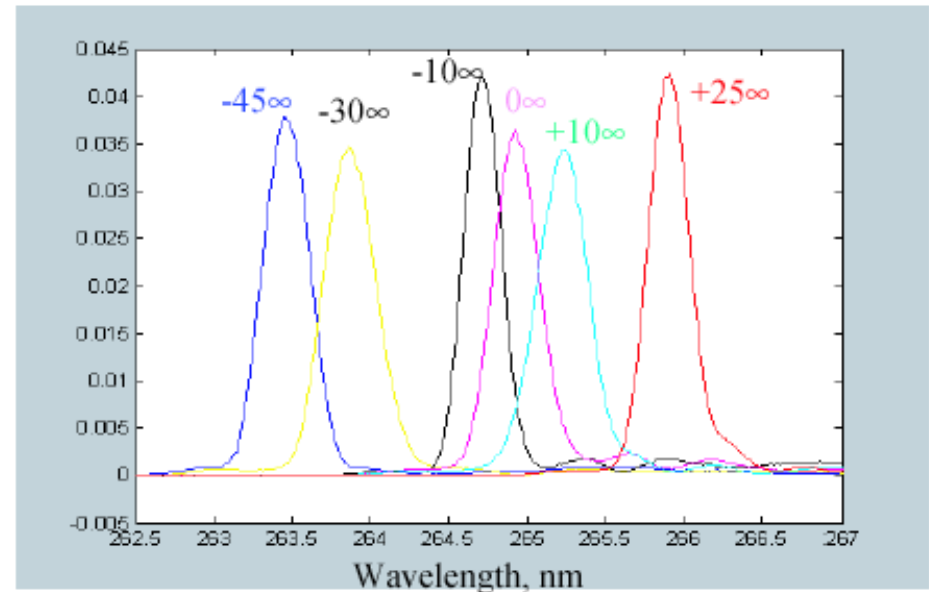
# Making use of chirp for tuneability

- Tuneability of the seed
- Combination of the chirp on the electron beam and the laser

*T. Shaftan, BNL workshop, 04*

*S. Biedron et al., NIM A475, 401; T. Shaftan, PRE 71, 2005, 046501*

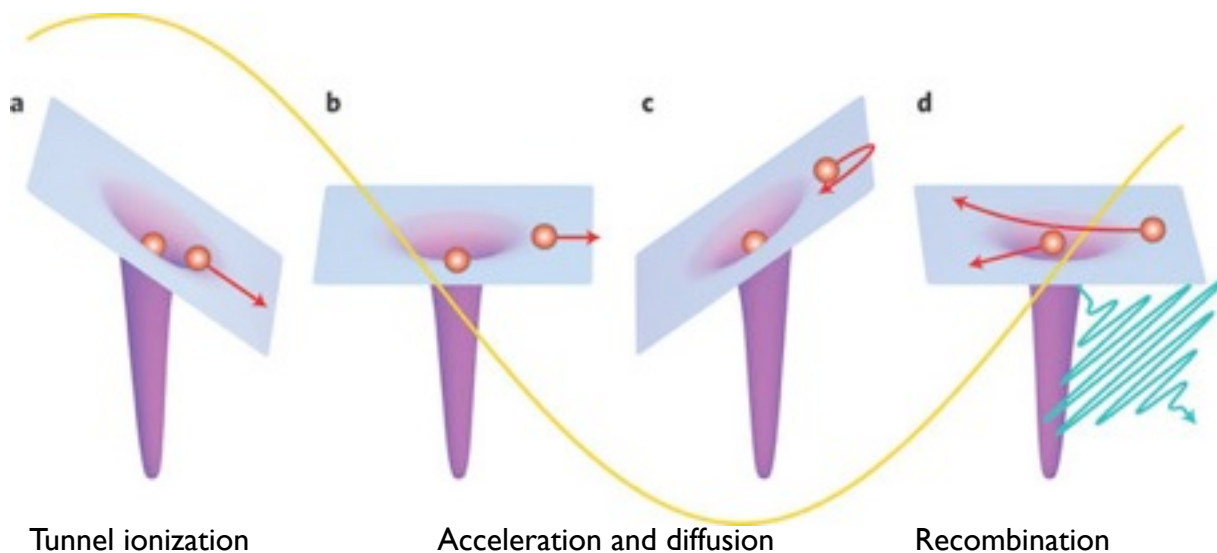
HGHG intensity, a.u.



# High Harmonic in Gas source

Mc Pherson et al., *JOSA B* 4, 595 (1987) (Chicago)

M. Ferray et al., *J. Phys. B* 21, L31 (1988) (Saclay)



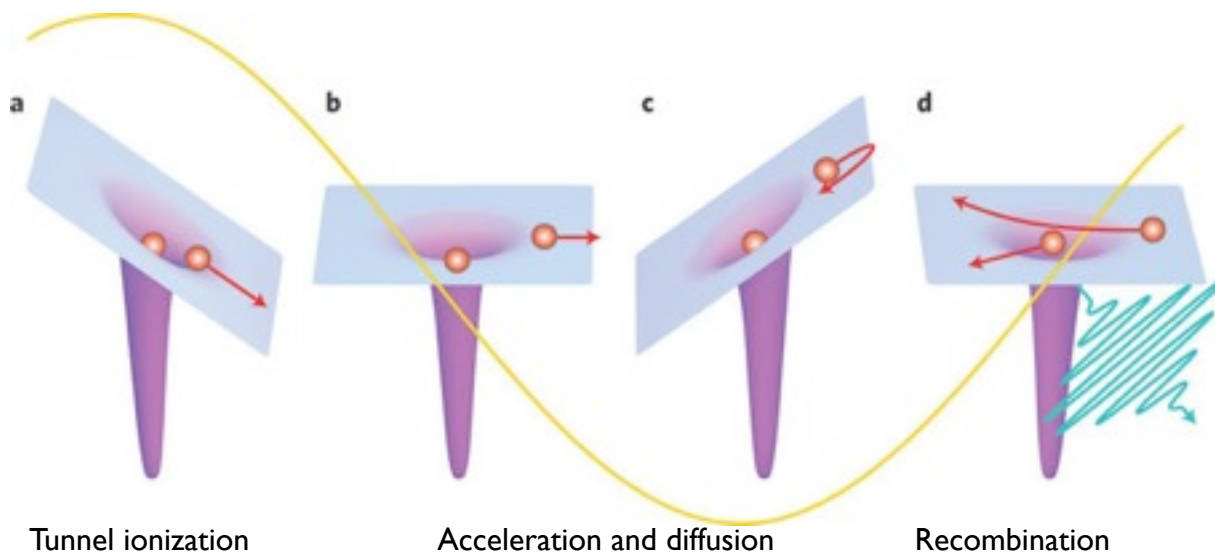
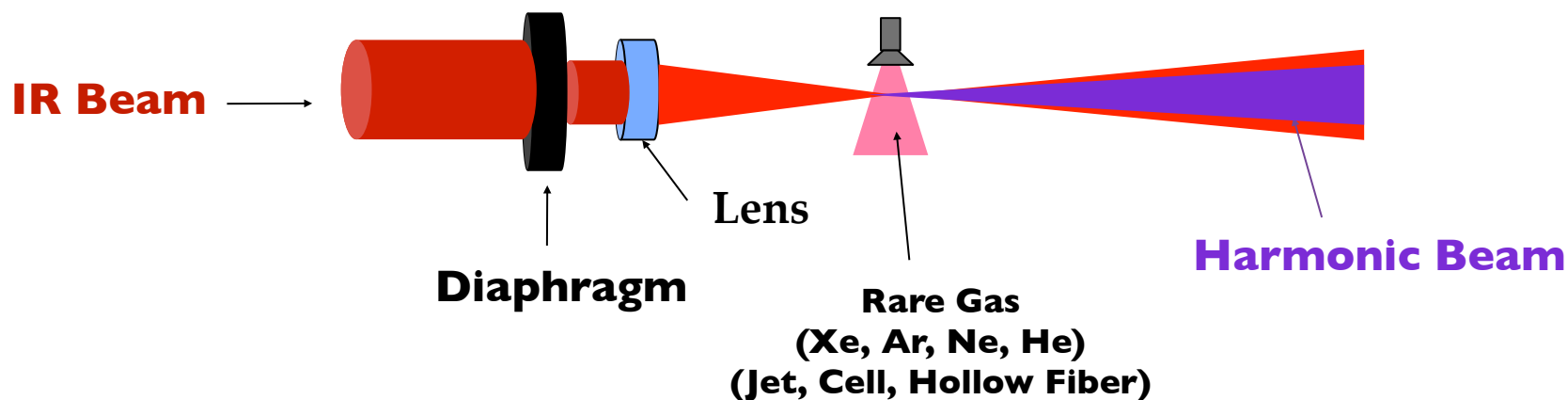
K. Midorikawa, *Ultrafast dynamic imaging, Nature Photonics* 5, 640–641 (2011)

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

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Mc Pherson et al., *JOSA B* 4, 595 (1987) (Chicago)

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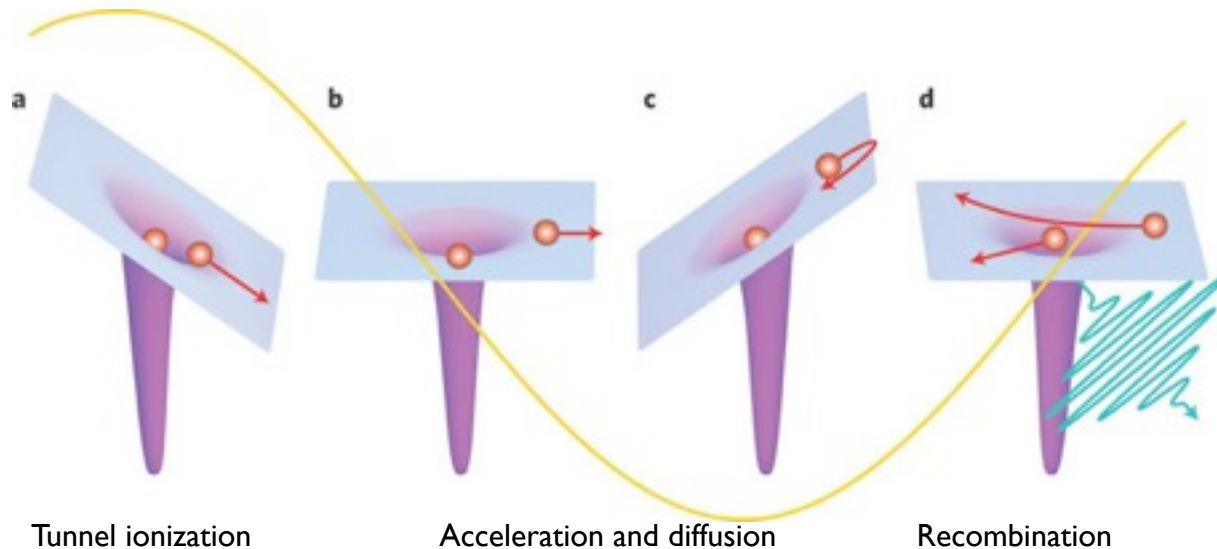
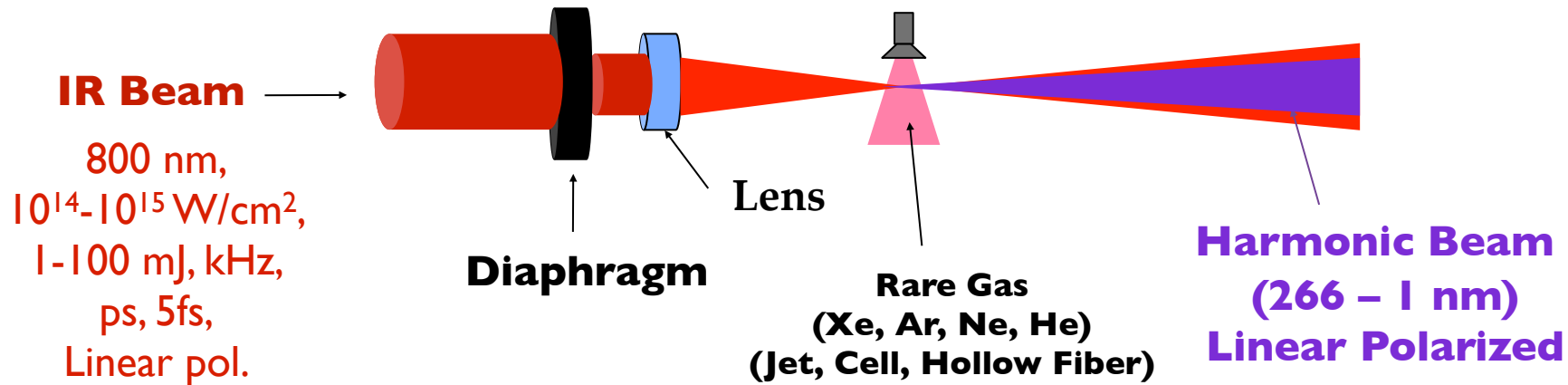


K. Midorikawa, *Ultrafast dynamic imaging*, *Nature Photonics* 5, 640–641 (2011)

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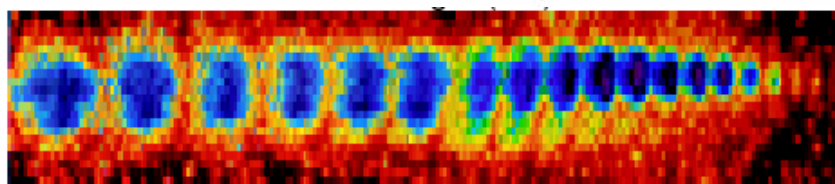
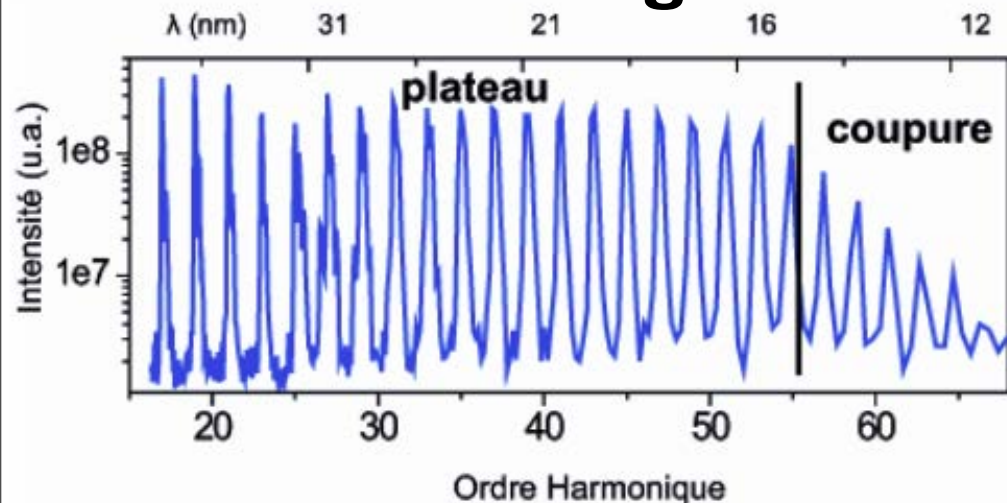
Mc Pherson et al., *JOSA B* 4, 595 (1987) (Chicago)

M. Ferray et al., *J. Phys. B* 21, L31 (1988) (Saclay)



K. Midorikawa, *Ultrafast dynamic imaging, Nature Photonics* 5, 640–641 (2011)

## High Harmonic in Gas source



Water window (Ti:Sa 26 fs, 2.7 nm (460 eV)  
He, 5.2 nm (239 eV) Ne)

## Tuneability of the HHG seed:

- by the laser tuneability
- by frequency mixing ~70% tunability from 180 to 18 nm  
with Ti:Sa + OPG *Garde et al., J. Phys B 29, L163 (1996)*
- by adjustment of the laser energy, chirp

•*Kim et al., PRA 67 (2003)*

Full tunability from 220 nm to 8 nm with 1.1 to 1.6 μm pump  
(Ti:Sa + OPA) *Chang et al., Phys. Rev. A 65 (2001)*

longer wavelength  
atoms (Ne  $I_p=21.6$  eV, He  
 $I_p=24.6$  eV), ions

$$h\nu_{\text{cutoff}} = I_p + \frac{0.5I_p^{3+} \lambda^2}{\left[ \ln \frac{0.86 \Delta t^{2n^*-1} G_{lm} C_{n^*}^2 I_p}{-\ln(1-p)} \right]^2}$$

short pulses (5-7 fs: approaching the  
one cycle limit)

## Reduction of the cut-off wavelength:

- ionisation potential: atoms (Ne  $I_p=21.6$  eV, He  
 $I_p=24.6$  eV), => ions *Milosevic et al. PRA63 (2000)*
- laser of long wavelength *Chang et al., PRA 65 (2001)*
- short pulses

*Z. Chang et al, PRL 79, 1997, 2967*

*Gibson, Science, 2003. Seres et al, Nature 2005*

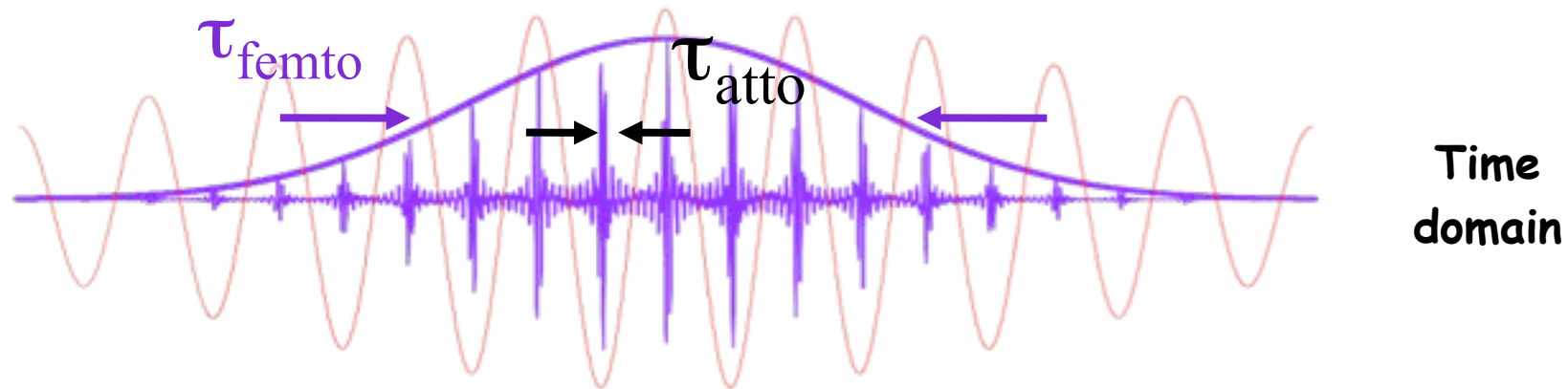
*B. Shang et al, PRA 65, 2001, 011804*

> 10 nJ/ pulse @ 4.37 nm

*M. Zepf et al, PRL 99, (2007), 143901*

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# High Harmonic in Gas source

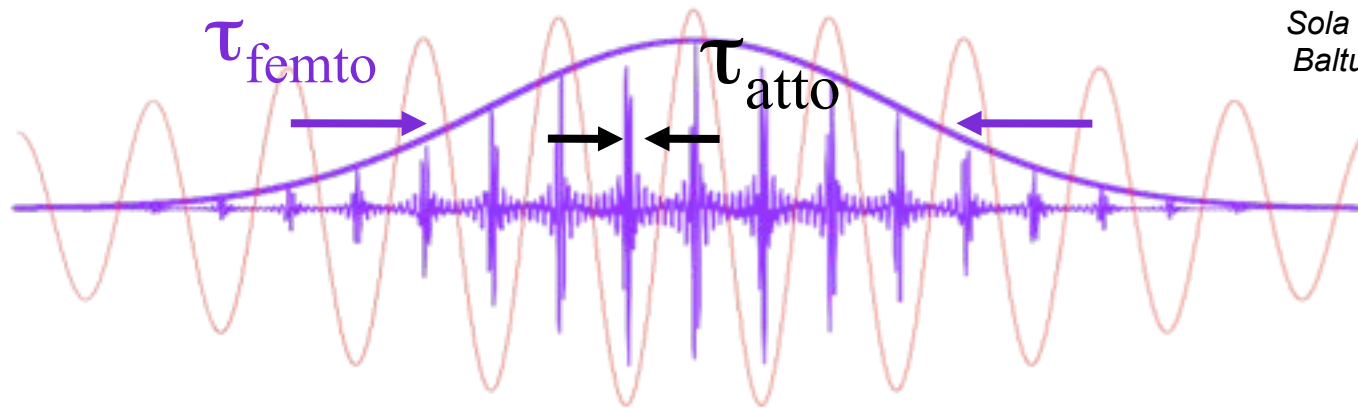


# High Harmonic in Gas source

Isolated atto pulse :

« Amplitude » or « Polarization » gating

*Sola et al., Nature Physics (2006)*  
*Baltuska et al. Nature 421 (2003)*



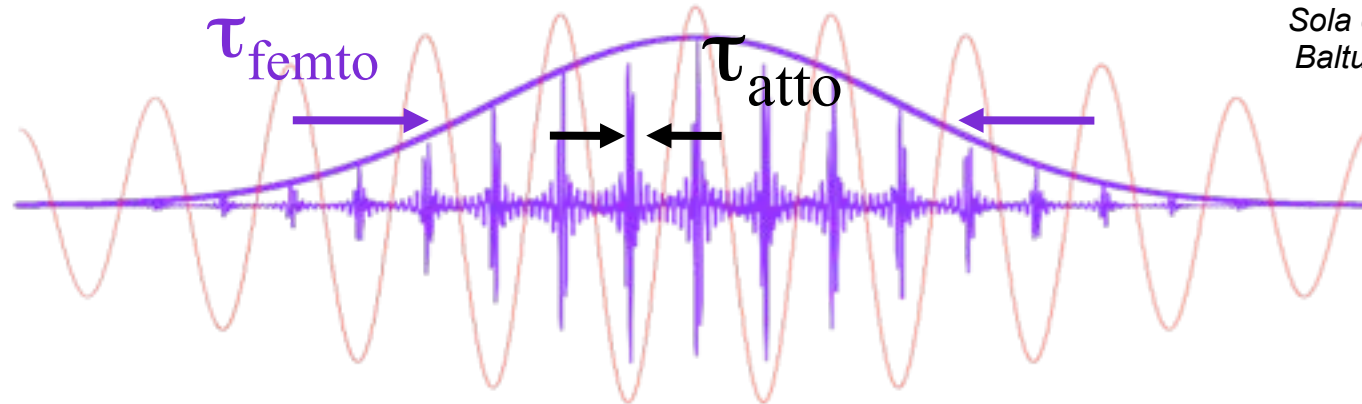
**Time  
domain**



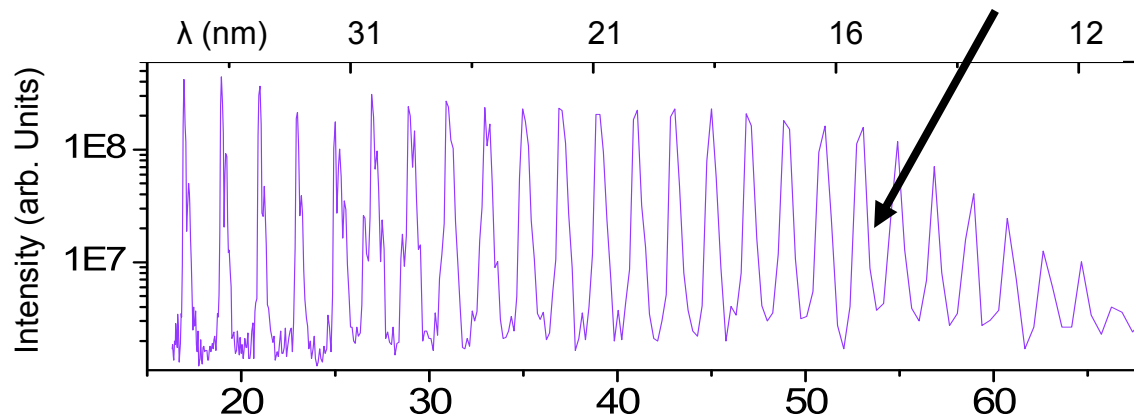
## High Harmonic in Gas source

Isolated atto pulse :

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*Sola et al., Nature Physics (2006)*  
*Baltuska et al. Nature 421 (2003)*Time  
domain

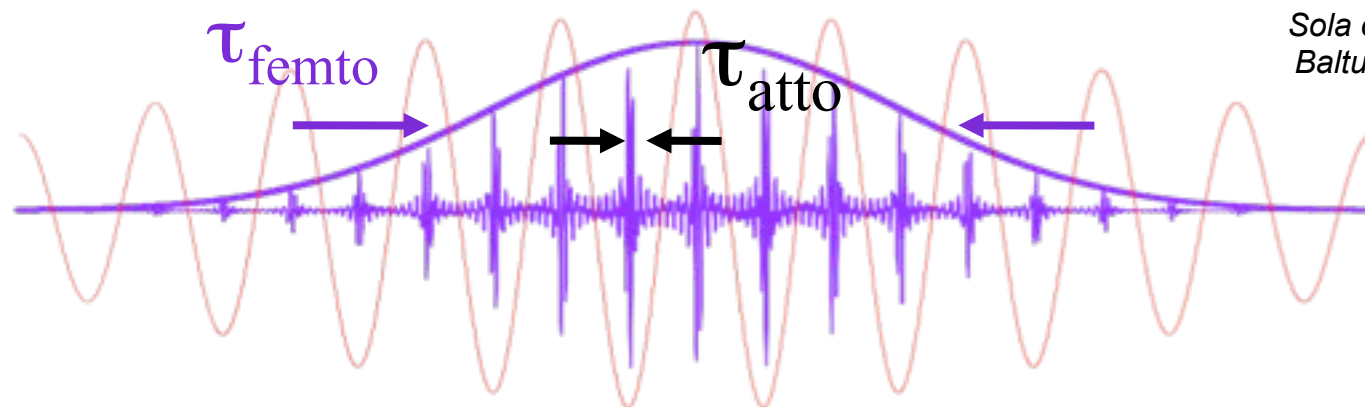
$$\Delta\nu_{\text{atto}} \sim 1/\tau_{\text{atto}}$$

Spectral  
domain

## High Harmonic in Gas source

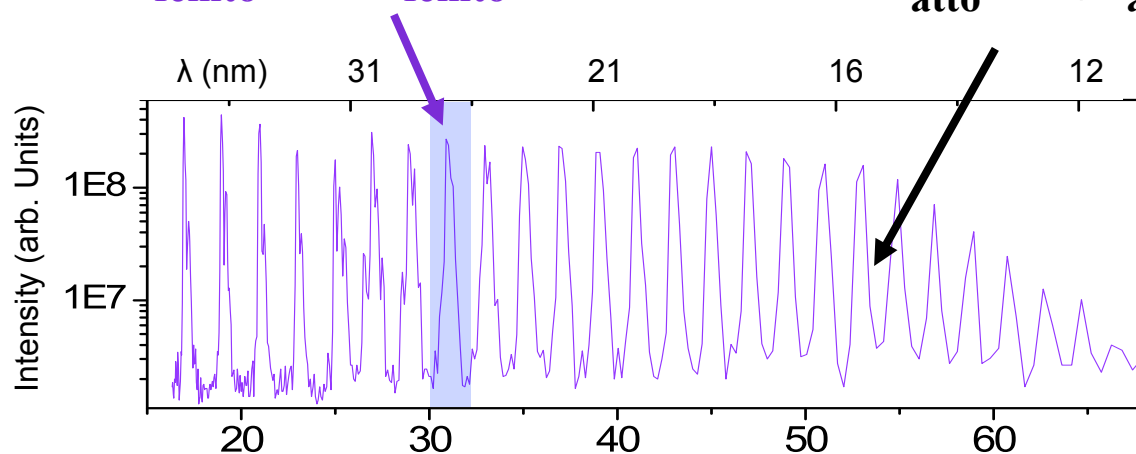
Isolated atto pulse :

« Amplitude » or « Polarization » gating

*Sola et al., Nature Physics (2006)*  
*Baltuska et al. Nature 421 (2003)*Time  
domain

$$\Delta\nu_{\text{femto}} \sim 1/\tau_{\text{femto}}$$

$$\Delta\nu_{\text{atto}} \sim 1/\tau_{\text{atto}}$$

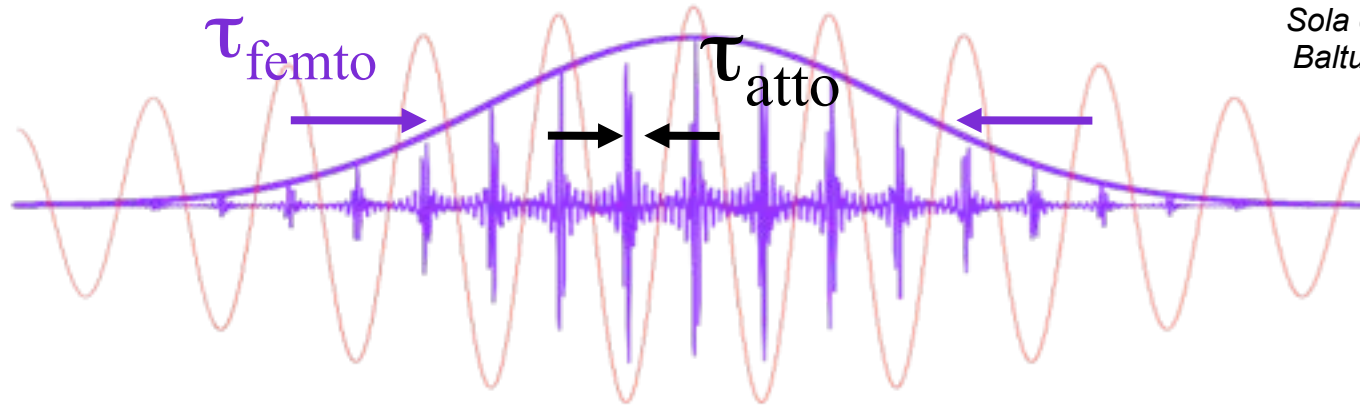
Spectral  
domain

## High Harmonic in Gas source

Isolated atto pulse :

« Amplitude » or « Polarization » gating

*Sola et al., Nature Physics (2006)*  
*Baltuska et al. Nature 421 (2003)*

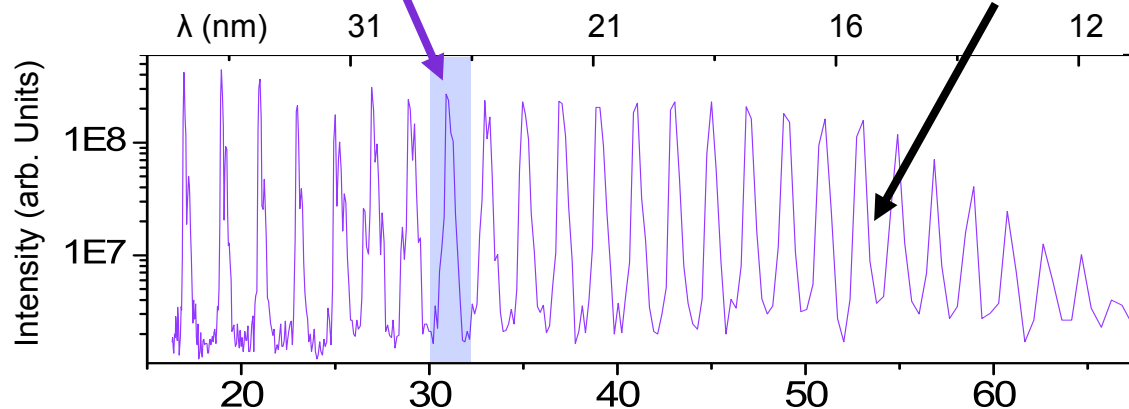


Time  
domain

$$\Delta\nu_{\text{femto}} \sim 1/\tau_{\text{femto}}$$

$$\Delta\nu_{\text{atto}} \sim 1/\tau_{\text{atto}}$$

Spectral  
domain



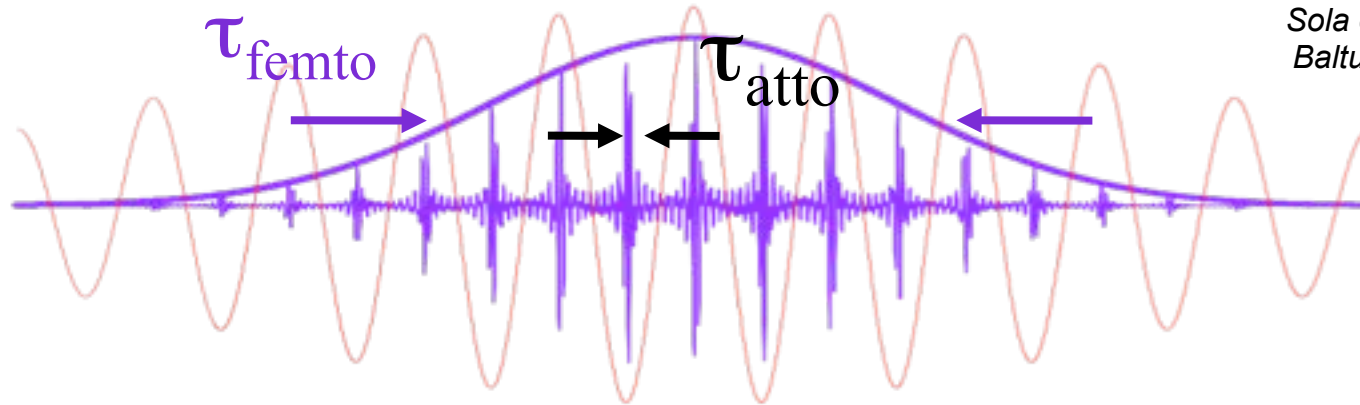
Harmonic train (atto pulse):  
linear phase

## High Harmonic in Gas source

Isolated atto pulse :

« Amplitude » or « Polarization » gating

*Sola et al., Nature Physics (2006)*  
*Baltuska et al. Nature 421 (2003)*

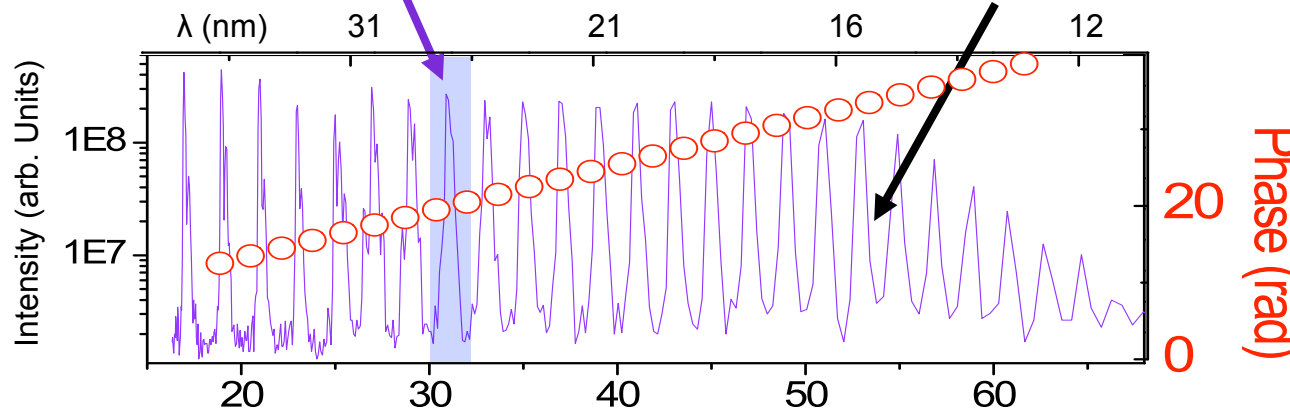


Time  
domain

$$\Delta\nu_{\text{femto}} \sim 1/\tau_{\text{femto}}$$

$$\Delta\nu_{\text{atto}} \sim 1/\tau_{\text{atto}}$$

Spectral  
domain



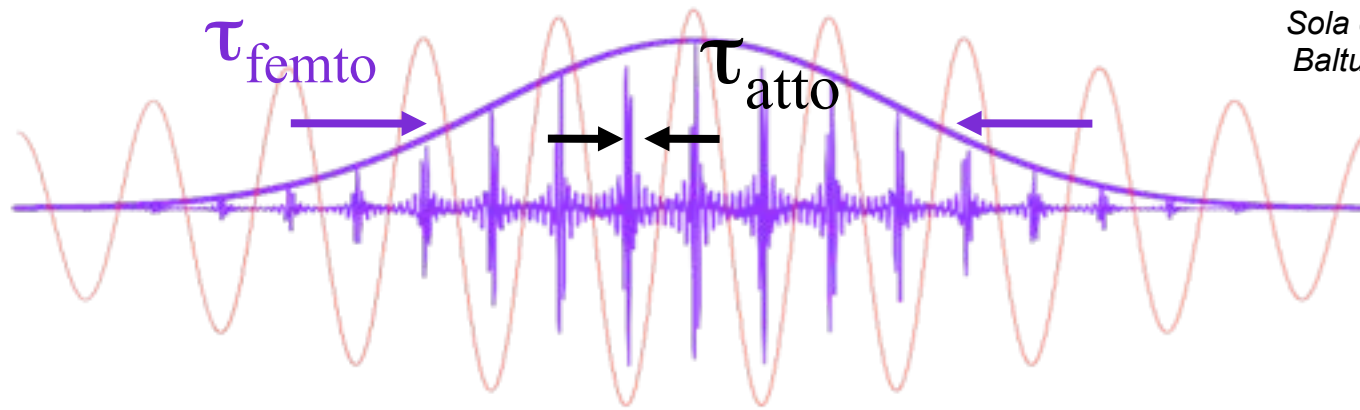
Harmonic train (atto pulse):  
linear phase

## High Harmonic in Gas source

Isolated atto pulse :

« Amplitude » or « Polarization » gating

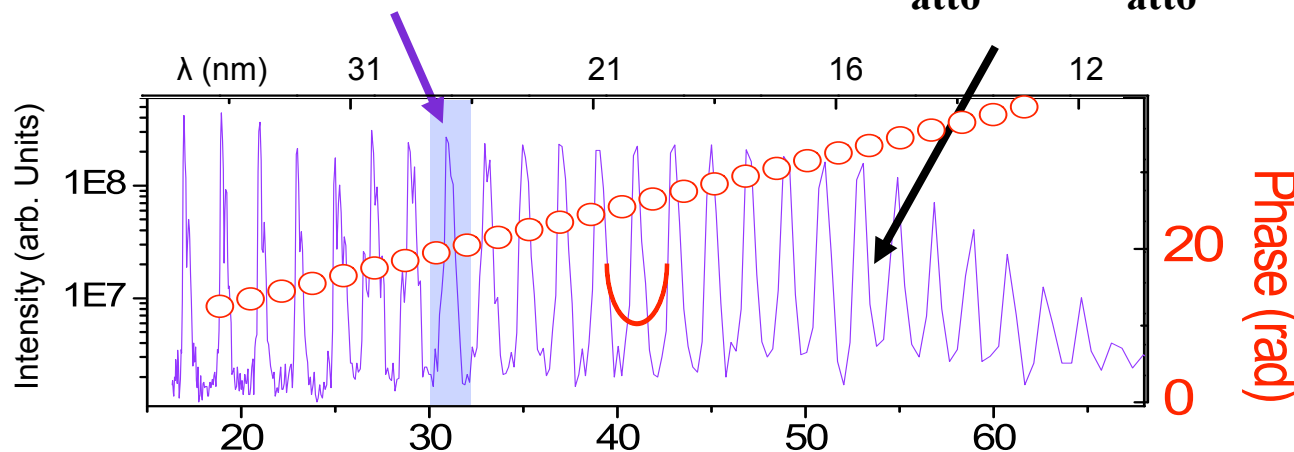
*Sola et al., Nature Physics (2006)*  
*Baltuska et al. Nature 421 (2003)*



**Time domain**

$$\Delta\nu_{\text{femto}} \sim 1/\tau_{\text{femto}}$$

$$\Delta\nu_{\text{atto}} \sim 1/\tau_{\text{atto}}$$



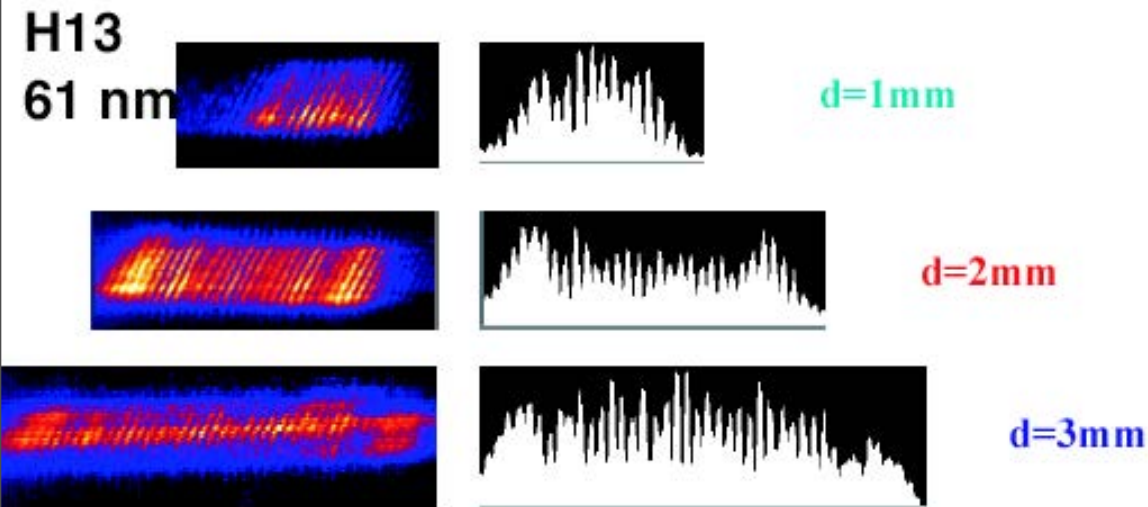
**Spectral domain**

Harmonic train (atto pulse):  
linear phase

Single harmonic (fs pulse):  
~ quadratic phase (can be compensated)

## High Harmonic in Gas source

Fresnel bi interferometer



Le Déroff et al., PRA61 (2000) 043802

Ne, 13.5 nm, 25 nJ, DV=0.35 mrad

He, 8.9 nm, 1 nJ, DV=0.3 mrad

E.J.Takahashi et al., Apply. Phys. Lett 84, 4 (2004)

Opt. Lett. 27, 2000, 1920

Phys. Rev A 68, 2003, 023808

Y.Tamaki et al, JJAP 40, 2001, L1154

Y. Mairesse et al, Science 302, 1540

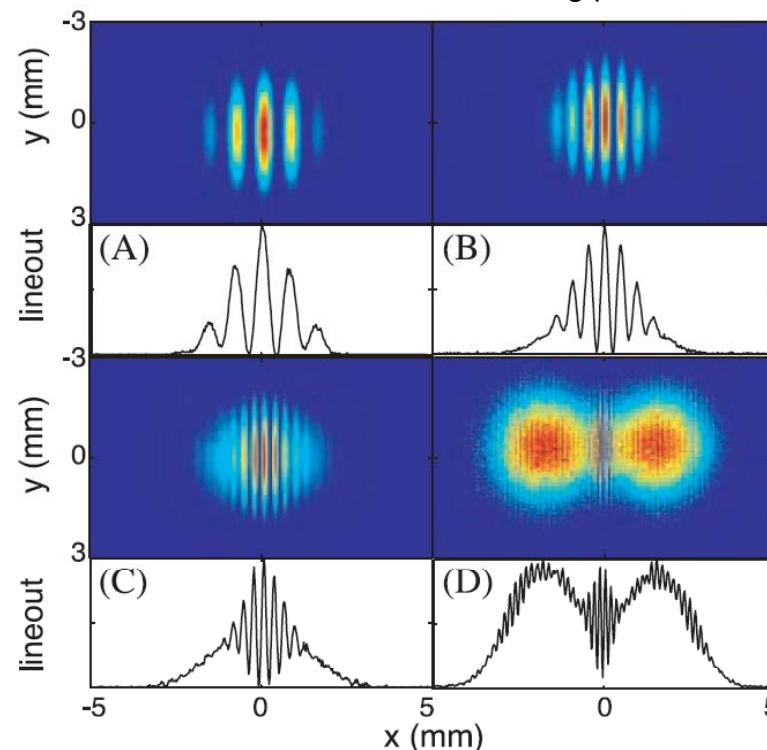
Sola et al., Nature Physics (2006)

Baltuska et al. Nature 421 (2003)

Young pinholes

H17-23

Young pinholes



Bartels et al., Science 297,376 (2002)

$$1.2 < M^2 < 4$$

Le Déroff et al., Optics Lett. 23 (1998)

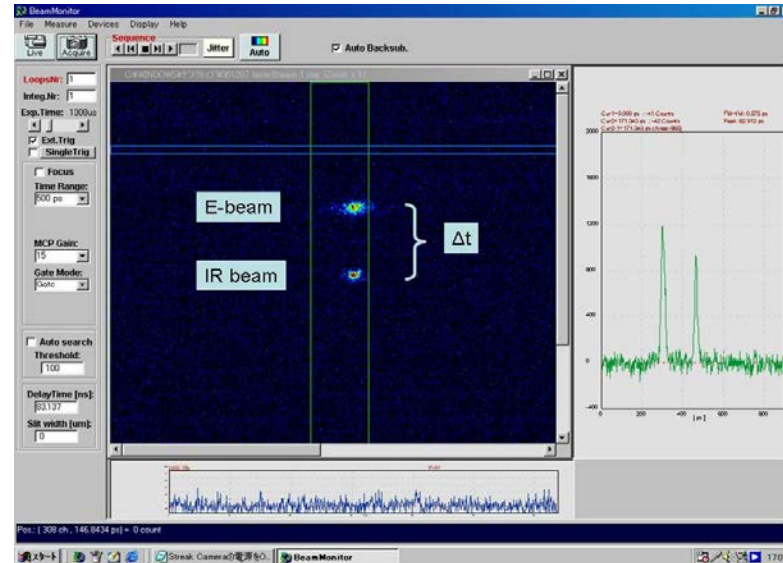
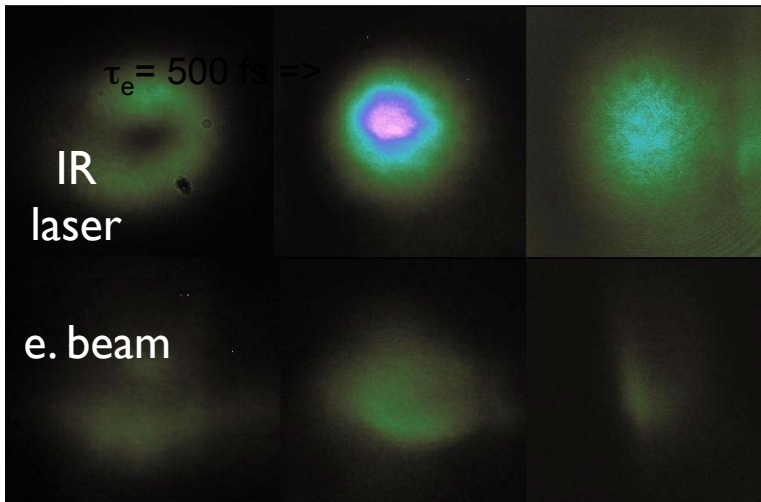
Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

## Issues for alignment and synchronisation

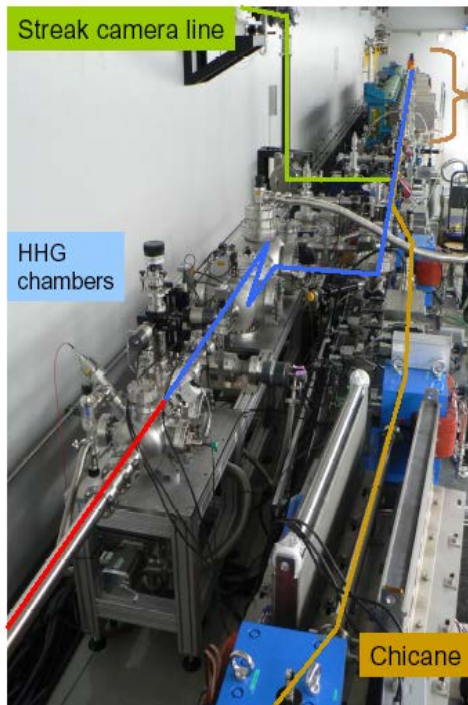
Monitoring the UV -VUV radiation of simply the generating laser?

Spatial overlap

Synchronisation

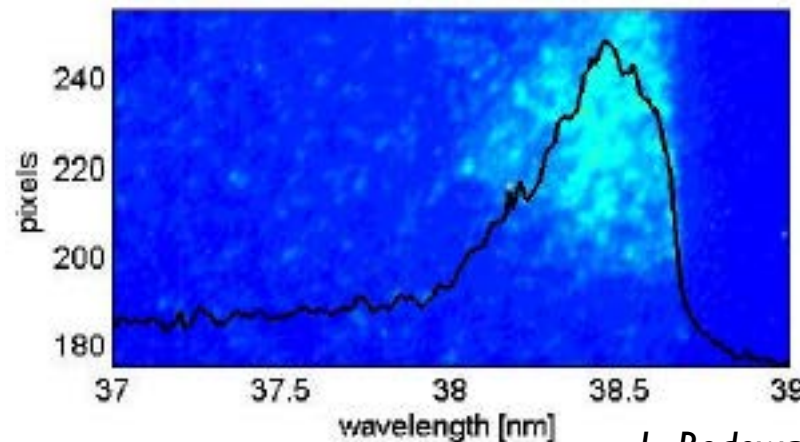


SCSS Test Accelerator



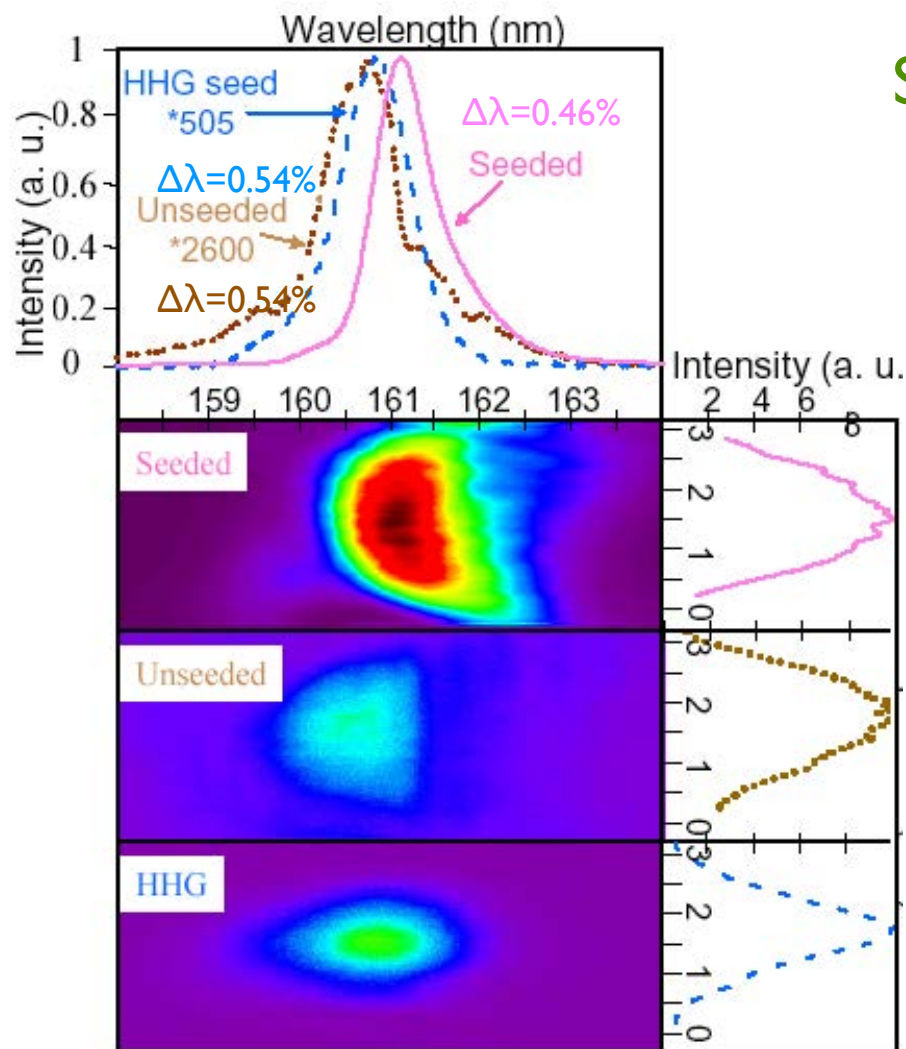
SCSS Test Accelerator

Spectral tuning



S-FLASH

# First demonstration of HHG seeding on the SCSS Test Accelerator (Japan)



Seed : 160 nm, 1 4.5 m long section of undulator, 150 MeV

Rather good alignment and spectral tuning, 0.3 nJ seed

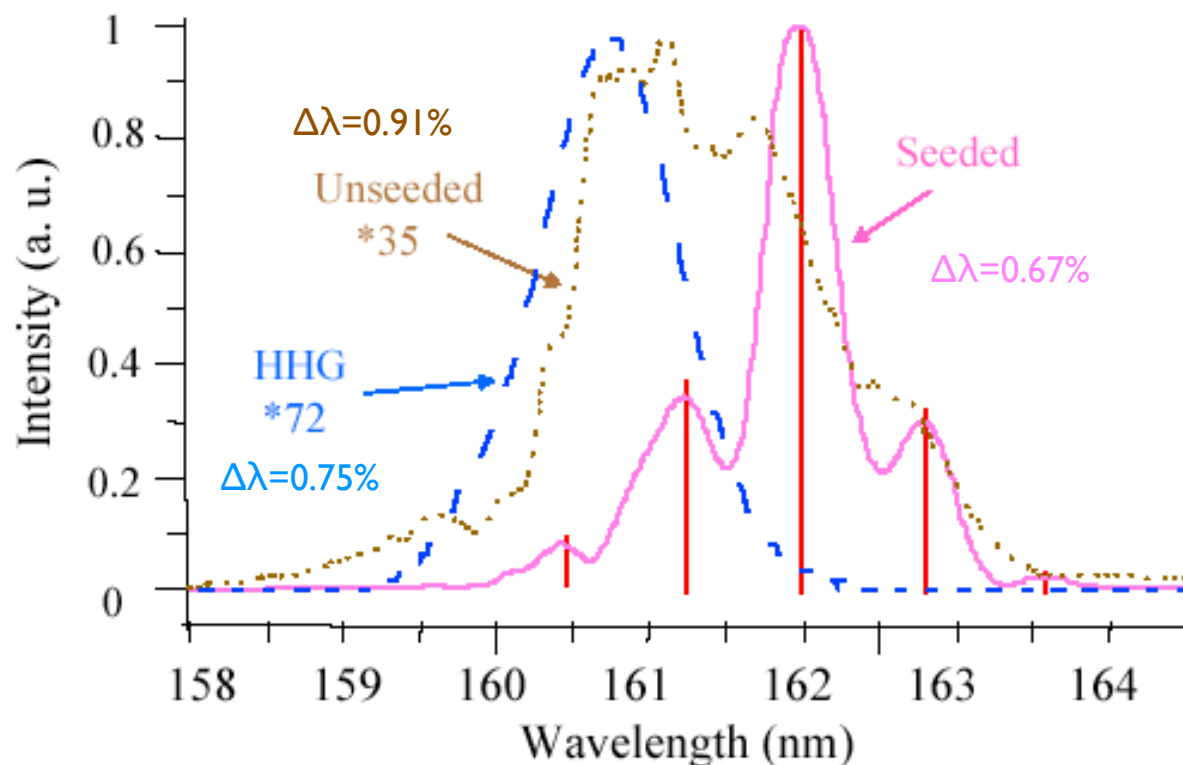
- Larger amplification
- Spectral narrowing => Improvement of the longitudinal coherence

G. Lambert et al., *Nature Physics Highlight*, (2008) 296-300



# First demonstration of HHG seeding on the SCSS Test Accelerator (Japan)

Seed : 160 nm, 2 sections of 4.5 m long undulator, 150 MeV

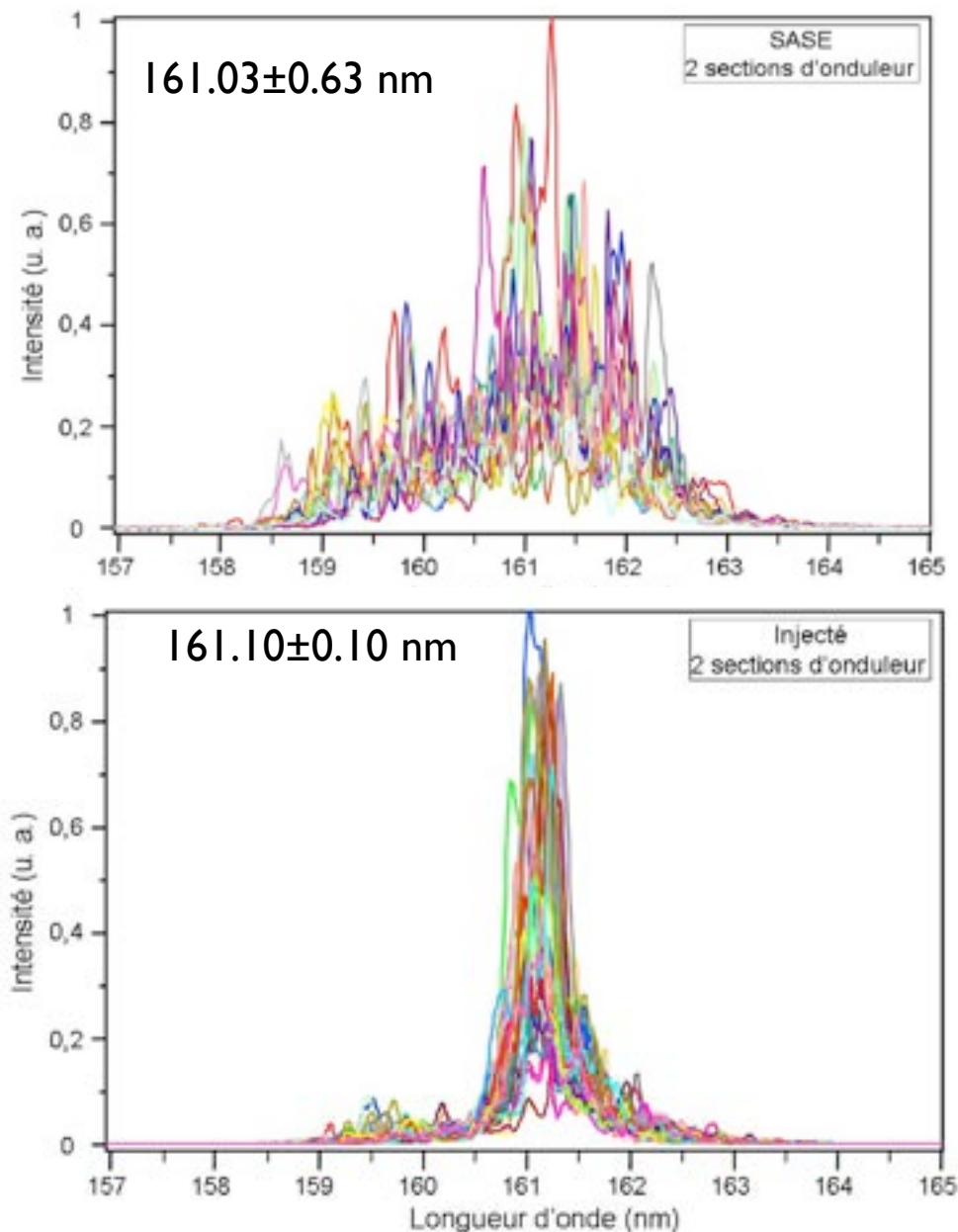
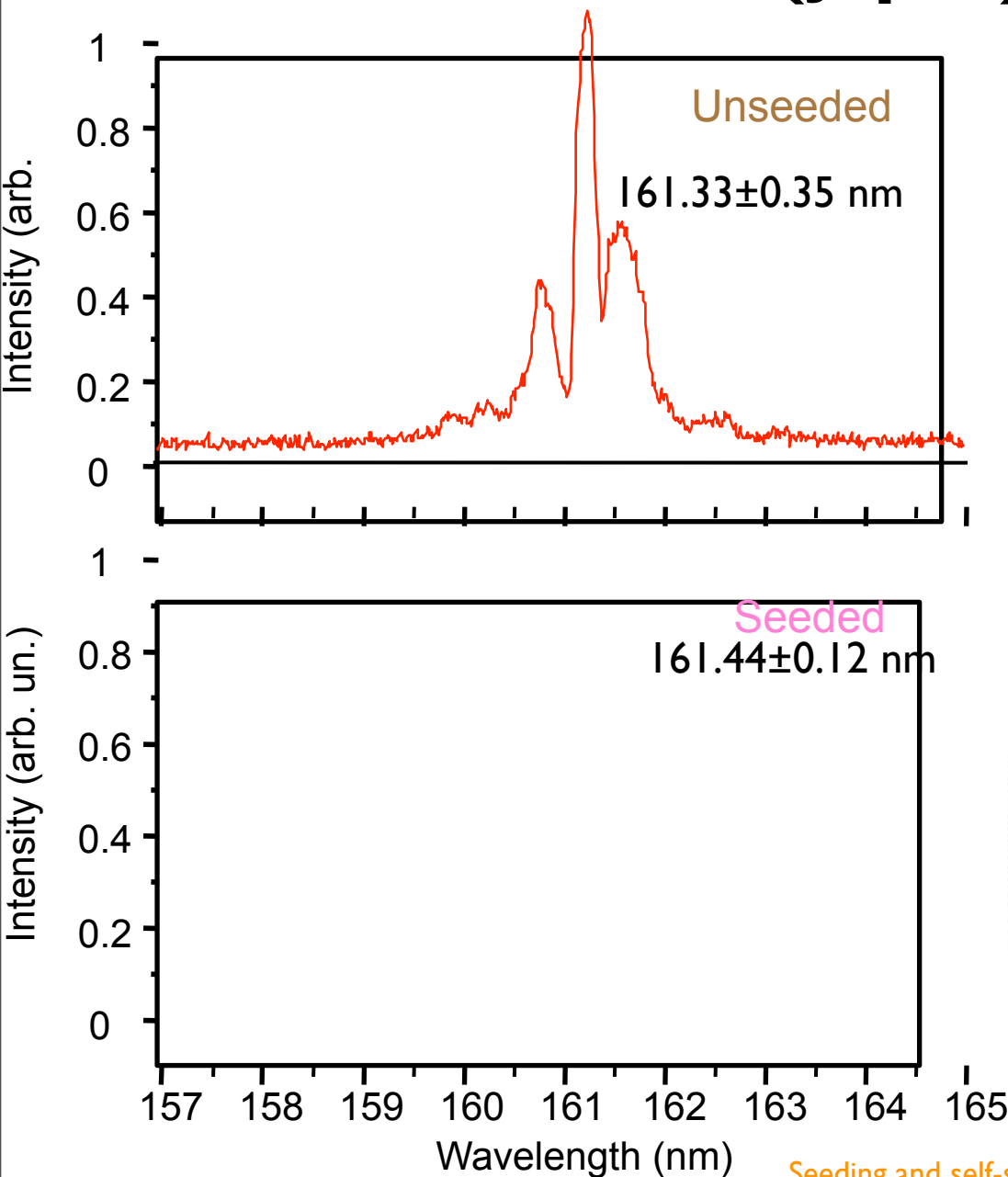


Imperfect alignment, seed 3 nj

- saturation for the HHG seeded FEL (sidebands)
- larger redshift
- smaller spectral narrowing
- limited amplification with respect to SASE

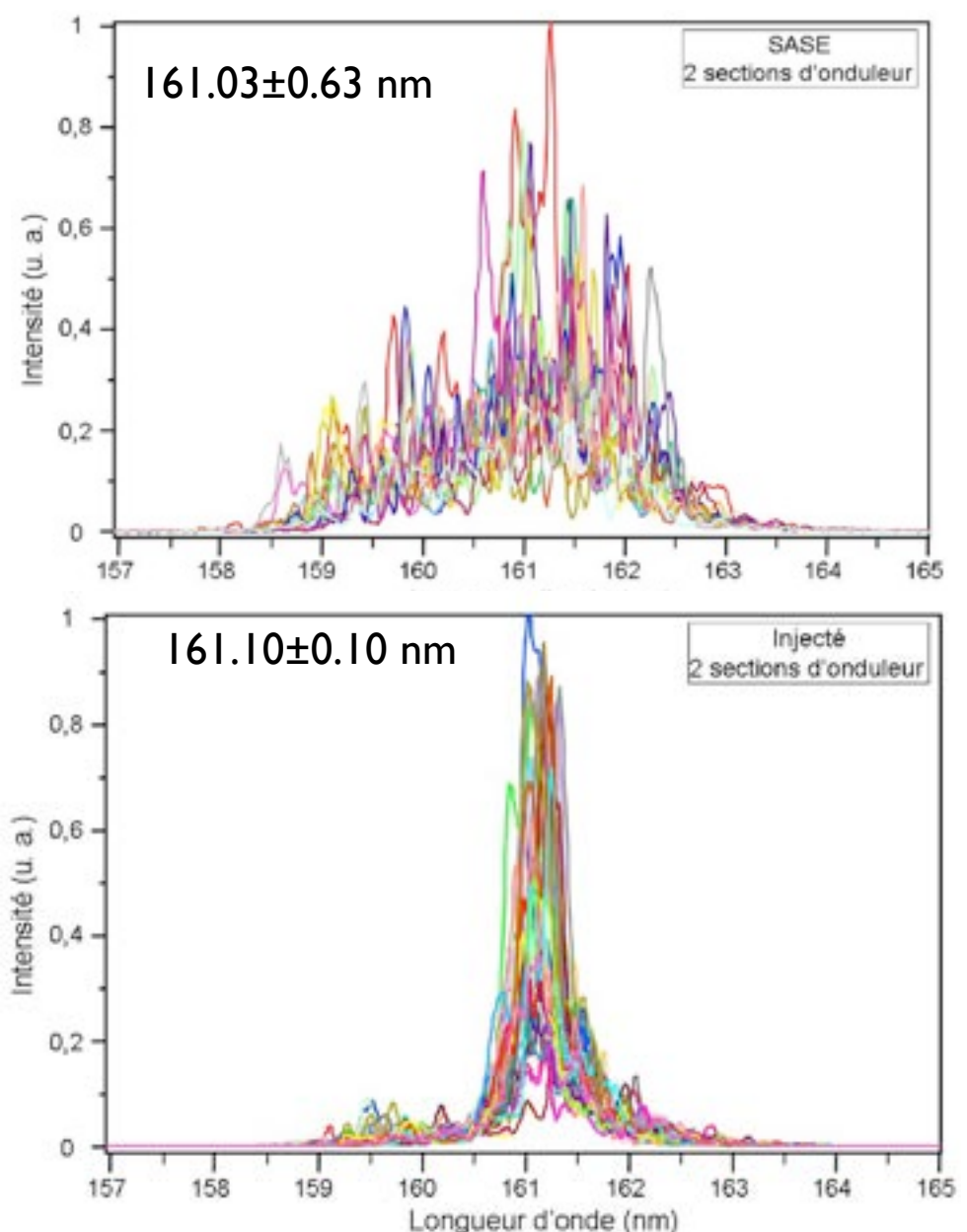
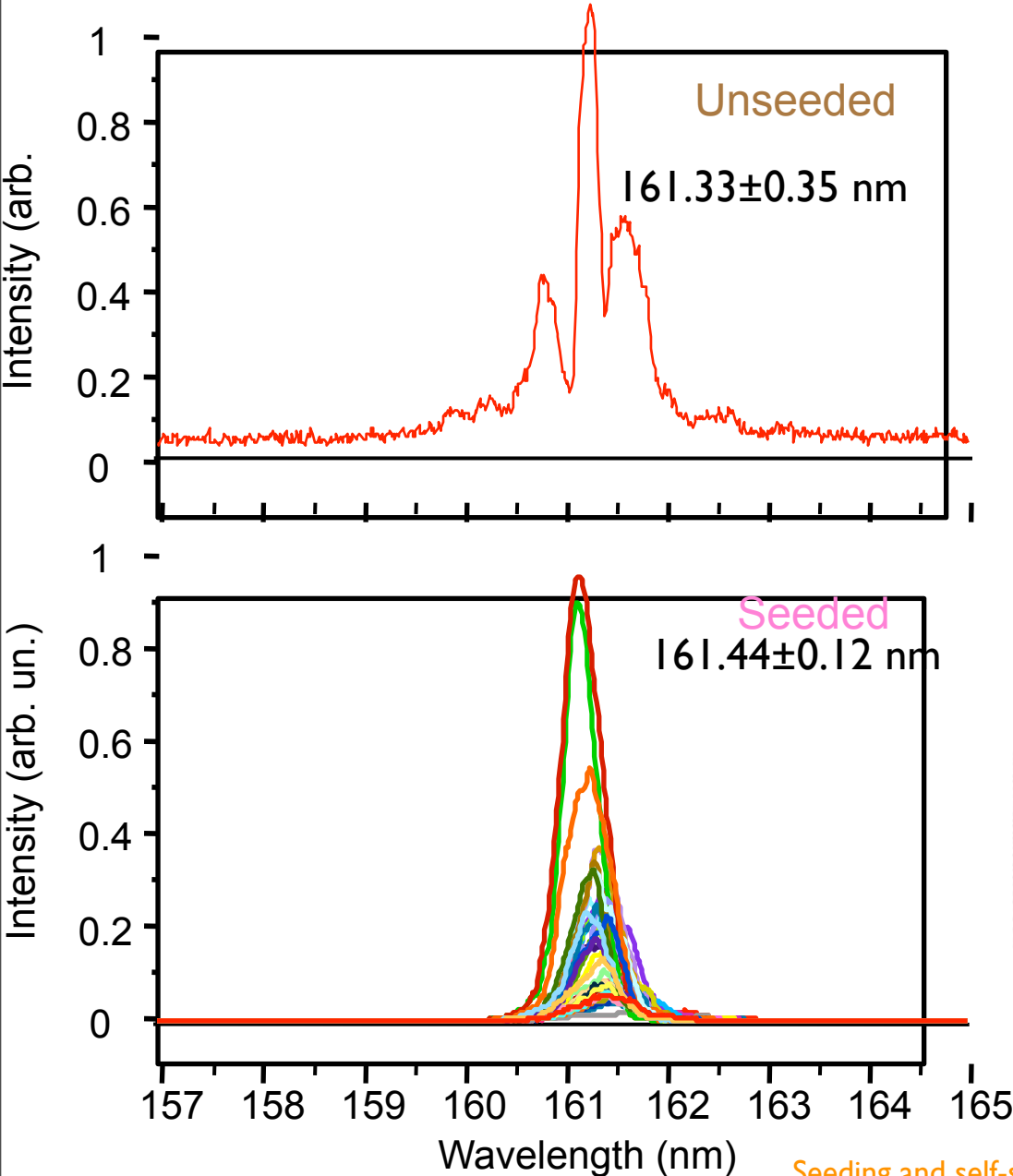
G. Lambert et al., *Nature Physics* 889 (2008)  
296-300

# First demonstration of HHG seeding on the SCSS Test Accelerator (Japan) : wavelength stability



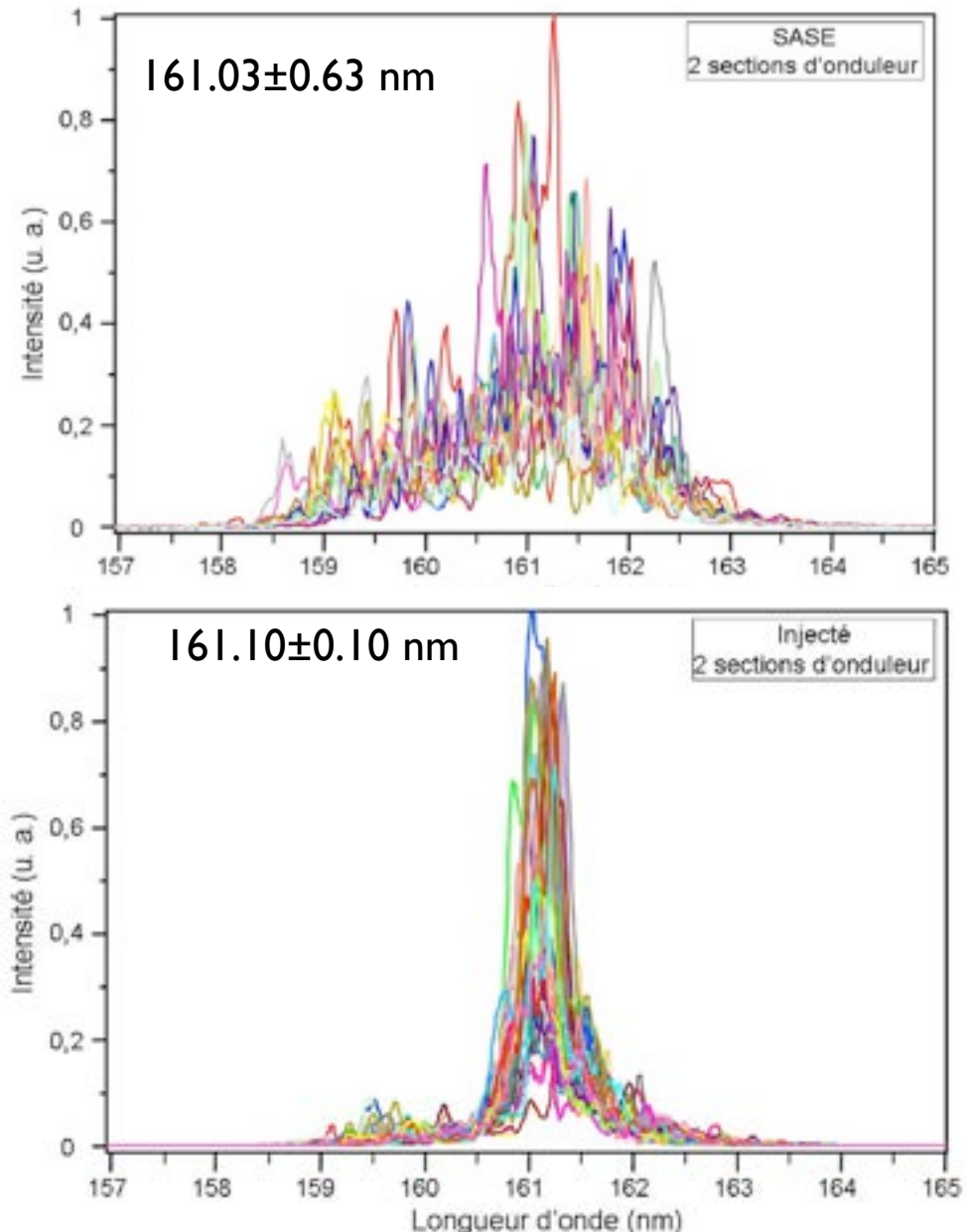
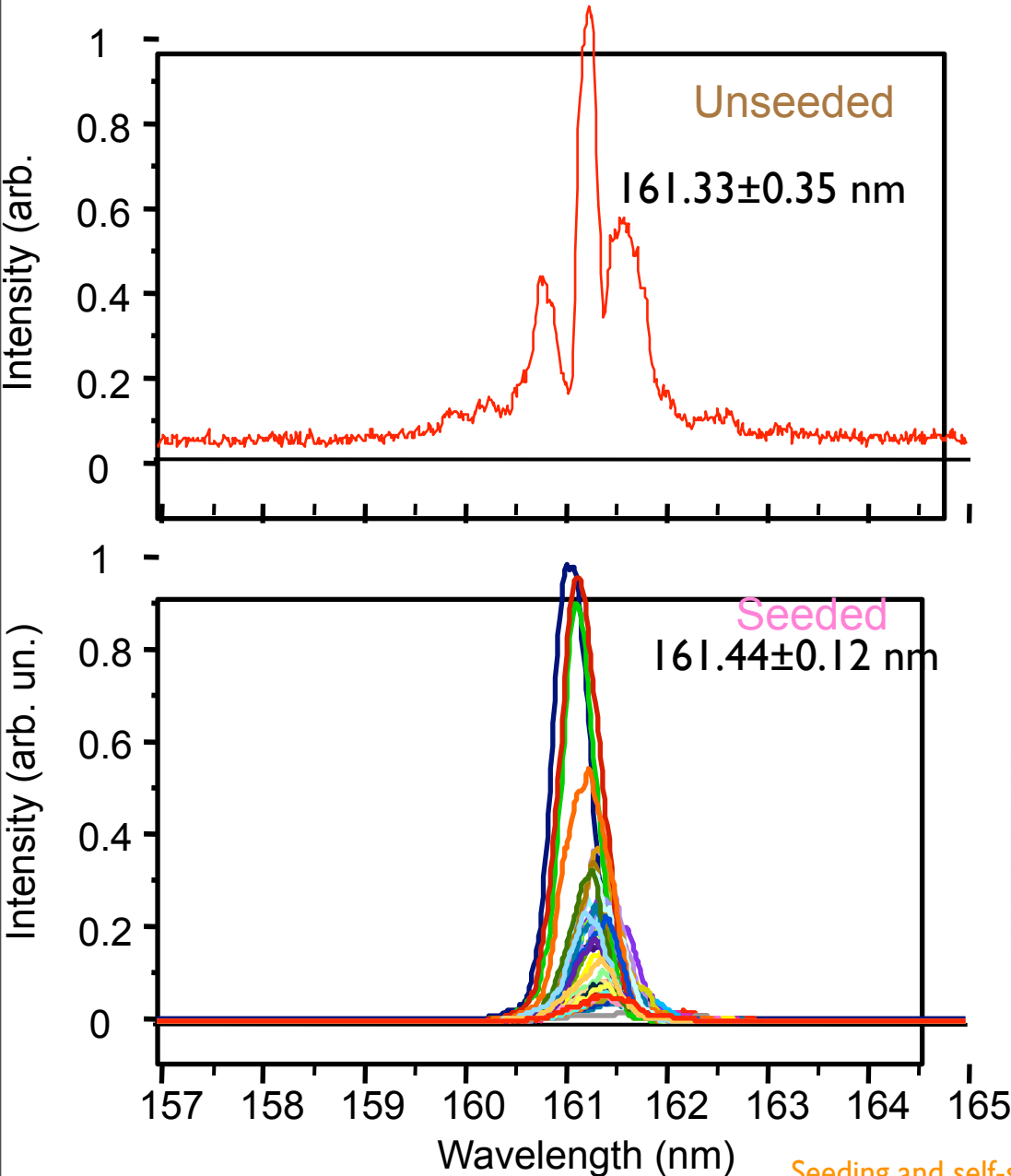
Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

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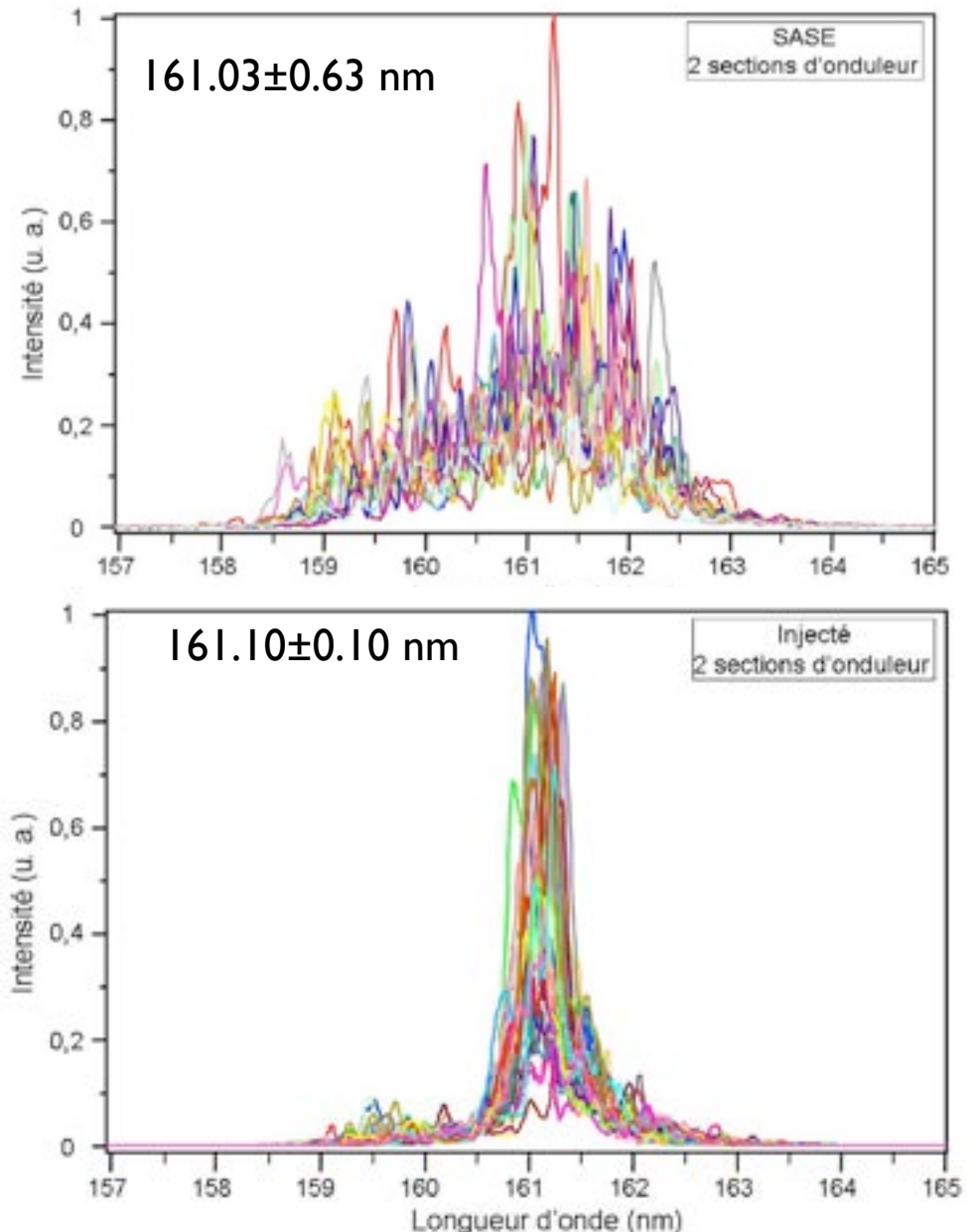
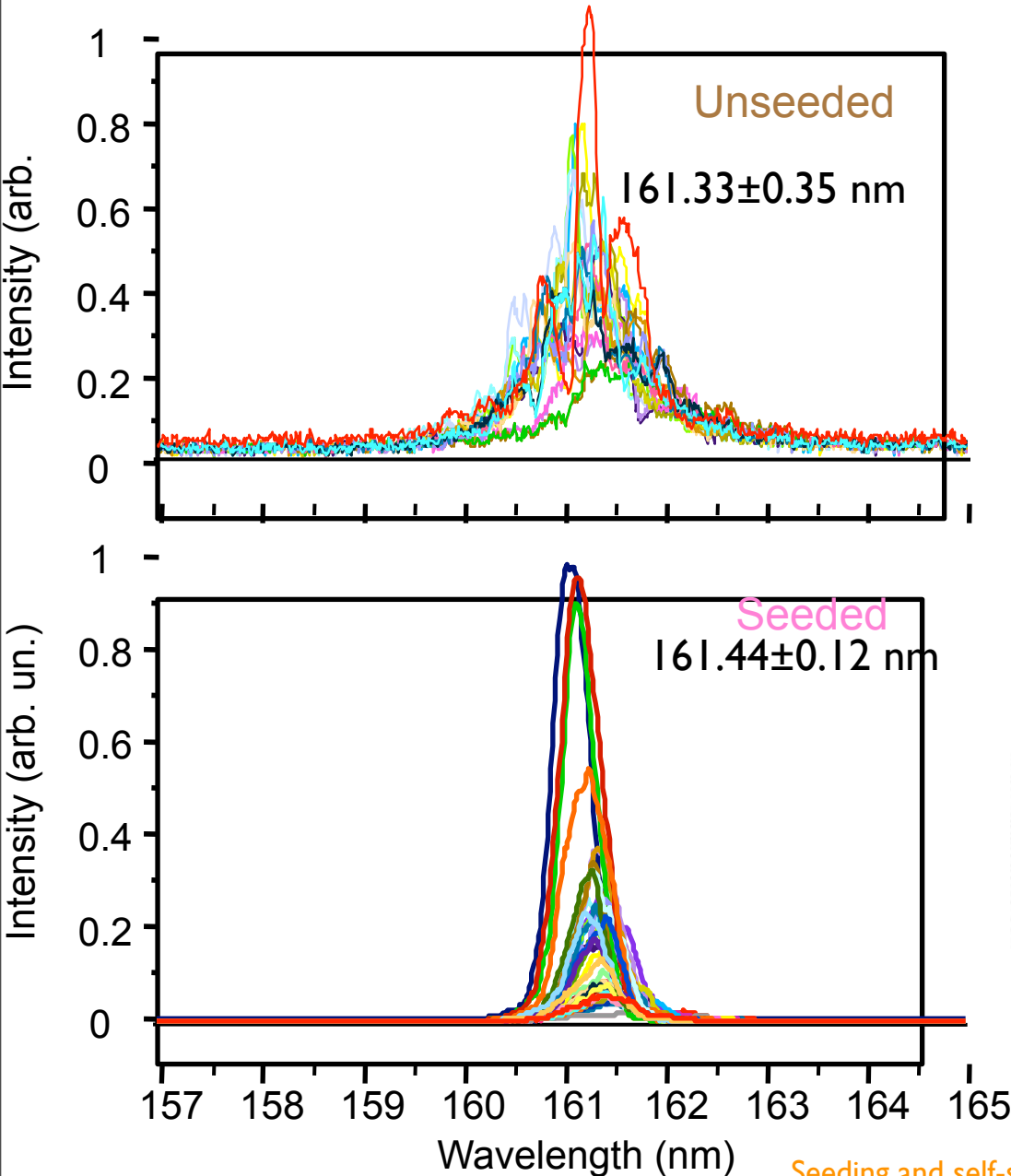
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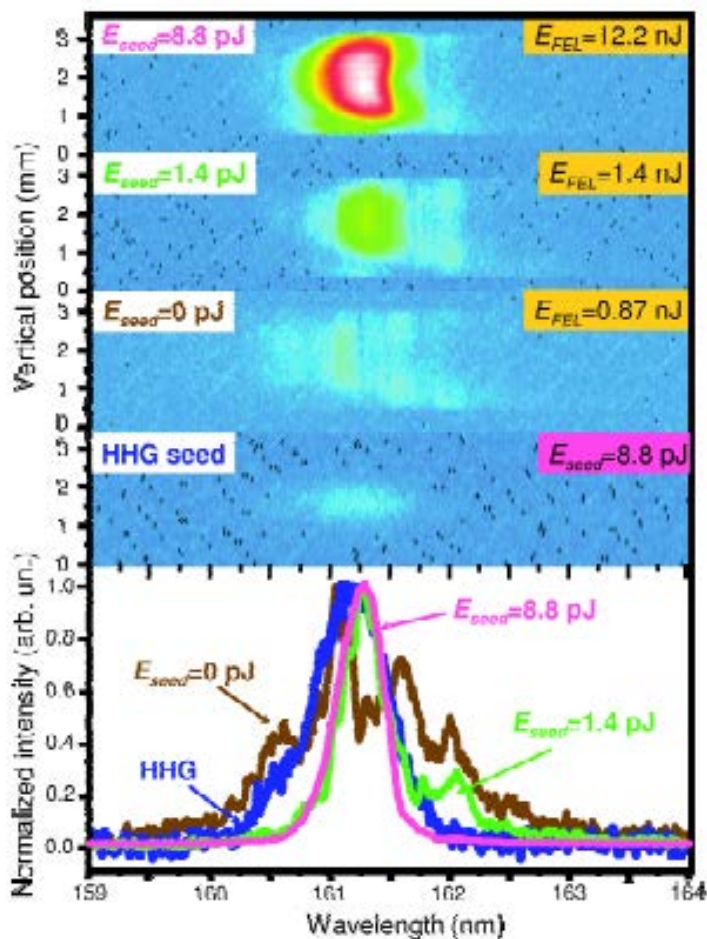


Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

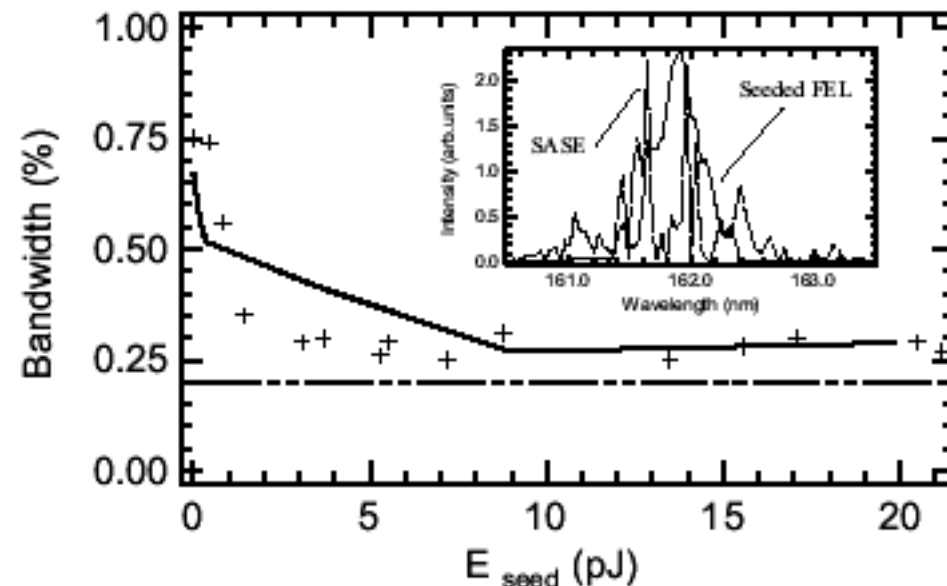
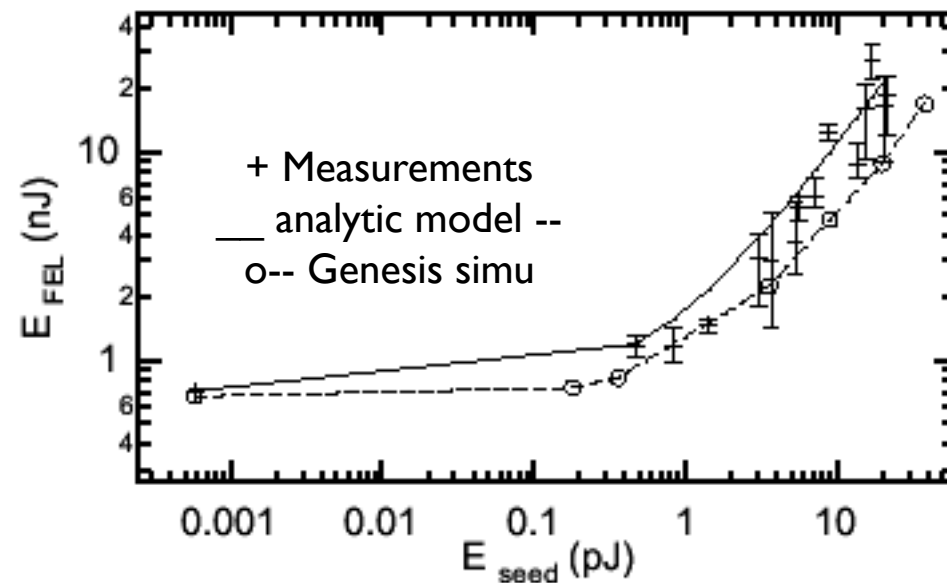
# Seed level study

Rather low HHG seed intensity required:

Debunched mode (3 ps, 1 undulator)



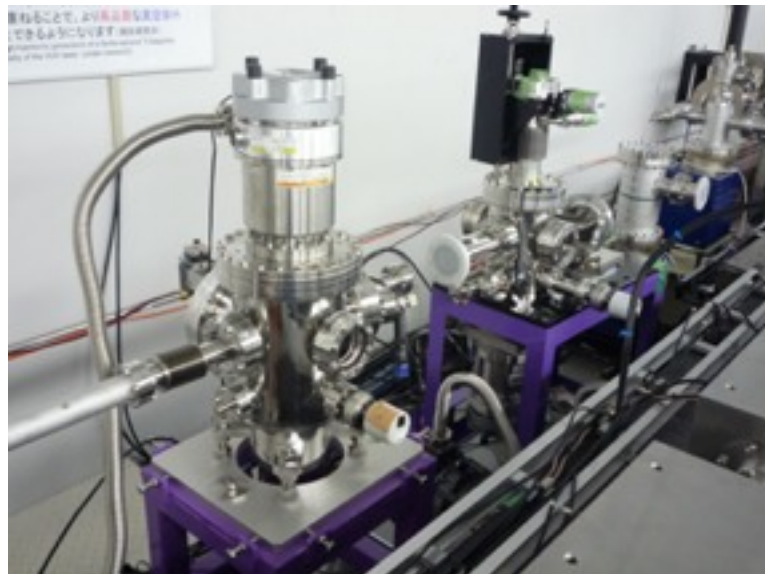
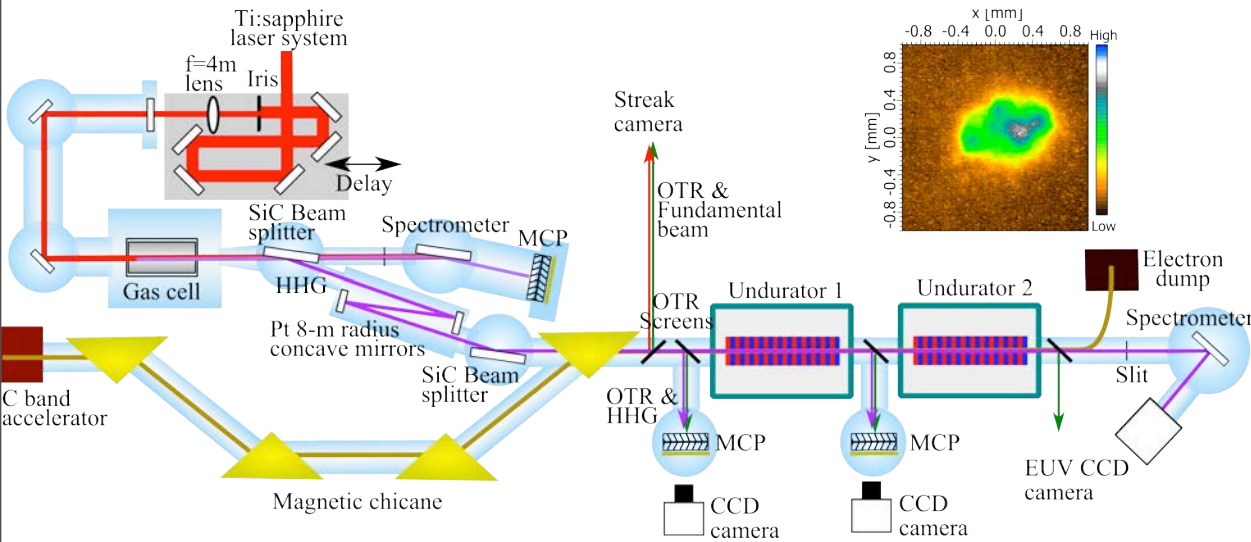
Two thresholds



G. Lambert, T. Hara, M. Labat, T. Tanikawa, Y. Tanaka, Y. Yabashi, D. Garzella, B. Carre, M. E. Couprie, Seed level requirement for improving temporal coherence of a FEL, *EPL* 88-5-54002, 2009

Source: <http://www.slac.stanford.edu/beamline/beamline.html>, Sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

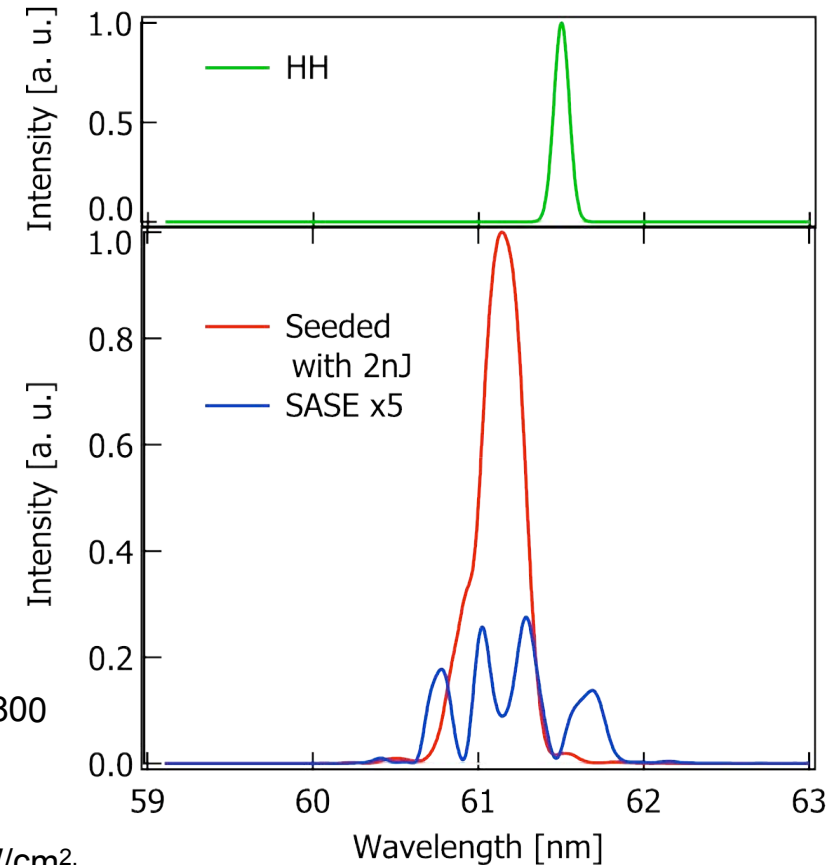
## Seeded FEL with plateau HHG @ 60 nm



Laser upgrade : ~200 mJ, 800 nm, 200 fs,  $\phi$ 50mm,  $M^2$ :~1.4  
Jitter: <200fs

HHG : 70 ~ 50 nm, > 5 MW/cm<sup>2</sup>.  
~ 200 fs,  $\phi$  1mm, > 10 nJ, 30 Hz

Extreme ultraviolet free electron laser seeded with high-order harmonic of Ti:sapphire laser  
T. Togashi, E. J. Takahashi, K. Midorikawa, M. Aoyama, K. Yamakawa, T. Sato, A. Iwasaki, S. Owada, T. Okino, K. Yamanouchi, F. Kannari, A. Yagishita, H. Nakano, M.E. Couprie, Kenji Fukami, T. Hatsui, I. T. Hara, T. Kameshima, H. Kitamura, N. Kumagai, I. S. Matsubara, M. Nagasono, I. H. Ohashi, T. Ohshima, Y. Otake, T. Shintake, K. Tamasaku, H. Tanaka, T. Tanaka, K. Togawa, H. Tomizawa, T. Watanabe, M. Yabashi, T. Ishikawa *Optics Express*, 1, 2011, 317-324

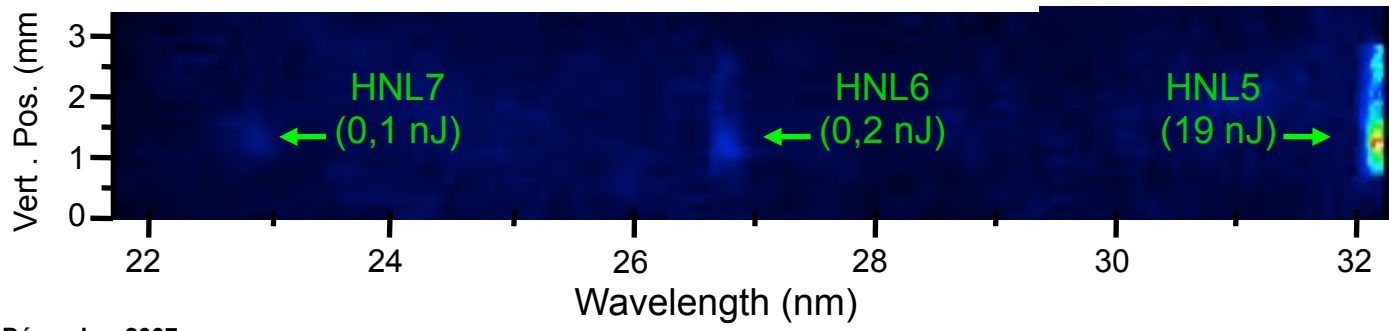
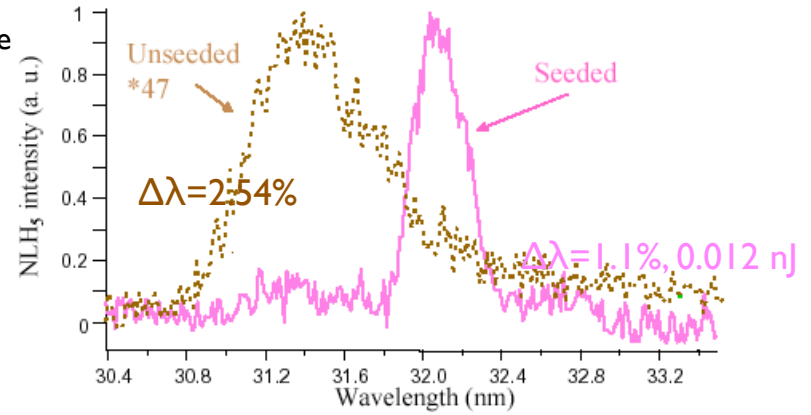
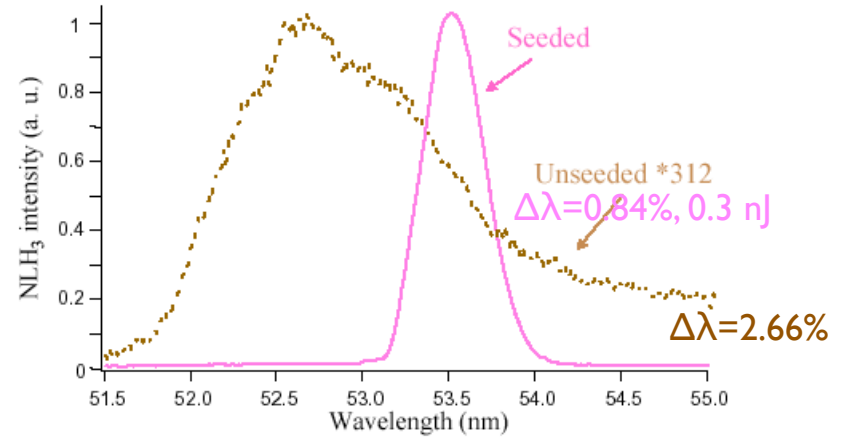
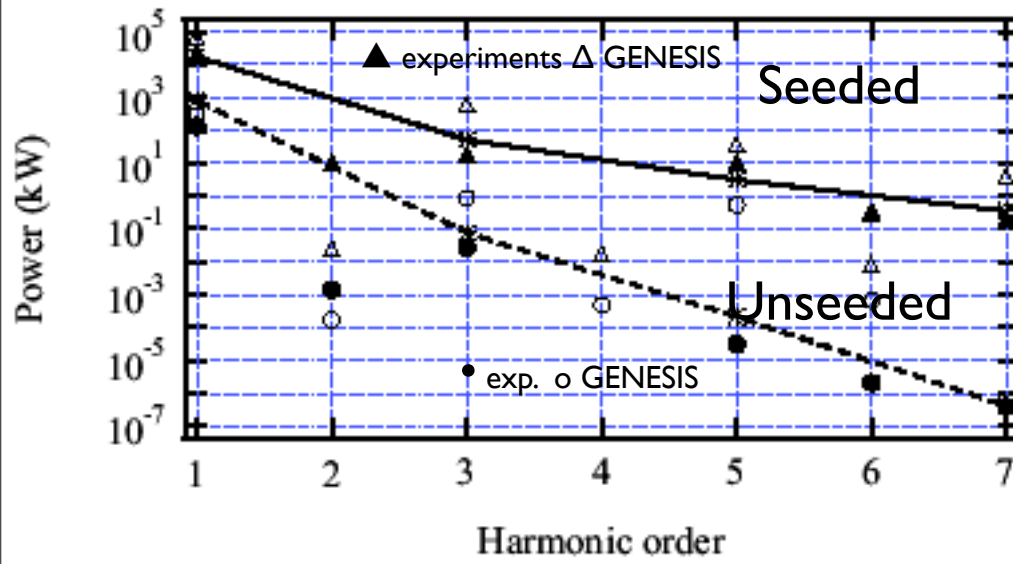


## Non linear harmonics

Seed : 160 nm, 2 sections of 4.5 m long undulator, 150 MeV, bunched beam 400 fs, 0.3 nJ

- Large redshift
- spectral narrowing

Debunched beam 400 fs, 2 und. sections



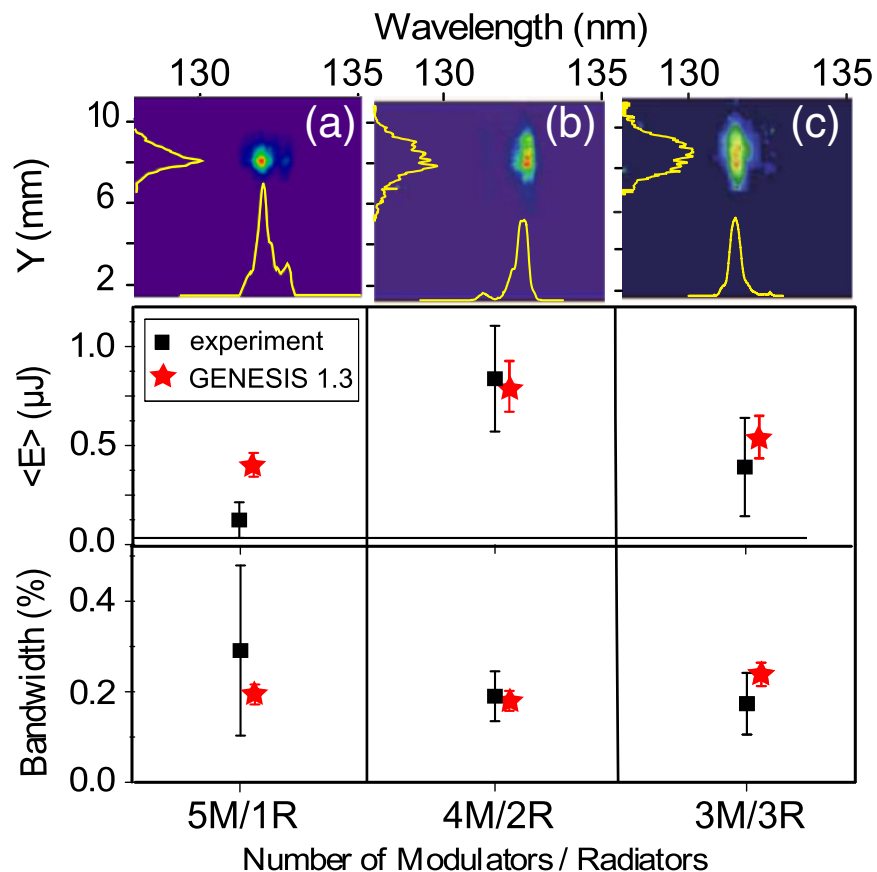
Décembre 2007

Nonlinear harmonic generation in a free-electron laser seeded with high harmonic from gas T. Tanikawa, G. Lambert, T. Hara, M. Labat, Y. Tanaka, M. Yabashi, O. Chubar, M.E. Couprie, EPL 2013 Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012



# HHG seeding + cascaded configuration

SPARC  
266 nm  
133 nm



M. Labat et al., Phys. Rev. Lett. 107, 224801 (2011)

# Further prospects opened now by HHG in atomic/ molecular gas

Major evolutions 2008-2012:

## 1. HHG driven in IR (800 nm) by a few-cycle pulse

Generation of Quasi-continuum (near single atto pulse) → tunability

## 2. HHG driven by mid-IR OPA/OPCPA lasers sources (1-4 μm)

- Cutoff :  $h\nu_{max} = I_p + 3.17U_p$ ;  $U_p \propto I_L \lambda_L^2$  → extension to keV range with good efficiency
- Generation of Quasi-continuum (near single atto pulse) → tunability

## 3. HHG in aligned molecules

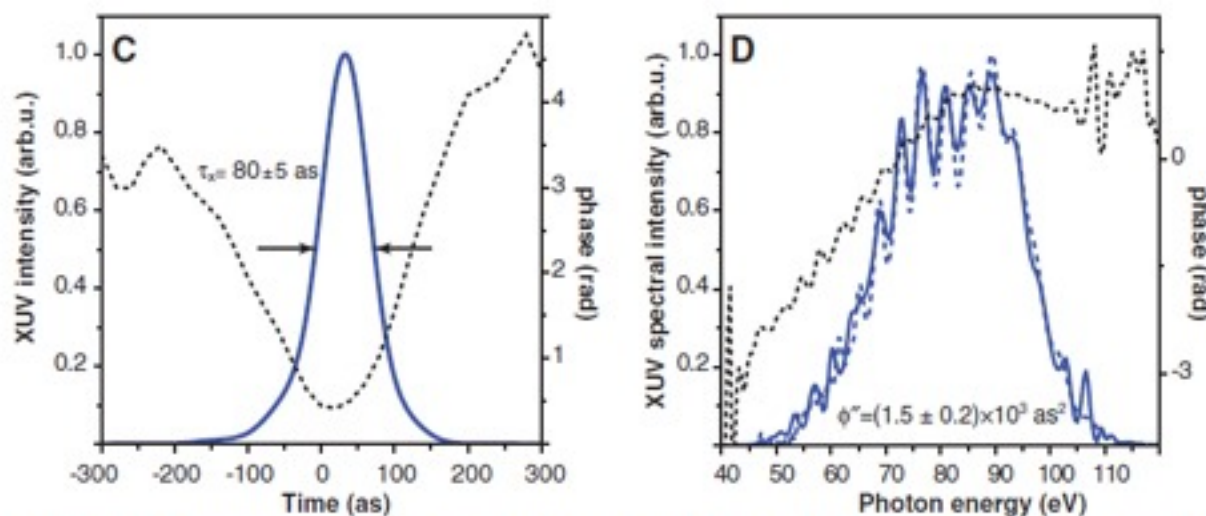
- Generation of elliptical polarization

# HHG driven in IR (800 nm) by a few-cycle pulse

## Single-cycle nonlinear optics

*E. Goulielmakis, M. Schultze, M. Hofstetter, V. S. Yakovlev, J. Gagnon, M. Uiberacker, A. L. Aquila, E. M. Gullikson, D. T. Attwood, R. Kienberger, F. Krausz, U. Kleineberg, Science 320, 1614-1617 (2008).*

- Generation of Quasi-continuum (near single atto pulse) → tunability



ATR : atomic transient recorder

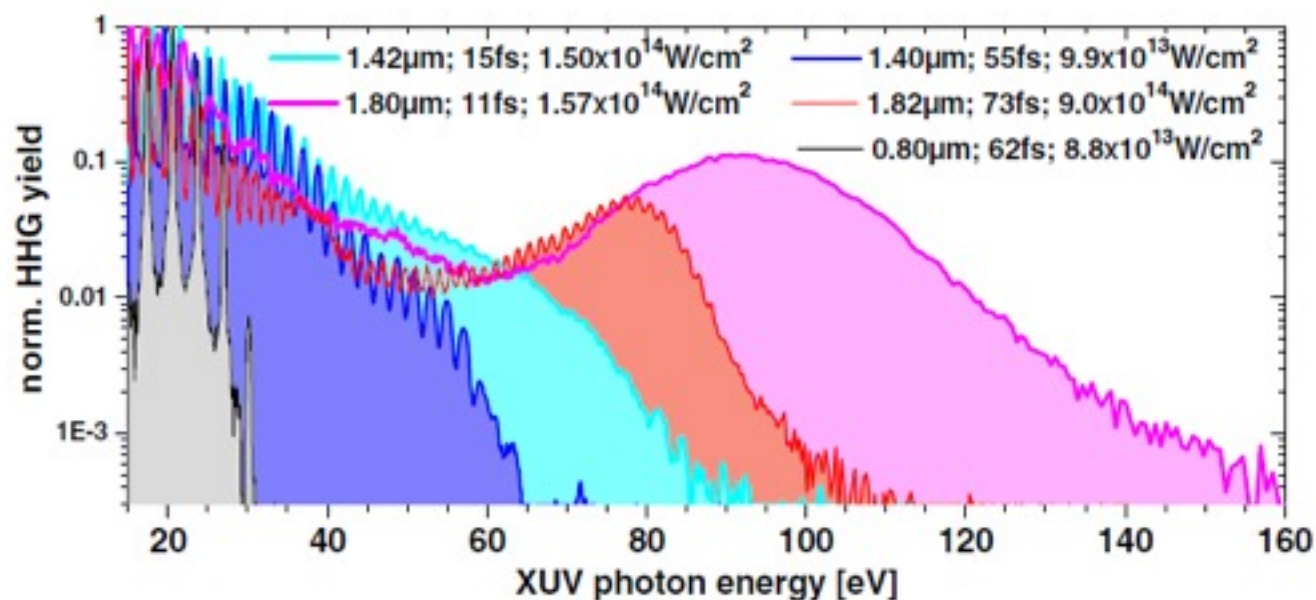
**Fig. 3.** Sub-100-as XUV pulse retrieval. **(A)** Measured ATR spectrogram compiled from 126 energy spectra of photoelectrons launched by an XUV pulse with a bandwidth of  $\sim 28$  eV (FWHM) and recorded at delay settings increased in steps of 80 as. Here, a positive delay corresponds to the XUV pulse arriving before the NIR pulse. The high flux of the XUV source allows this spectrogram to be recorded within  $\sim 30$  min. **(B)** ATR spectrogram reconstructed after  $\sim 10^3$  iterations of the FROG algorithm (17). **(C)** Retrieved temporal intensity profile and spectral phase of the XUV pulse. The intrinsic chirp of the XUV emission (Fig. 4B) is almost fully compensated by a 300-nm-thick Zr foil introduced into the XUV beam between the attosecond source and the ATR measurement. Arrows indicate the temporal FWHM of the XUV pulse. **(D)** XUV spectra evaluated from the measurement of the XUV-generated photoelectron spectrum in the absence of the NIR streaking field (blue dashed line) and from the ATR retrieval (blue solid line). The black dotted line indicates the retrieved spectral phase.

# HHG driven by mid-IR laser

- Cutoff extension :  $h\nu_{max} = I_p + 3.17U_p$ ;  $U_p \propto I_L \lambda_L^2$
- Generation of Quasi-continuum (near single atto pulse) → tunability

## High harmonic generation with long-wavelength few-cycle laser pulses

B. E. Schmidt, A. D. Shiner, M. Giguère, P. Lassonde, C. a Trallero-Herrero, J.-C. Kieffer, P. B. Corkum, D. M. Villeneuve, and F. Légaré, *J. Phys. B* 45, 074008 (2012)



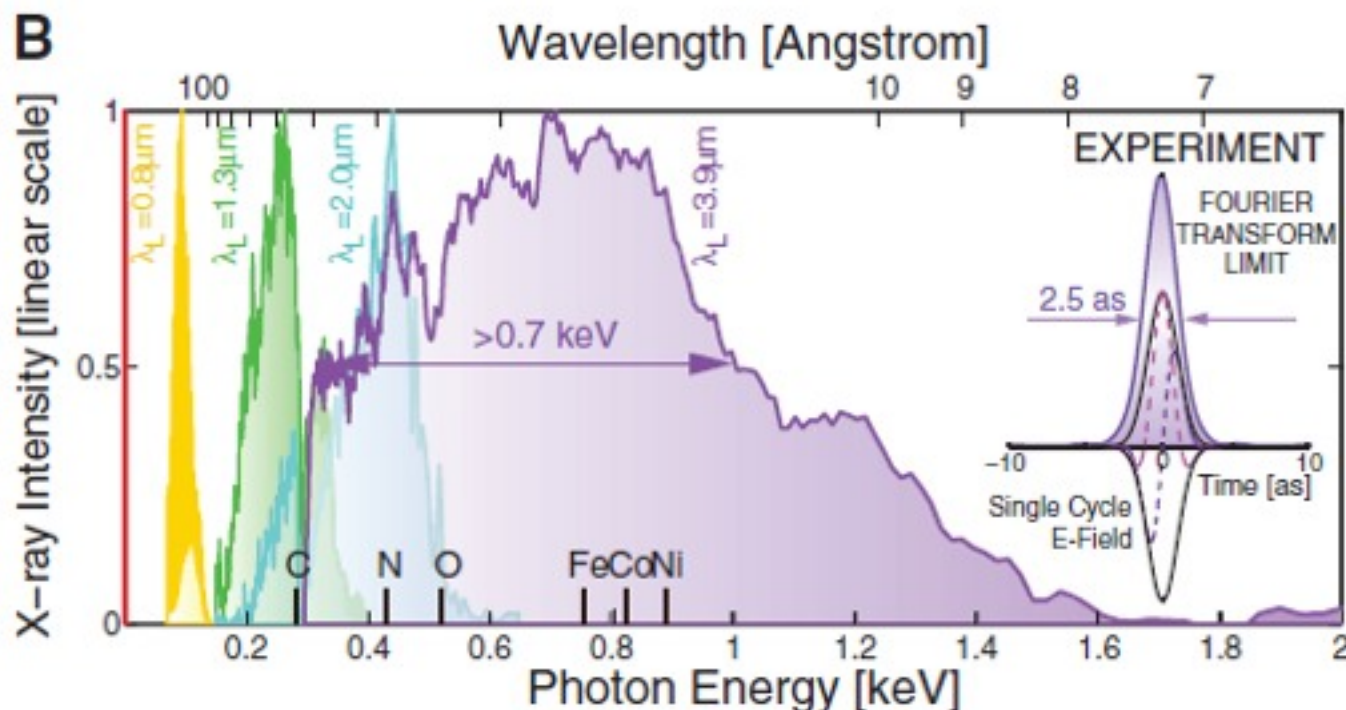
**Figure 5.** Semi-log plot of high harmonic yield in xenon at different driving laser wavelengths (grey: 0.8  $\mu\text{m}$ ; bluish: 1.4  $\mu\text{m}$ ; reddish: 1.8  $\mu\text{m}$ ). Each measurement was performed at intensity close to the corresponding saturation intensity and shows the benefit of increased driving wavelength and decreased pulse duration, respectively. Pulse durations and intensities are given in the legend. The striking enhancement around 95 eV originates from the giant resonance in xenon, an effect due to multi-electron correlation upon the recombination step in HHG.

# HHG driven by mid-IR laser

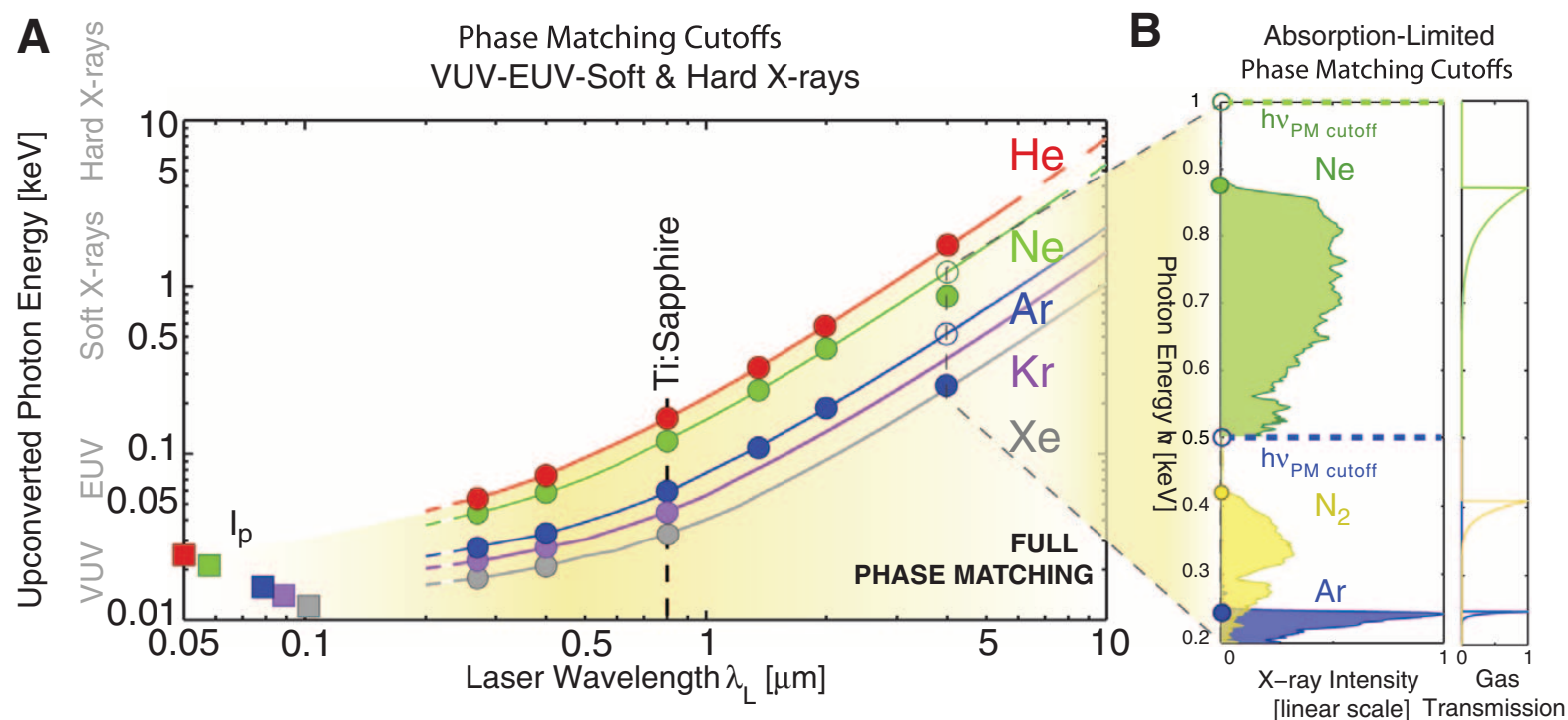
Bright Coherent Ultrahigh Harmonics in the keV X-ray Regime from Mid-Infrared Femtosecond Lasers

Popmintchev, T; Chen, MC; Popmintchev, D; Arpin, P; Brown, S; Alisauskas, S; Andriukaitis, G; Balciunas, T; Mucke, OD; Pugzlys,; Baltuska, A; Shim, B; Schrauth, SE; Gaeta, A; Hernandez-Garcia, C; Plaja, L; Becker, A; Jaron-Becker, A; Murnane, MM; Kapteyn, HC, Science 336, 1287-1291 (2012)

- HHG driven in (high pressure) He by mid-IR laser at  $3.7 \mu\text{m}$  : cutoff at 1.6 keV
- Quasi continuum  $\rightarrow$  near single atto pulse

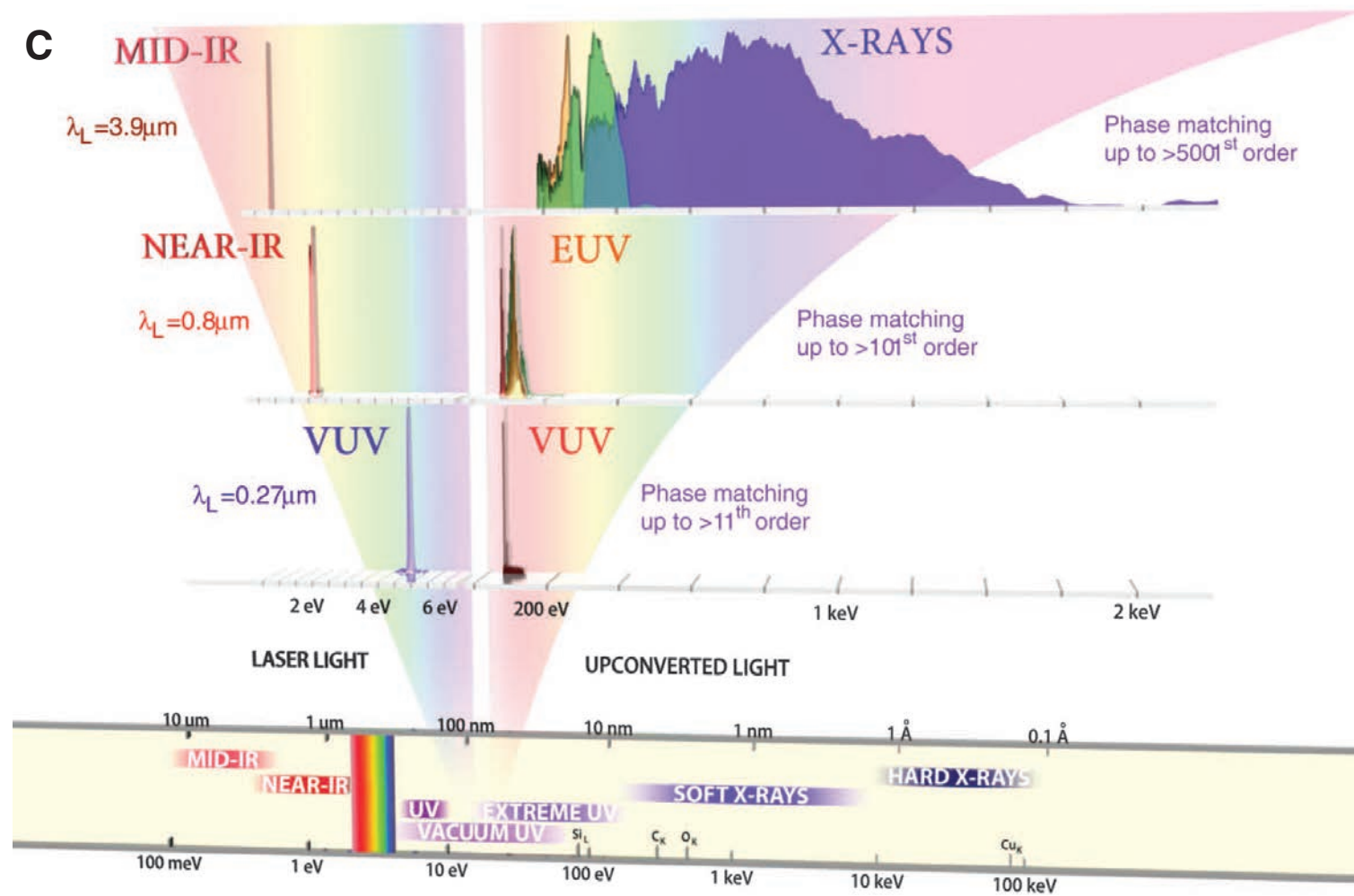


# HHG driven by mid-IR laser



**Fig. 2.** (A) Predicted and observed HHG phase-matching cutoffs as a function of laser wavelength from the UV to the mid-IR. Solid circles show the observed cutoffs; open circles show the predicted cutoffs for Ar and Ne [which cannot be reached due to inner-shell absorption, as shown in (B)]. Solid squares on the left show the ionization potentials ( $I_p$ ) of the different atoms. (C) Unified picture of optimal phase-matched high-harmonic upconversion, including microscopic and macroscopic effects.

## HHG driven by mid-IR laser



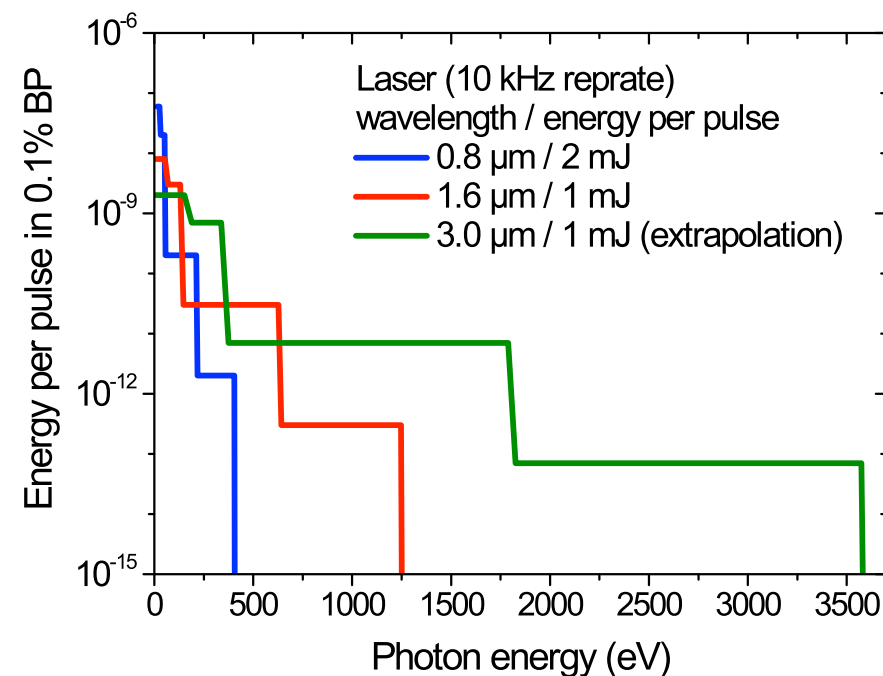
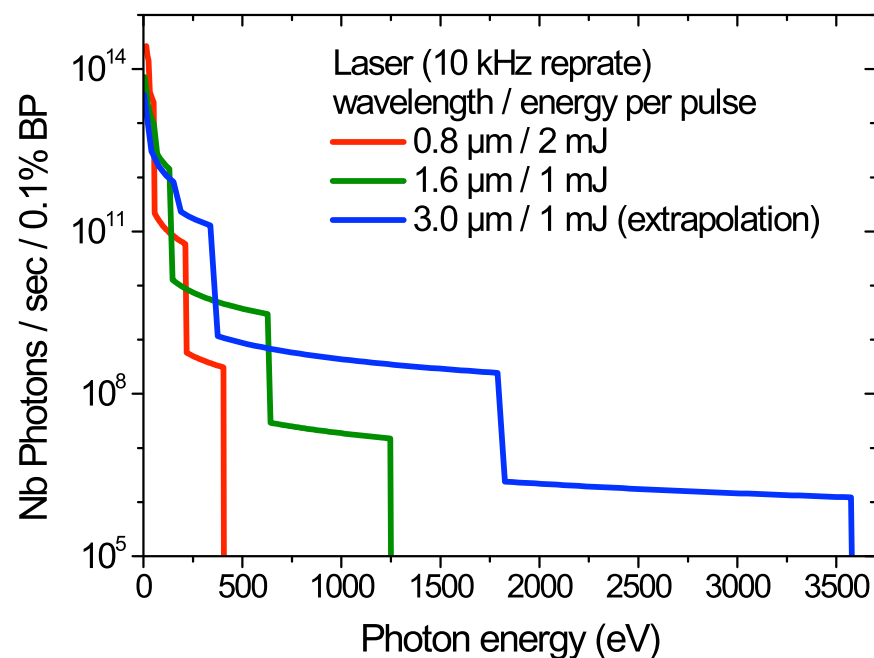
# Flux : orders of magnitude

Laser (10 kHz replate) wavelength / energy per pulse

0.8  $\mu\text{m}$  / 2 mJ

1.6  $\mu\text{m}$  / 1 mJ

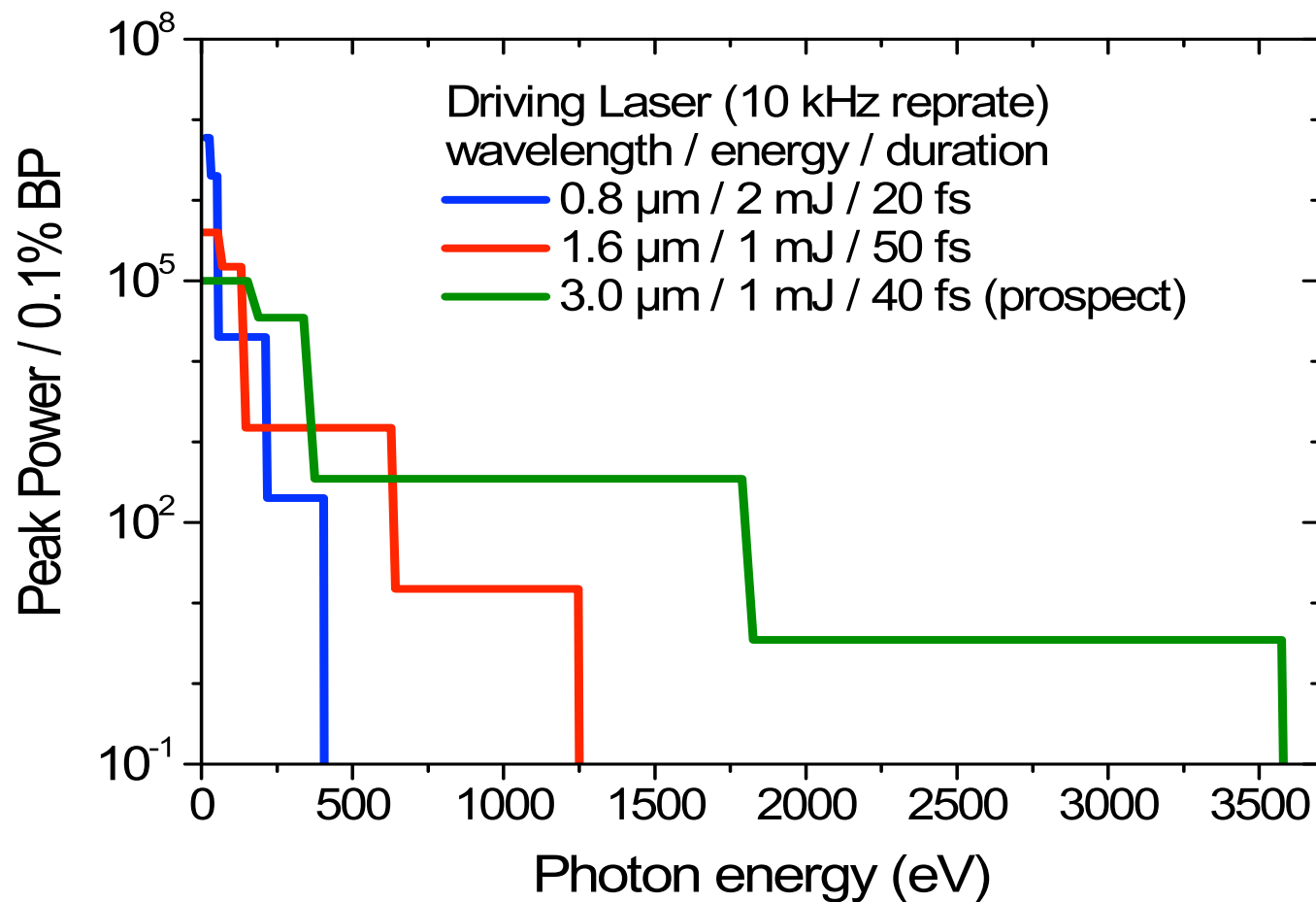
3.0  $\mu\text{m}$  / 1 mJ (extrapolation)



(origin graphs by BC)



# Flux : orders of magnitude



(origin graphs by BC)

# Polarization

Elliptical polarization can be produced in HHG from aligned molecules

Origin : Phase difference between  $d_{\parallel}$  and  $d_{\perp}$  components of the dipole

(due to non isotropic Coulomb potential, recolliding electronic wave  $\neq$  plane wave, multiple orbital dynamics)

Elliptically Polarized High-Order Harmonic Emission from Molecules in Linearly Polarized Laser Fields

X. Zhou, R. Lock, N. Wagner, W. Li, H. C. Kapteyn, and M. M. Murnane, PRL 102, 073902 (2009)

Ellipticity close to 0.4 in  $N_2$

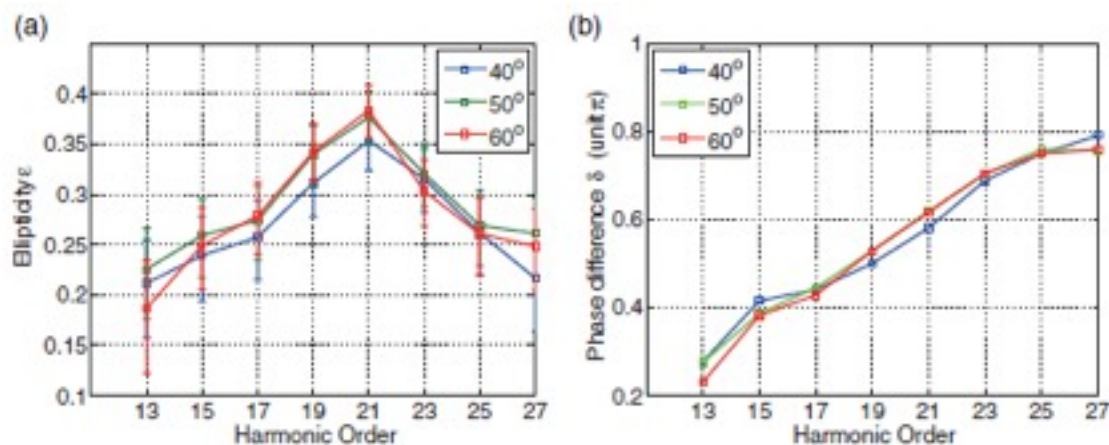


FIG. 4 (color online). (a) Ellipticity  $\varepsilon$  for relative angles between the pump and probe of  $\theta = 40^\circ$ ,  $50^\circ$ , and  $60^\circ$ , as a function of harmonic order for  $N_2$ . (b) Calculated phase difference  $\delta$  between the parallel and perpendicular components of HHG emission in  $N_2$ .

cannot be explained with single active electron model and strong field approximation

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

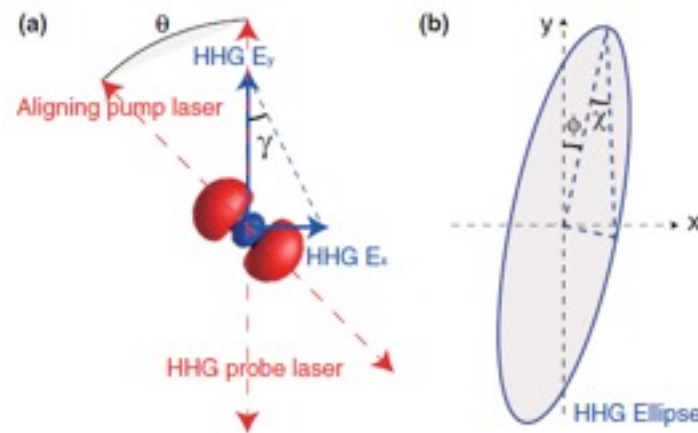
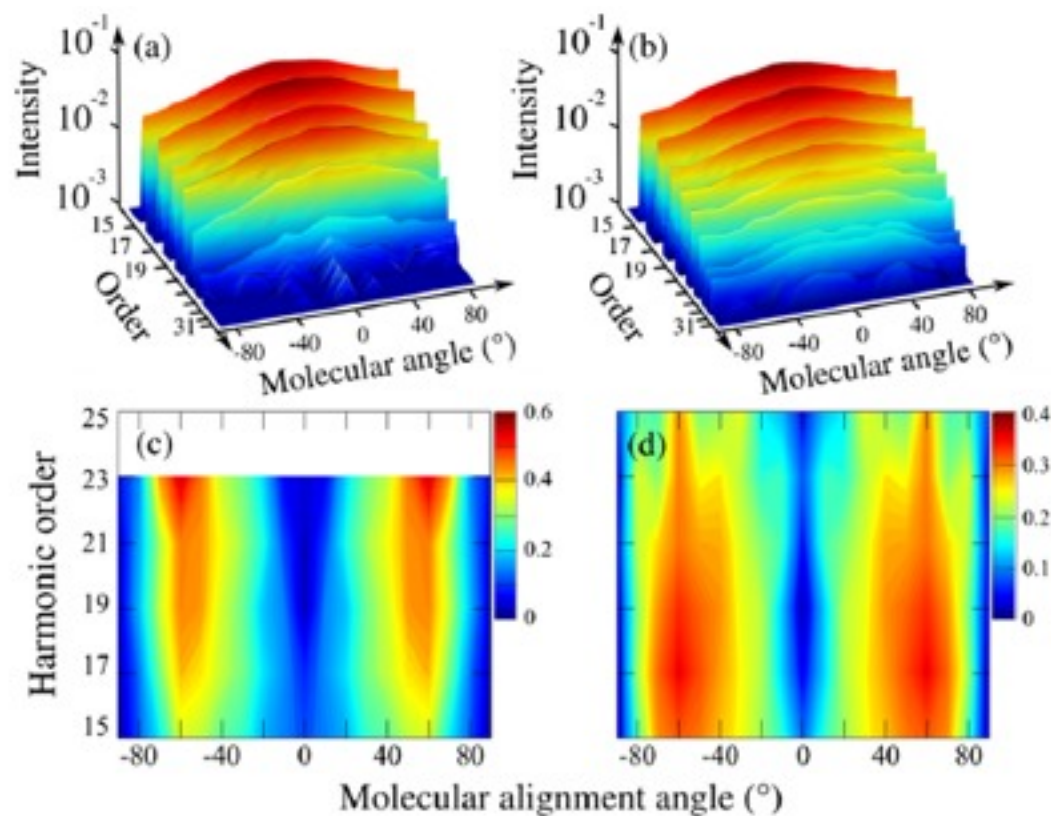


FIG. 1 (color online). High harmonic generation from aligned  $N_2$ . (a) HHG components parallel and perpendicular to the probe laser polarization can be generated. The angle  $\gamma$  is defined by  $\tan(\gamma) = |E_x|/|E_y|$ , where  $x$  and  $y$  are the lab frame axes. (b) The orientation angle  $\phi$  of the HHG ellipse is defined as the angle between the major axis of the ellipse (max. transmission direction through polarizer) and the  $y$  axis. The angle  $\chi$  is defined by  $\tan(\chi) = \varepsilon$ , where  $\varepsilon$  is the ellipticity.  $\phi$  and  $\theta$  are positive for clockwise rotation from the  $y$  direction and negative for counterclockwise rotation.

# Polarization

## High Harmonic Spectroscopy of Multichannel Dynamics in Strong-Field Ionization

Y. Mairesse, J. Higué, N. Dudovich, D. Shafir, B. Fabre, E. Mével, E. Constant, S. Patchkovskii, Z. Walters, M. Yu. Ivanov, and O. Smirnova, PRL 104, 213601 (2010)

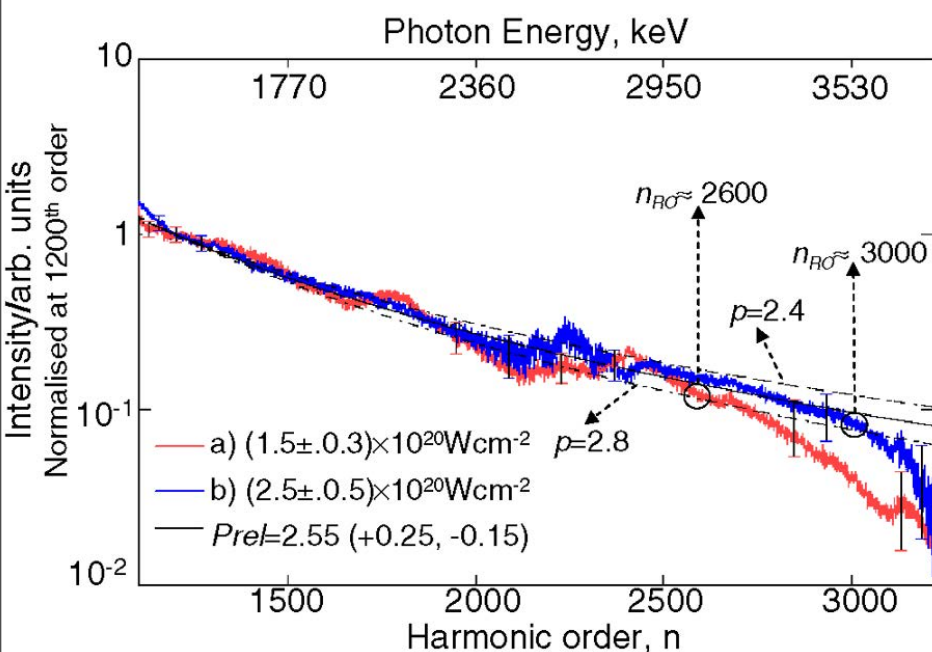


Ellipticity close to 0.4 in  $N_2$

FIG. 2 (color online). Measured harmonic spectra (a),(b) and ellipticity (c),(d) as a function of the molecular alignment angle  $\Theta$  at  $I = 8 \times 10^{13}$  W/cm<sup>2</sup> [(a),(c)] and  $I = 1 \times 10^{14}$  W/cm<sup>2</sup> [(b),(d)].

# High Harmonic Generation on solid target

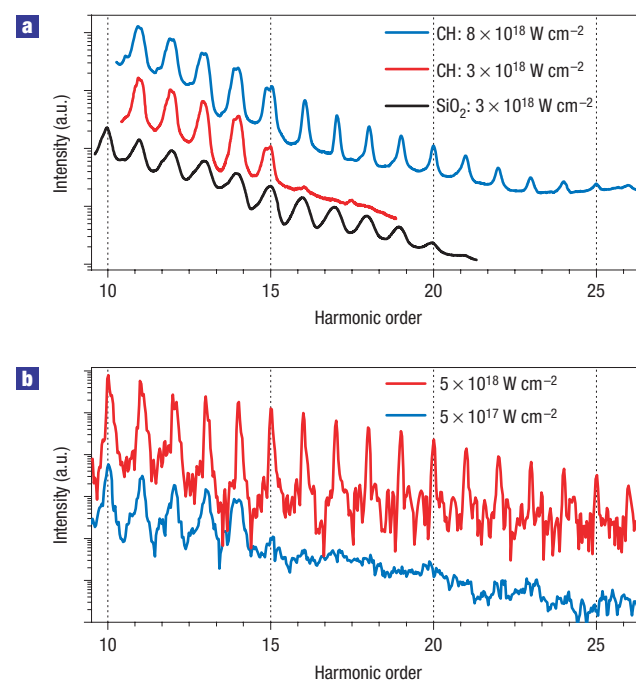
Intense laser pulse interacting with a near discontinuous plasma-vacuum boundary  
 $\Rightarrow$  the laser electric field can efficiently couple to the plasma surface  
 $\Rightarrow$  the electrons oscillate in phase : relativistic mirror oscillating at the laser frequency.  
 Position of this mirror surface : temporal function of the incident optical laser cycle  $\Rightarrow$  the phase of the reflected light wave is modulated  $\Rightarrow$  no longer sinusoidal  $\Rightarrow$  high order harmonics content



3.3 Å (3.8 keV), Harmonic order > 3200

4° cone emission (13° halo)

Vulcan PW laser (600 J, 500 fs)



C. Thaury et al., *Nature Physics* 3, 2007, 424-429

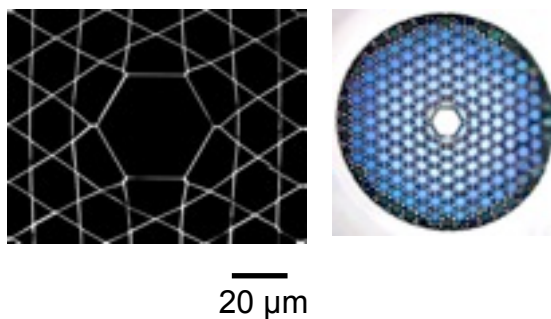
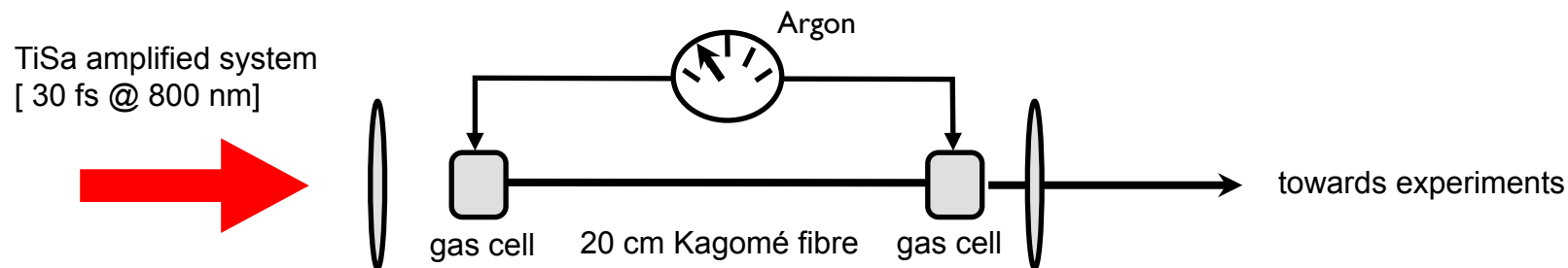
3 μJ @3 keV, B. Dromey. et al, *Phys. Rev. Lett.* 99 (2007) 085001

B. Dromey. et al. *Nature Physics* 2 (2006) 456

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

## Kagome fiber source

Medium : kagome hollow-core photonic crystal fibre filled with noble gas.



Diffraction-limited DUV pulses of  $>50$  nJ, fs-duration, continuously tunable from below 200 nm to above 300 nm

N.Y.Joly, et al., "Bright spatially coherent wavelength tunable deep UV laser source using an Ar-filled photonic crystal fibre," *PRL* 106, 203901 (2011)

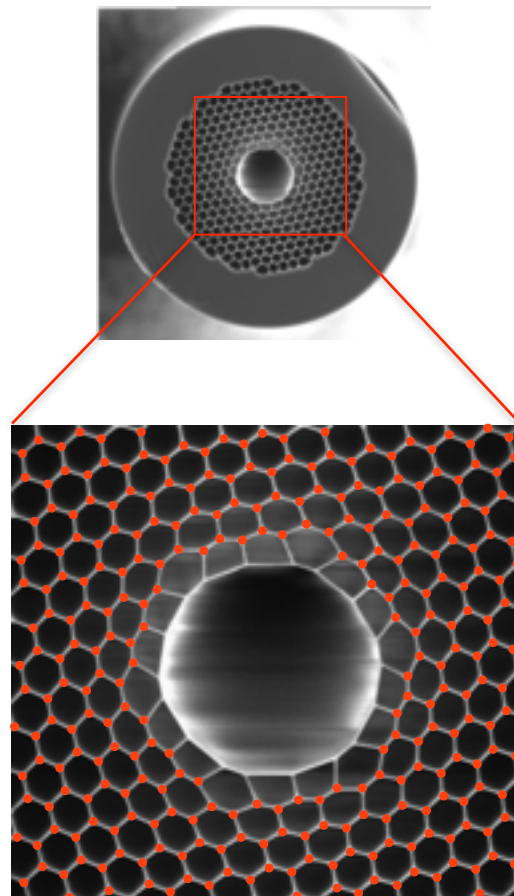
P. Hölzer et al., Femtosecond non linear fiber optics in the ionization regime, *PRL* 107, 203901 (2011)

J.Travers, W. Chang, J. Nold, N.Y.Joly, P. S. J. Russel, Ultrafast nonlinear optics in gas-filled hollow-core photonic crystal fibers, *J. Opt. Soc. Am. B* 28 (12), A11-A26 (2011)

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

## Hollow core family

bandgap fiber

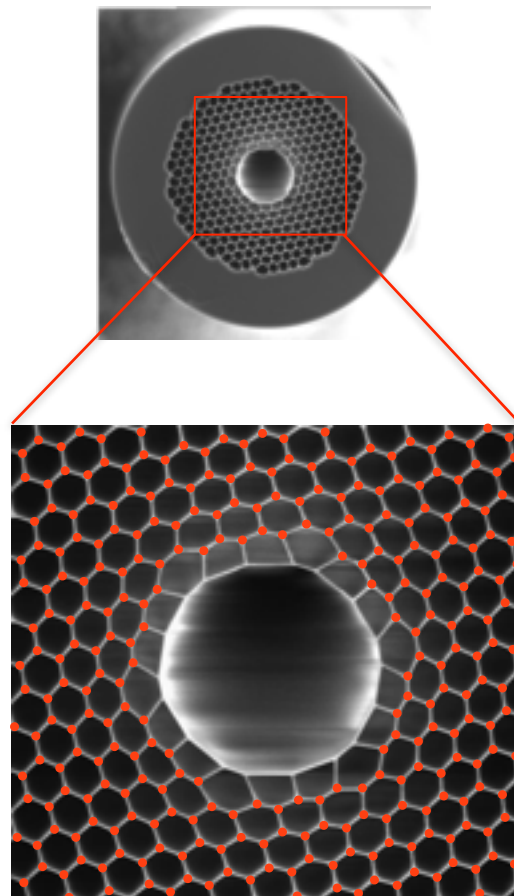


fibre kagomé

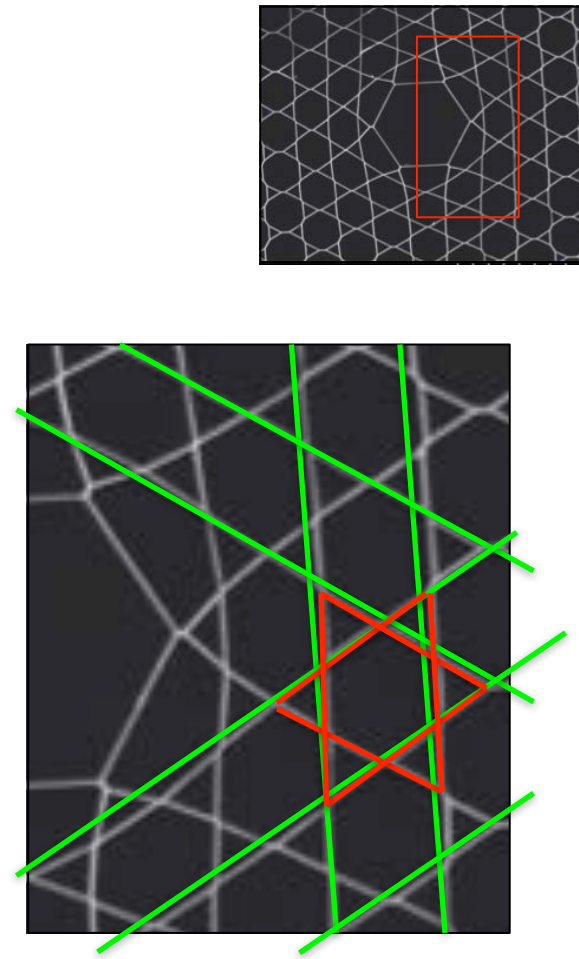


## Hollow core family

bandgap fiber

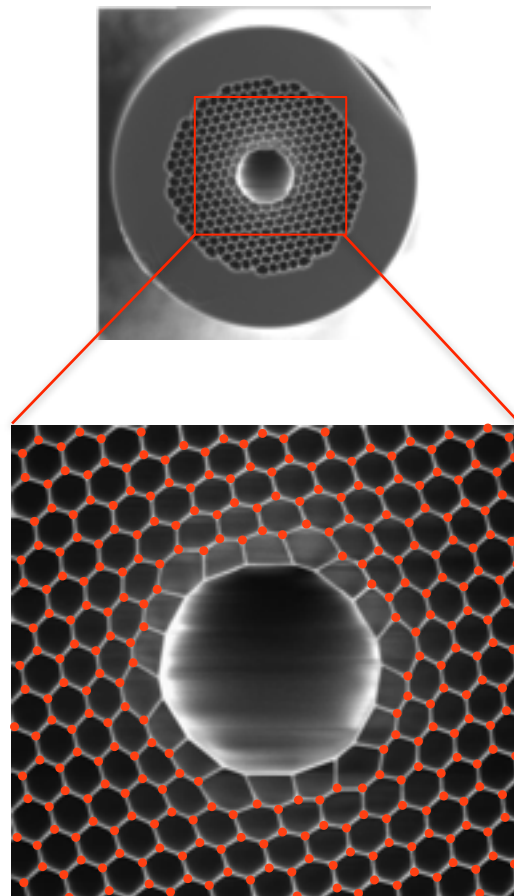


fibre kagomé

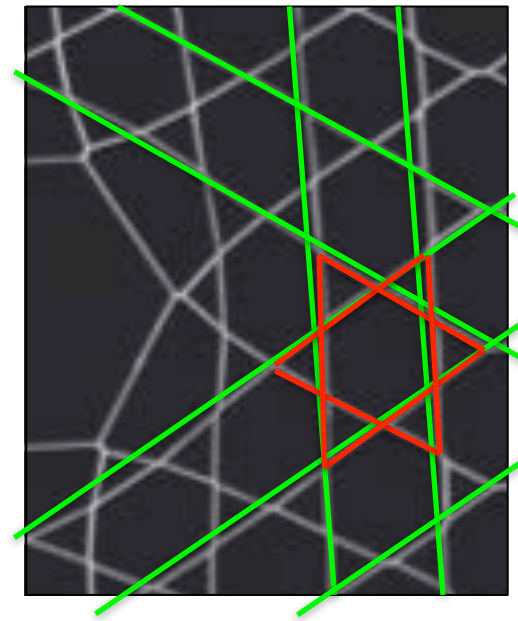
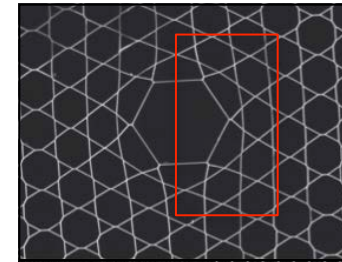


## Hollow core family

bandgap fiber



fibre kagomé

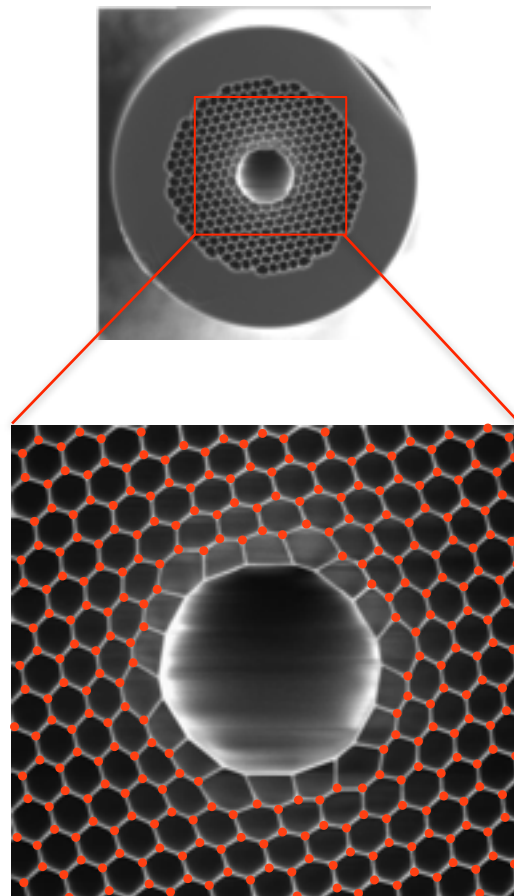


Kagomé basket [かごめ]

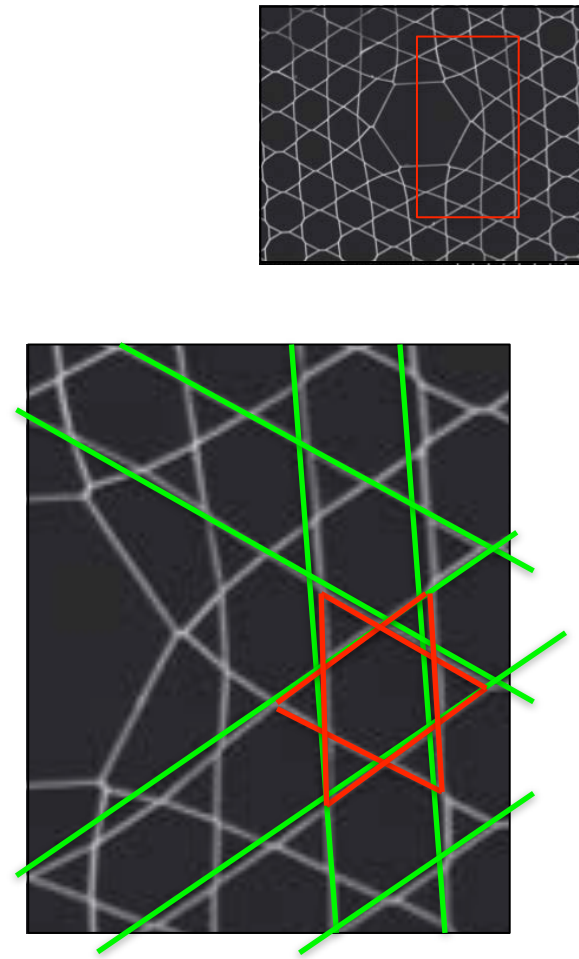


## Hollow core family

bandgap fiber

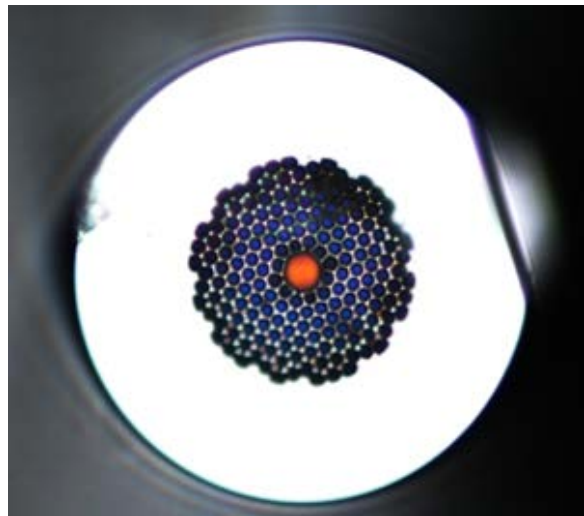
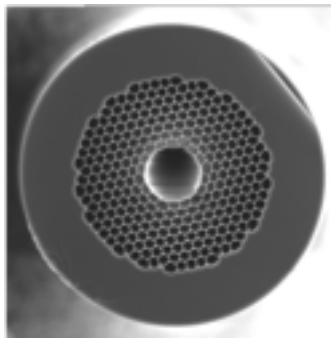


fibre kagomé

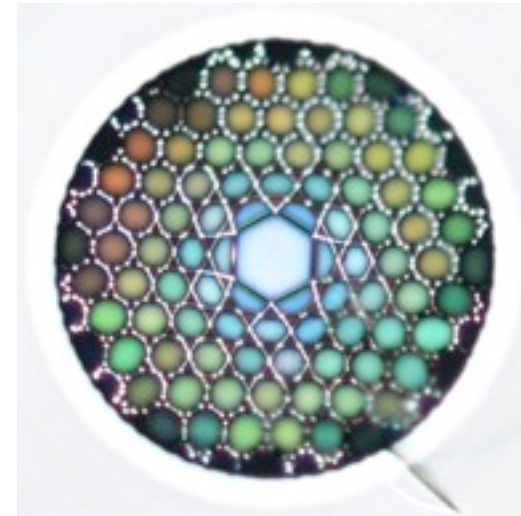


## Hollow core family

bandgap fiber

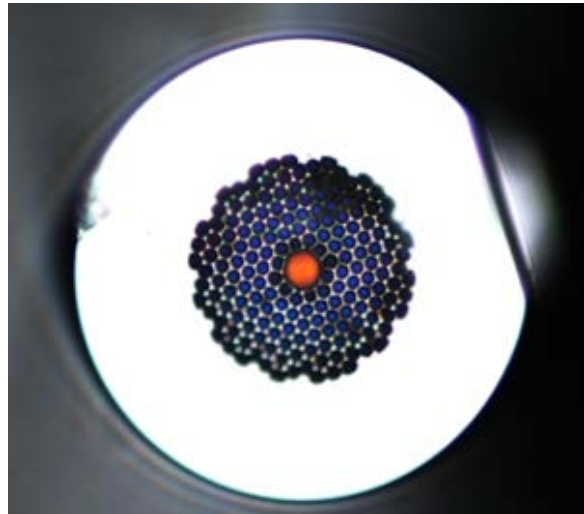
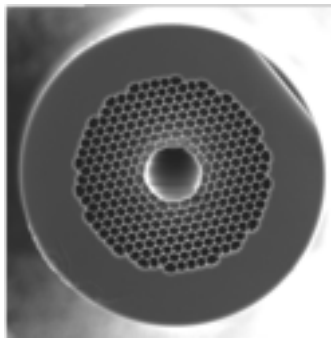


fibre kagomé

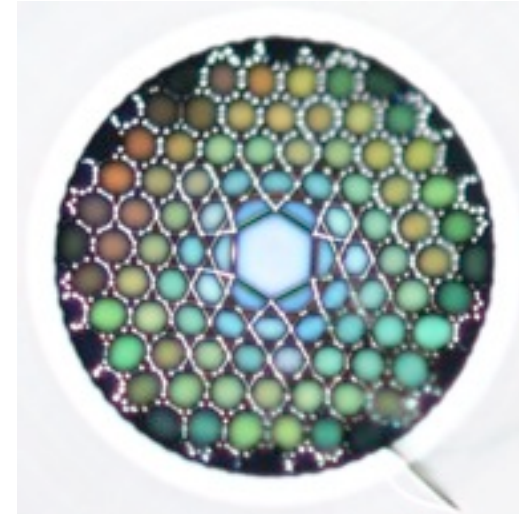
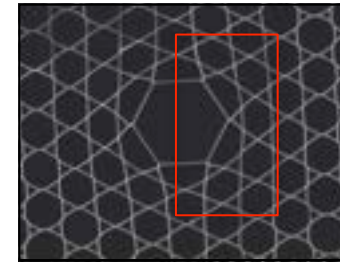


## Hollow core family

bandgap fiber

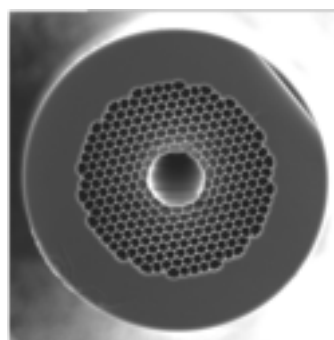


fibre kagomé

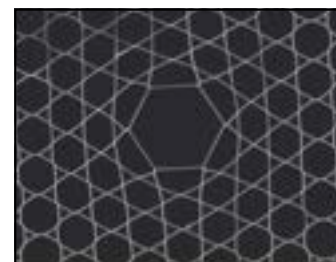


## Hollow core family

bandgap fiber



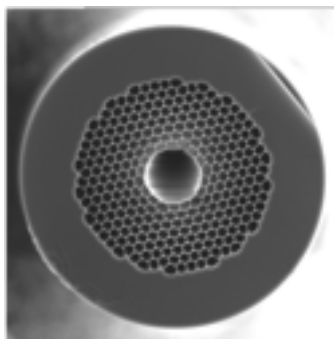
fibre kagomé



In hollow-core photonic crystal fibres (HC-PCF), light propagates in an effectively diffractionless manner, and can then interact efficiently with any material filling the core region

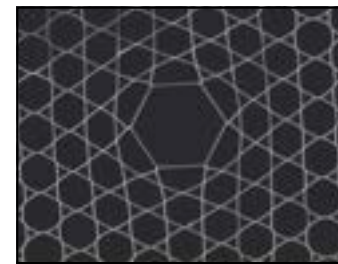
## Hollow core family

bandgap fiber



- narrow transmission band
- low loss (dB/km)
- strong variation of dispersion over the transmission band

fibre kagomé



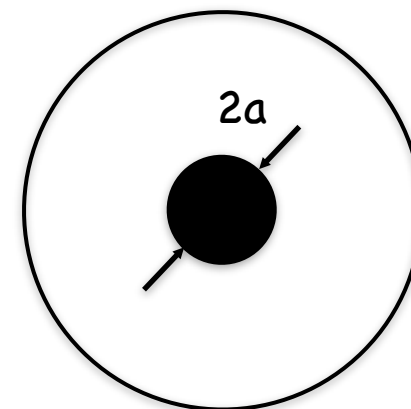
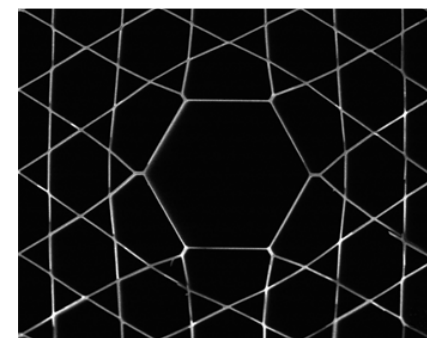
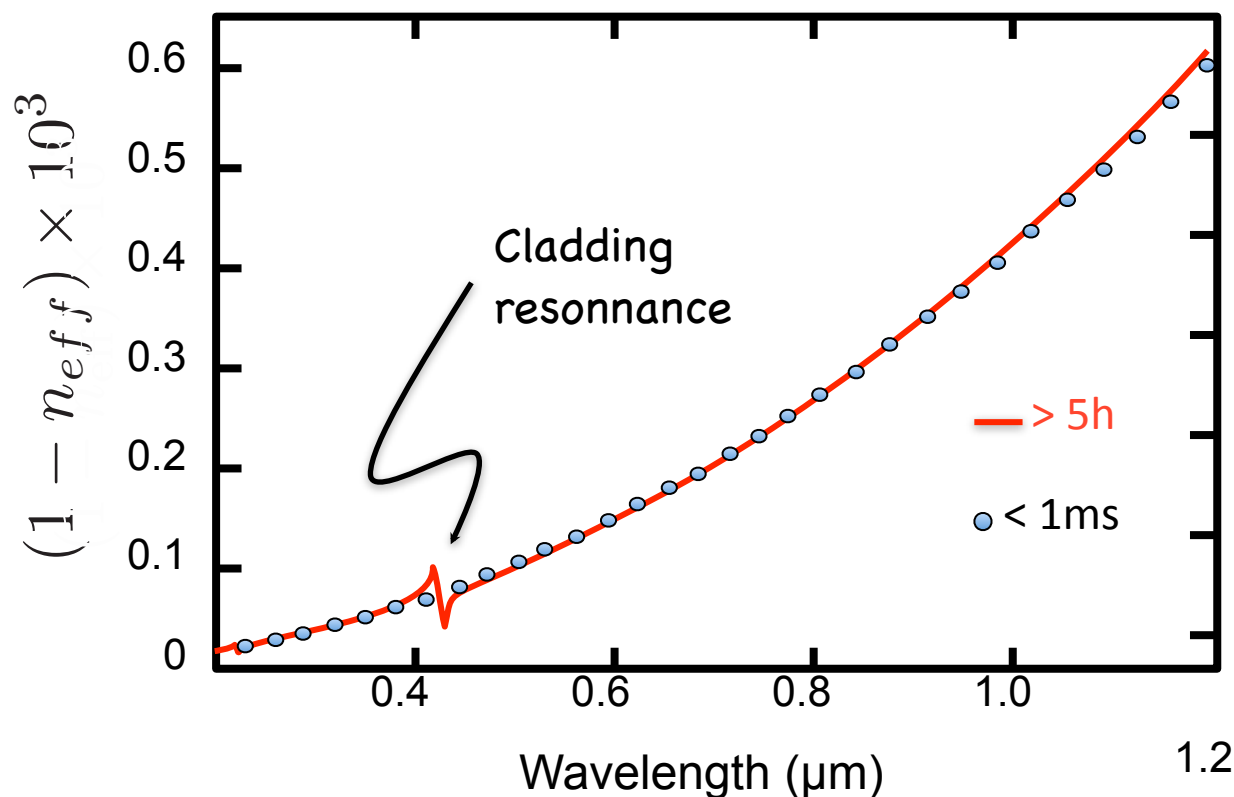
- no bandgap
- large spectral transmission
- high loss (dB/m)
- very low dispersion

In hollow-core photonic crystal fibres (HC-PCF), light propagates in an effectively diffractionless manner, and can then interact efficiently with any material filling the core region

# very broad band transmission associated to ultra-low dispersion

Analytical form for the dispersion

$$n_{mn}(\lambda, P, T) = \sqrt{n_{\text{gas}}^2(\lambda, P, T) - \frac{u_{mn}^2}{k^2 a^2}}$$



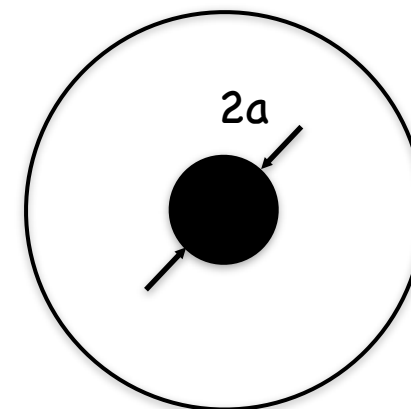
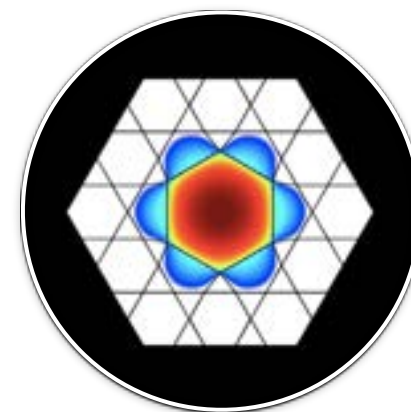
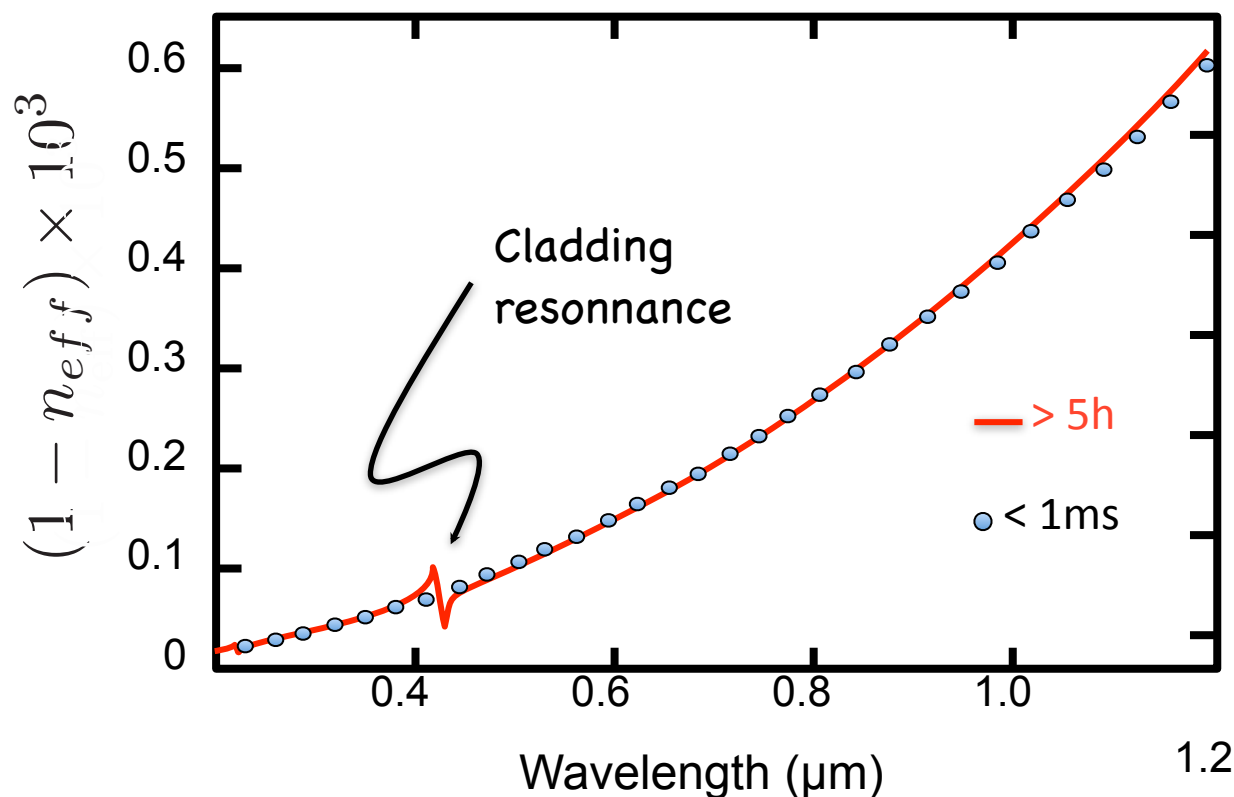
=> nonlinear effects and especially spectral broadening

Modelling : non linear Schrödinger equation including the non linear contribution of the gas (refractive index, effective area of the core)

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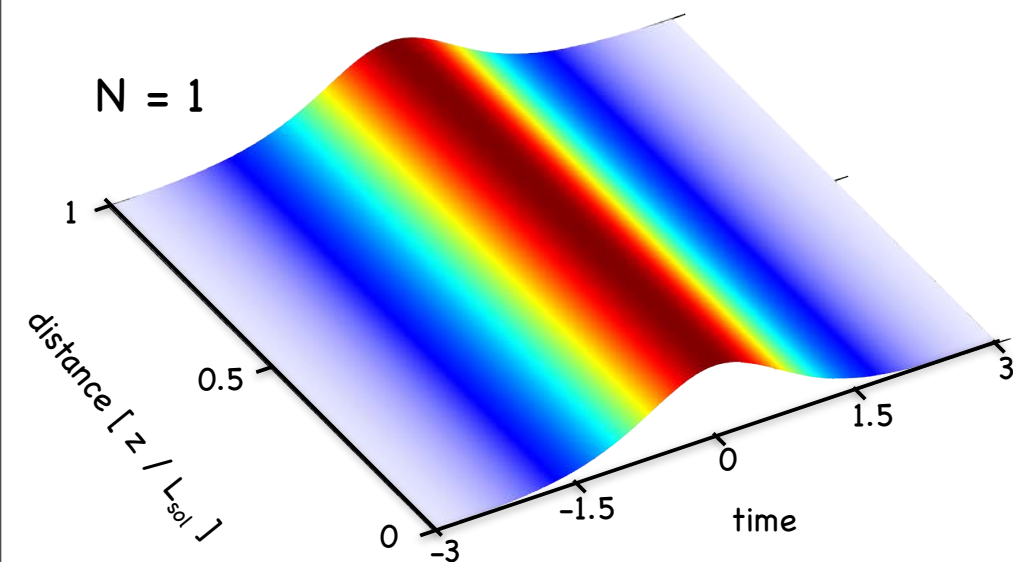


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# Soliton effect

In ideal case (only one term of dispersion + Kerr effect) we can have soliton. Its order is N.



$$N^2 = \frac{L_D}{L_{NL}} = \frac{\gamma P_0 t_0^2}{|\beta_2|}$$

Dispersion length

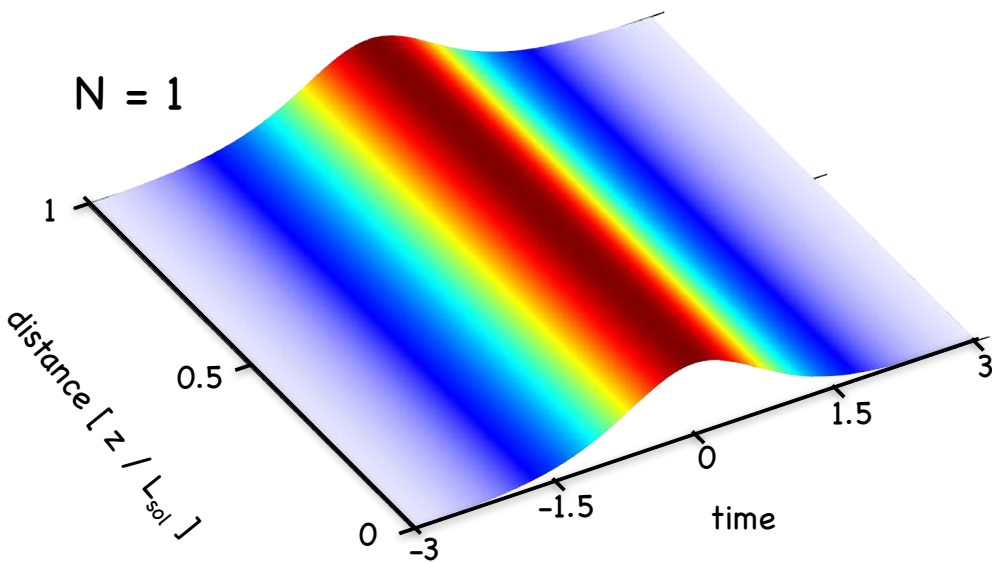
Nonlinear length



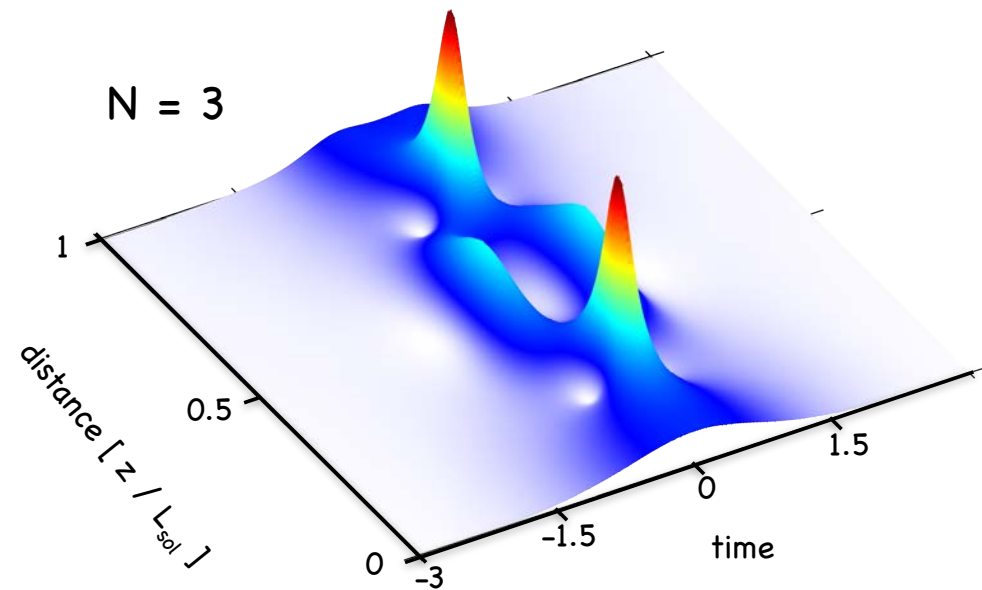
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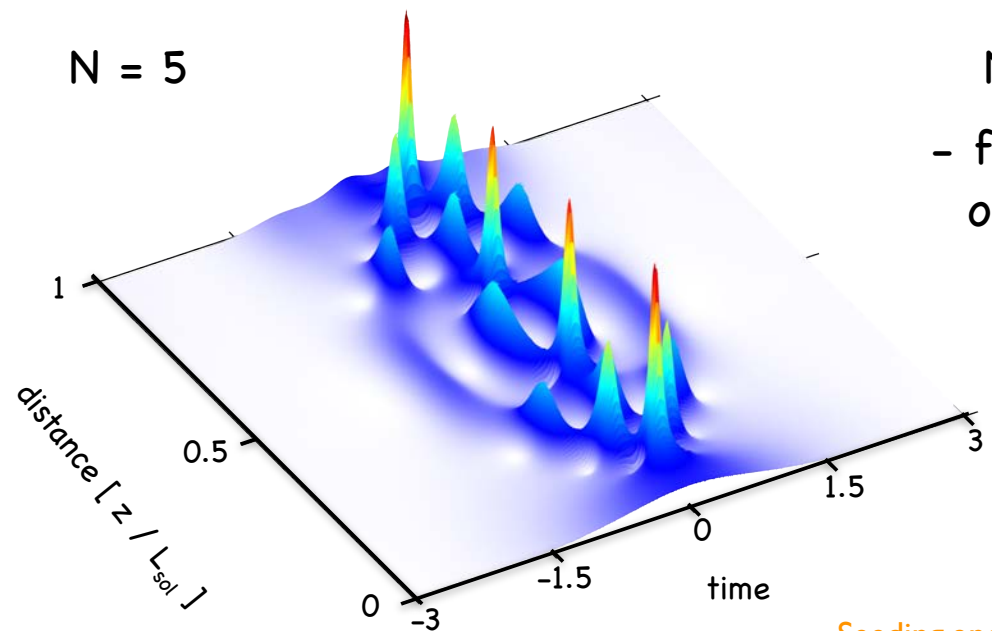
$N = 1$



$N = 3$

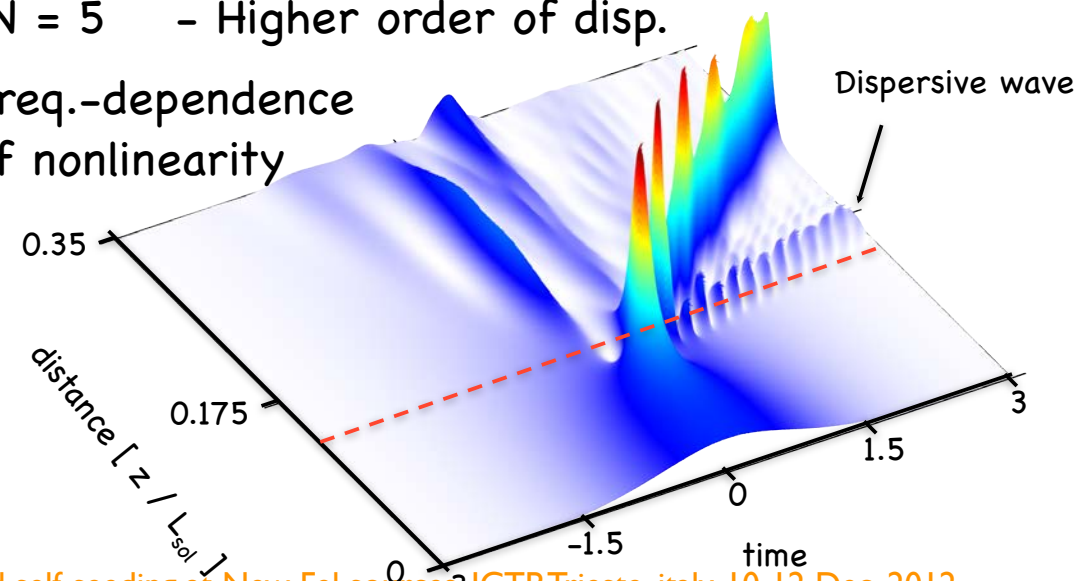


$N = 5$



$N = 5$  - Higher order of disp.

- freq.-dependence  
of nonlinearity

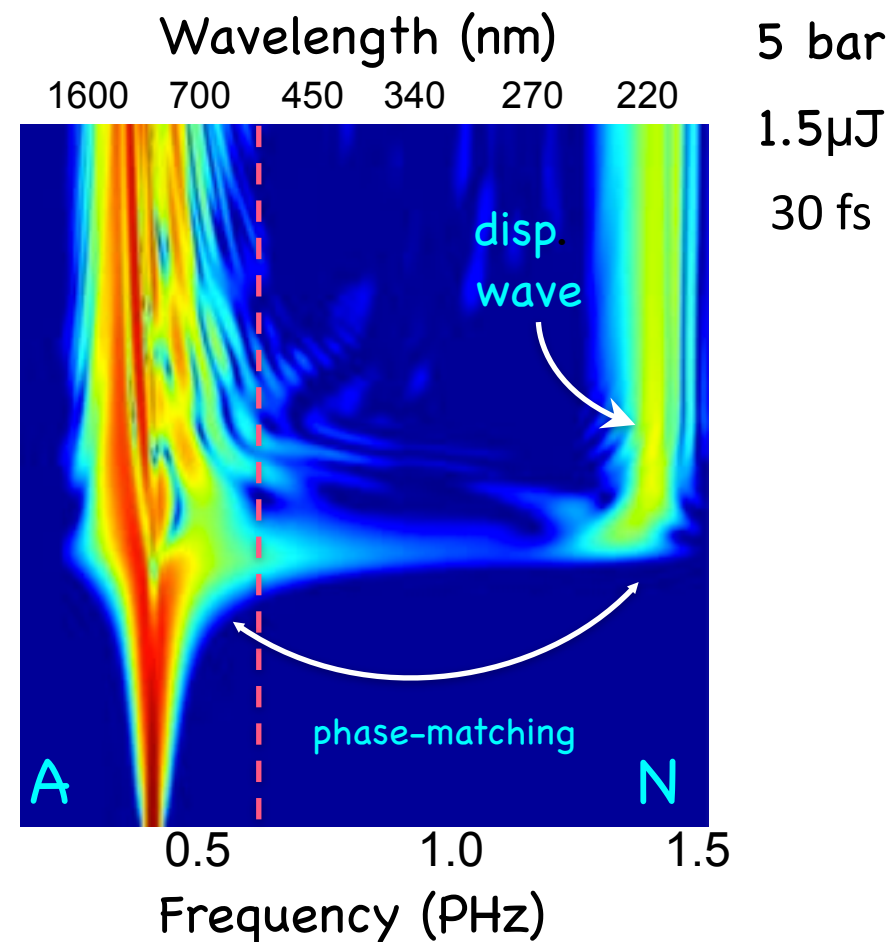


## Dispersion wave emission

When the self compression reaches a few cycles (10 fs), it emits resonant dispersive wave at phase-matched wavelength in the UV spectral range

$$\text{Quality factor of the compression } Q_C \sim \frac{3.7}{N}$$

$$\text{Length to max. compression } L_{\text{fiss.}} = \frac{L_D}{N} = \frac{t_0^2}{N |\beta_2|}$$



Modelling : non linear Schrödinger equation including the non linear contribution of the gas (refractive index, effective area of the core)

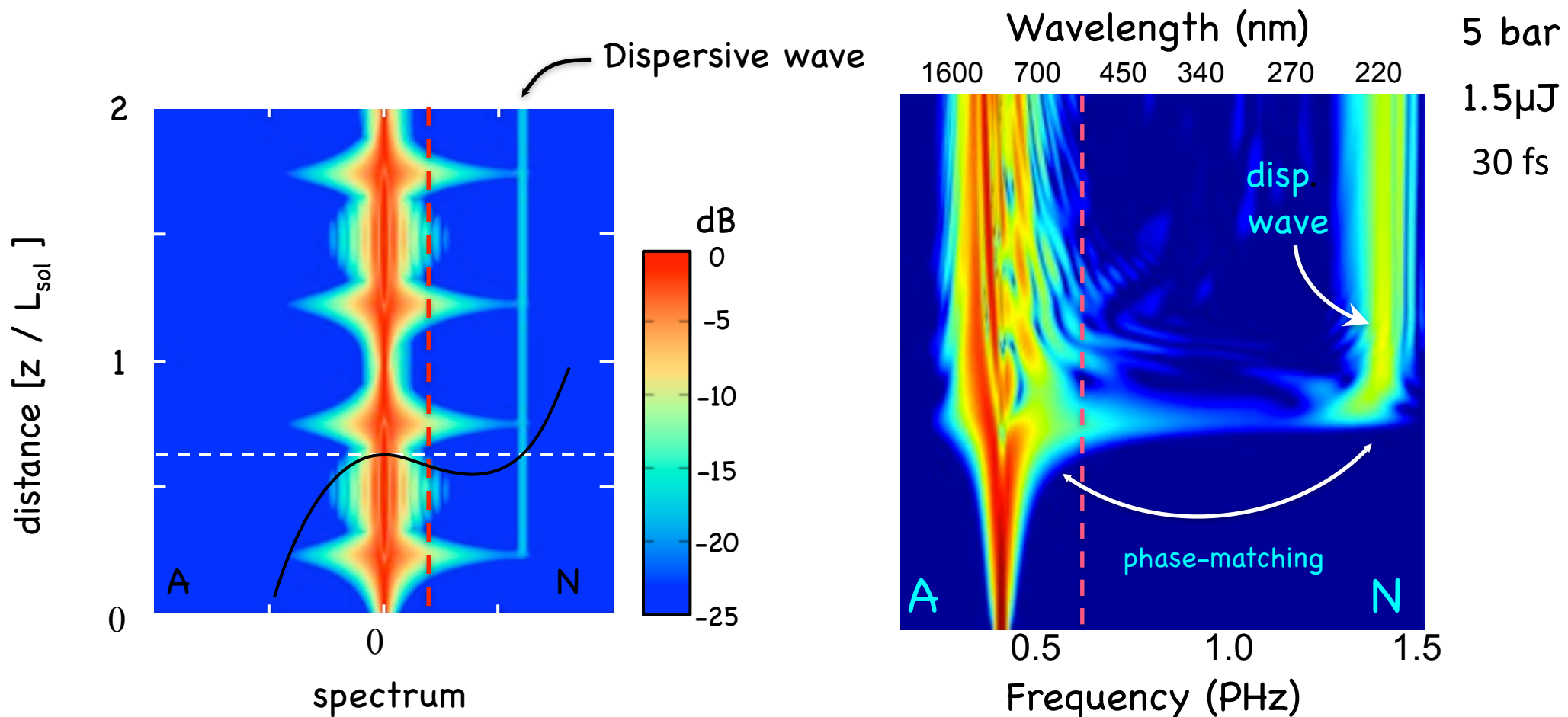
Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

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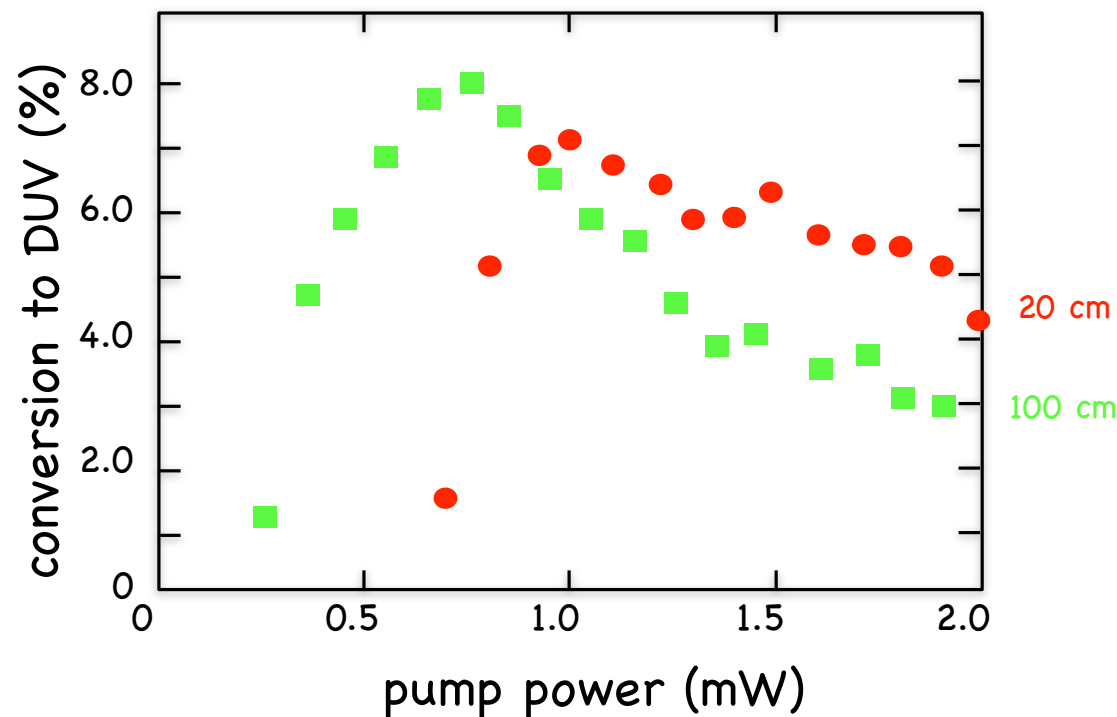
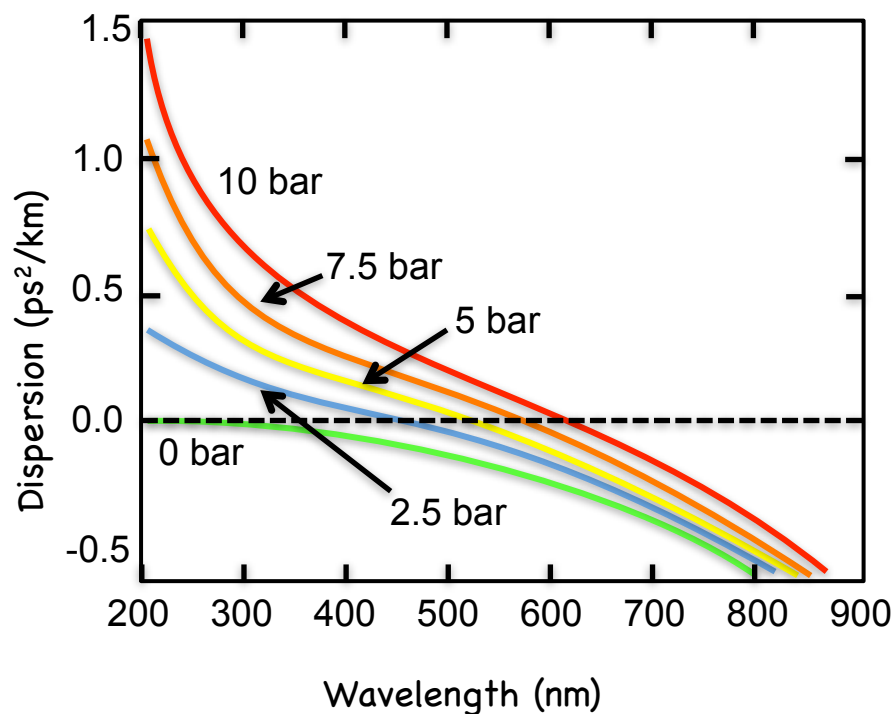


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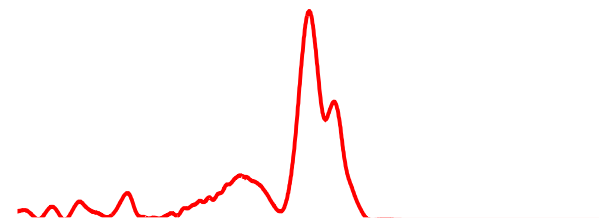
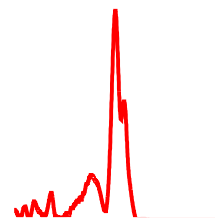
# Kagome fiber source properties

Tunability of the UV  
with the gas pressure



Tunability obtained by controlling both pressure and input power

## Influence of the input pulse duration



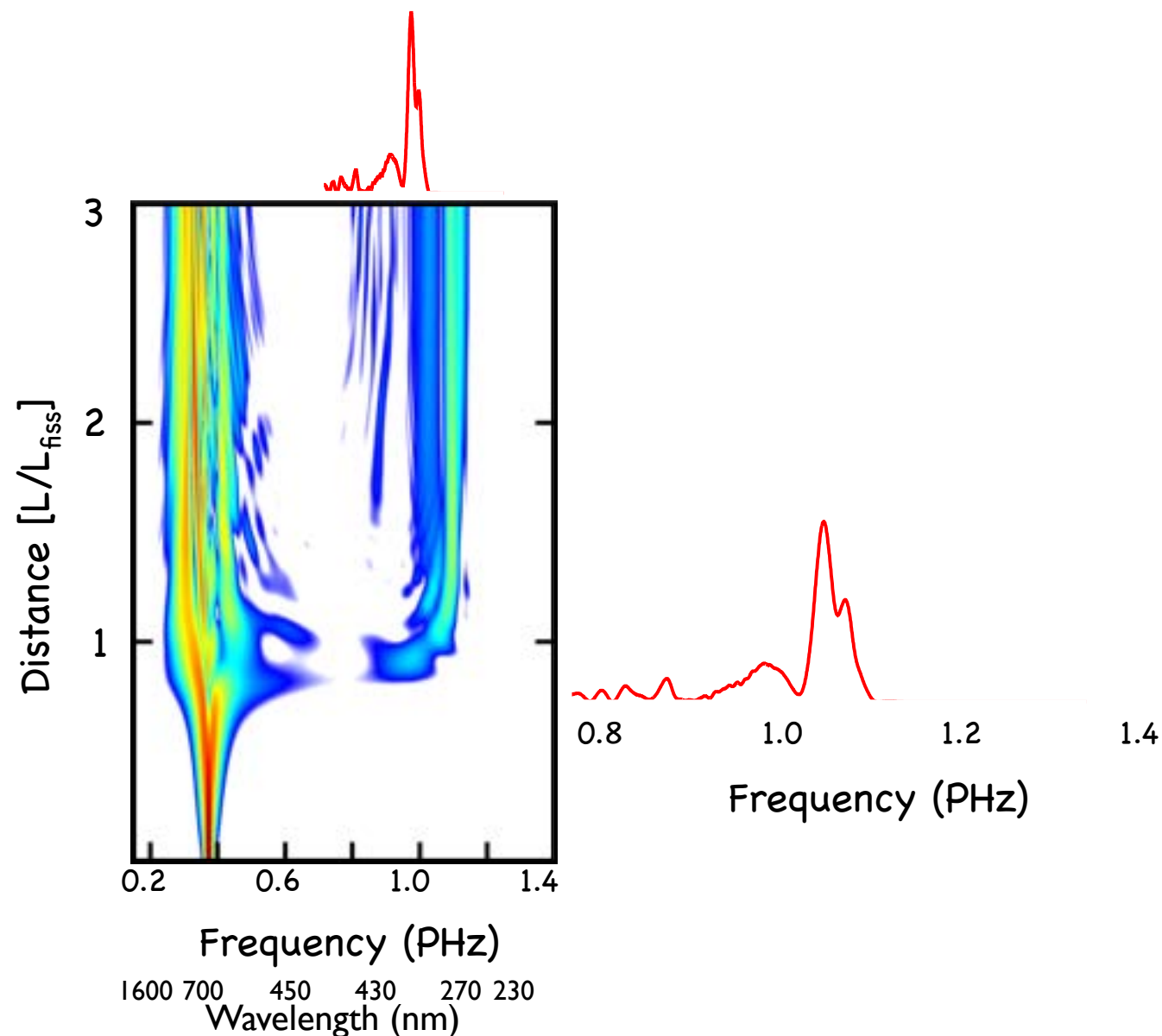
1600 700 450 430 270 230  
Wavelength (nm)

The quality of the compression is inversely proportional to the soliton order  
=>, the quality of the generated band is then strongly affected by N

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

# Influence of the input pulse duration

$$\delta t = 30 \text{ fs}$$

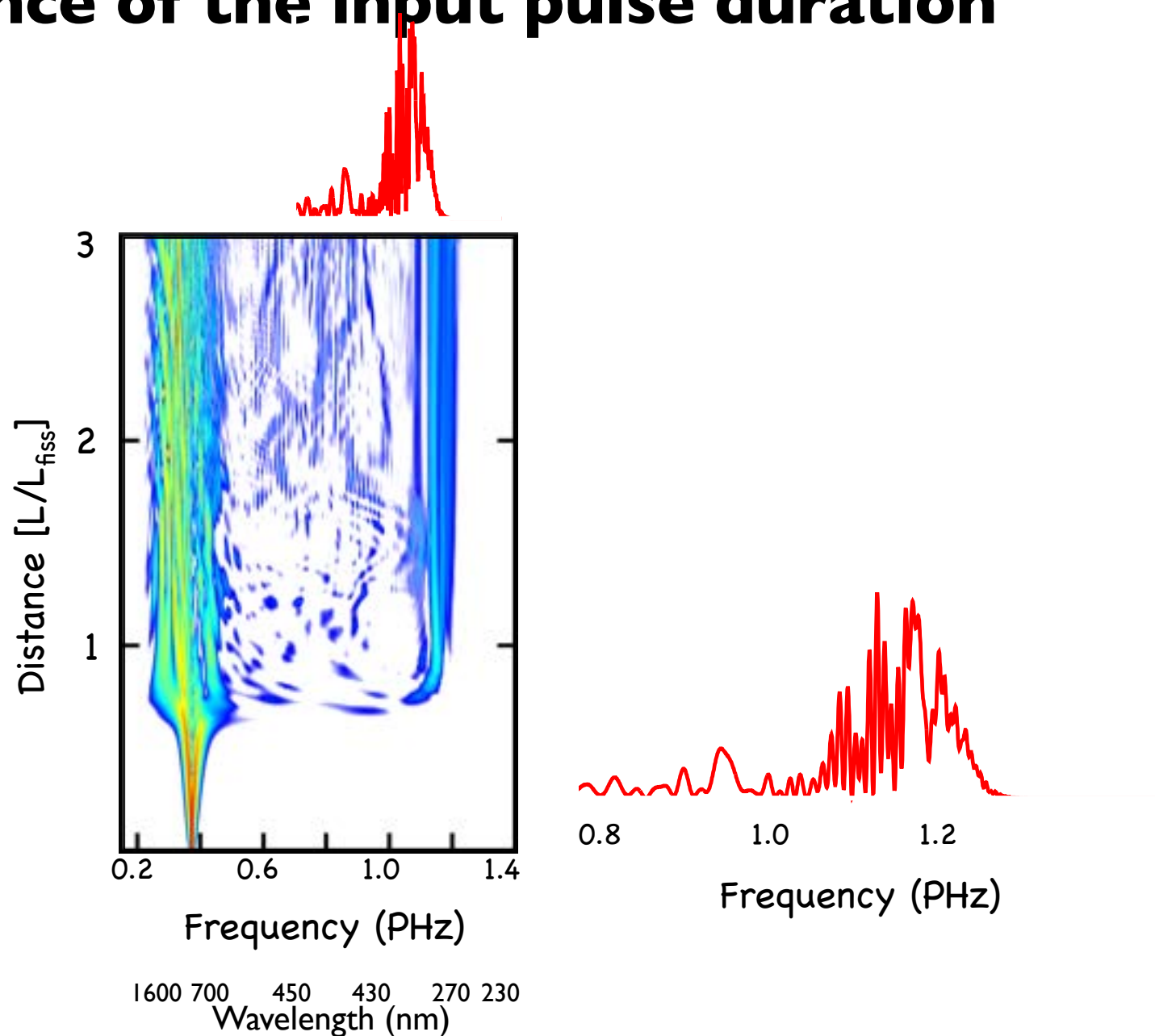


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Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

# Influence of the input pulse duration

$\delta t = 60 \text{ fs}$

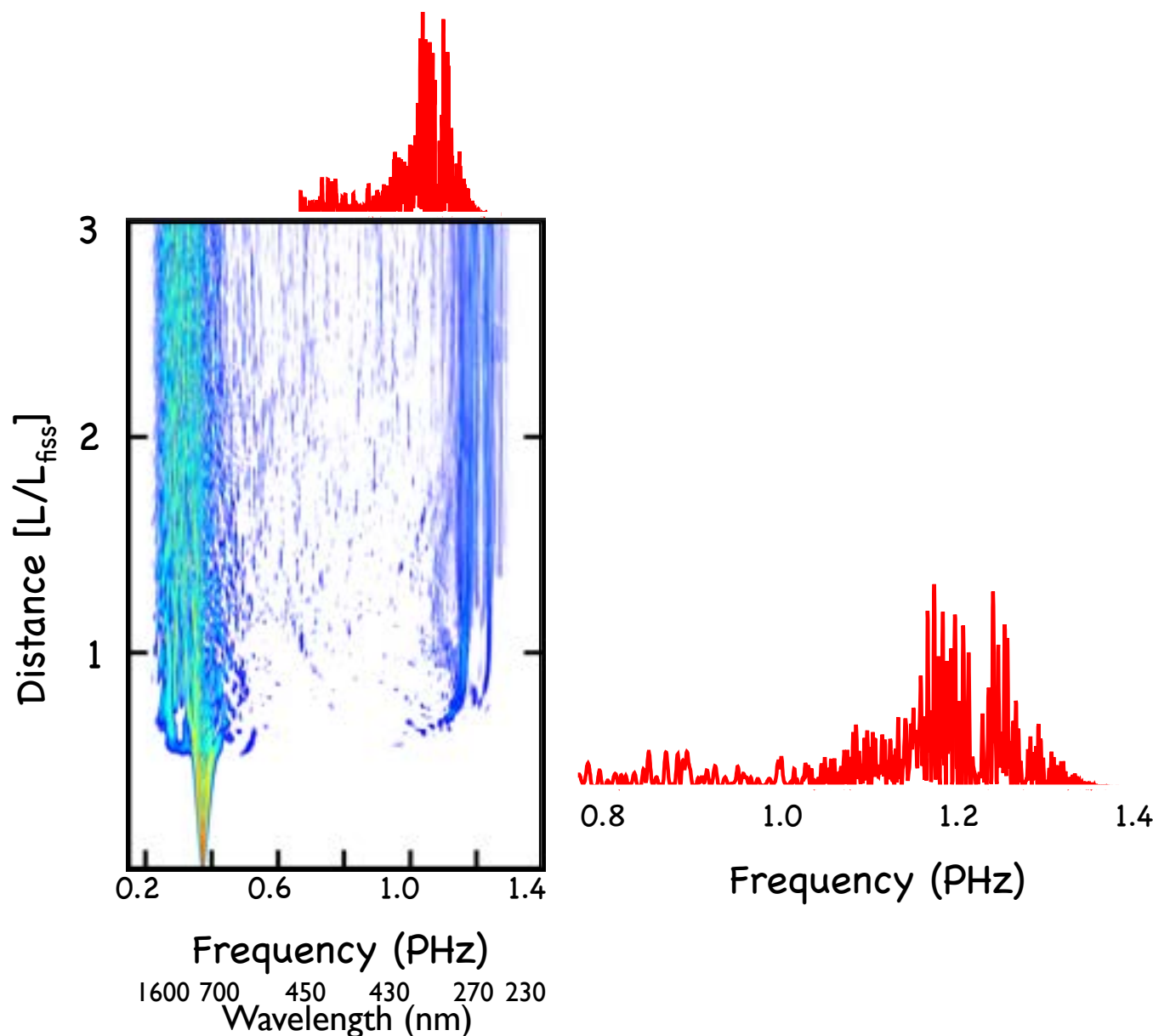


=> the longer the pulse duration, the less clean is the emerging UV band

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

# Influence of the input pulse duration

$\delta t = 120 \text{ fs}$



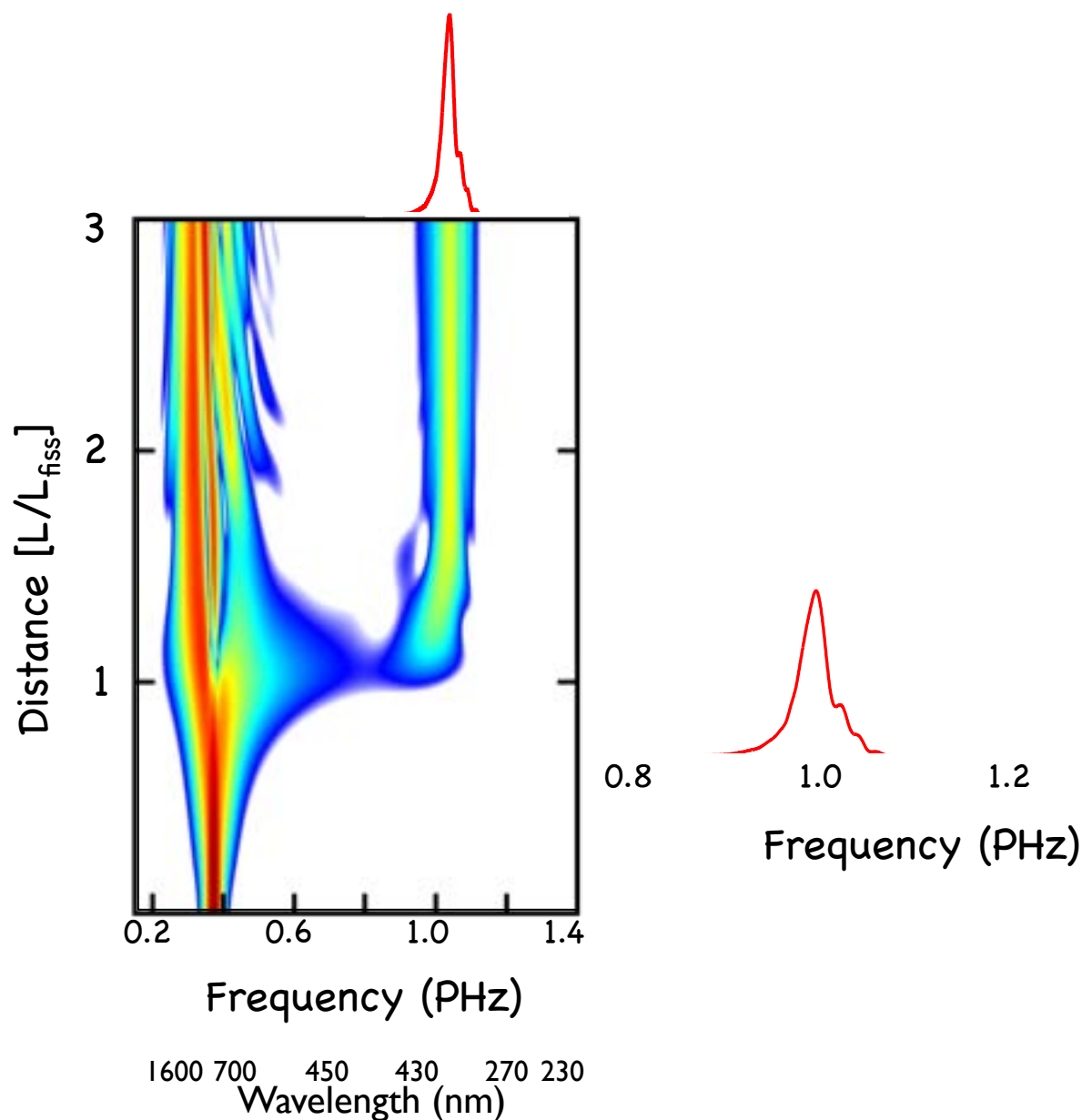
=> the longer the pulse duration, the less clean is the emerging UV band

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012



# Influence of the input pulse duration

$\delta t = 15 \text{ fs}$

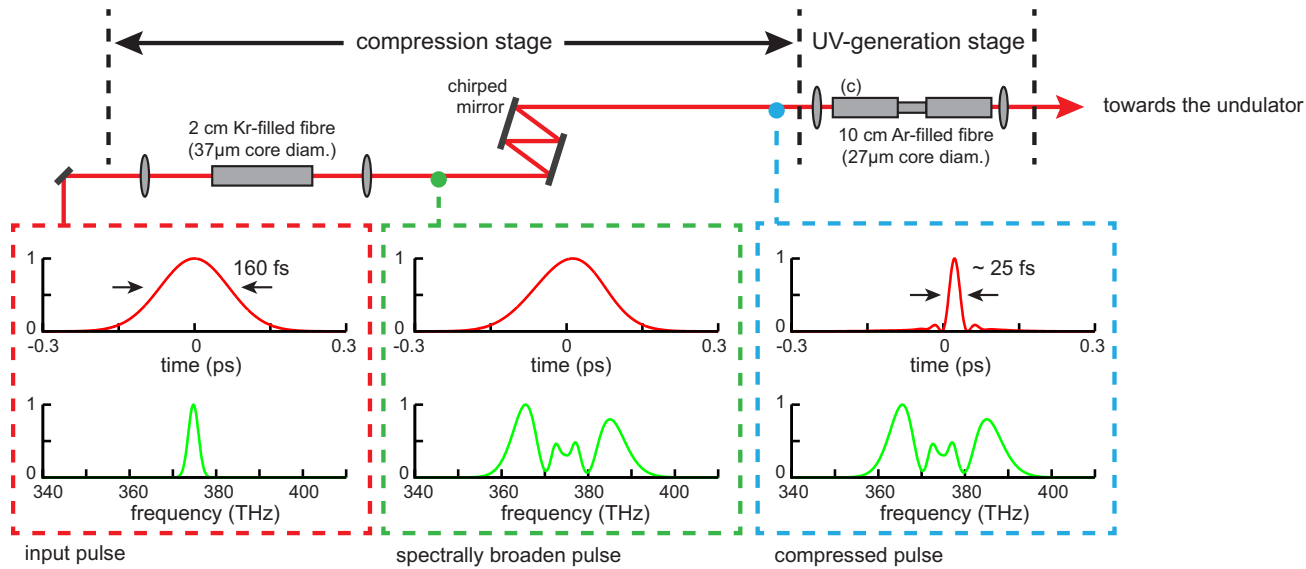


=> Ideally, pulses  $< 60 \text{ fs}$  should be used

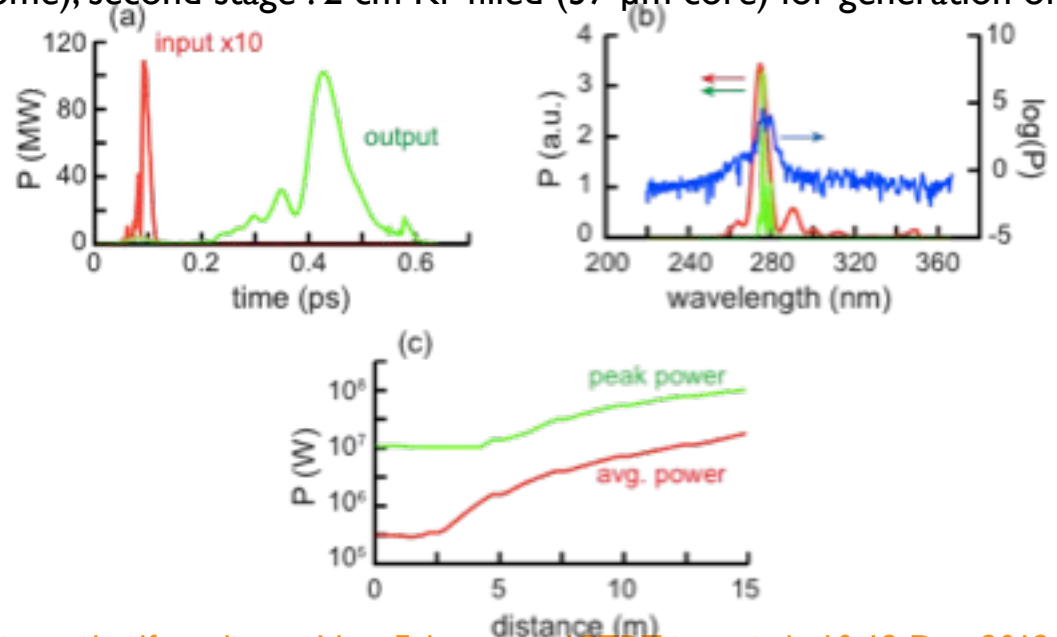
Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012

## Proposed experiment at SPARC

SPARC (Frascati, Italy), 176 MeV,  
 50 A,  $1 \pi$  mm.mrad, 0.01%  
 Kagome fiber UV at 275 nm, 10  
 nJ, 10 fs, 1 MW  
 => FEL at 280 nm



First stage: compress the pulse (10 cm Ar-filled Kagome), second stage : 2 cm Kr filled (37 μm core) for generation of the UV.



Seeding of SPARC-FEL with a Tunable Fibre-based source, N. Joly et al, TUPD17, Proceedings FEL conf, Nara, Japan, 2012

# Conclusion

Seeding : control of FEL performances via the seed (pulse duration... attosecond? ..., stability...)

Compactness

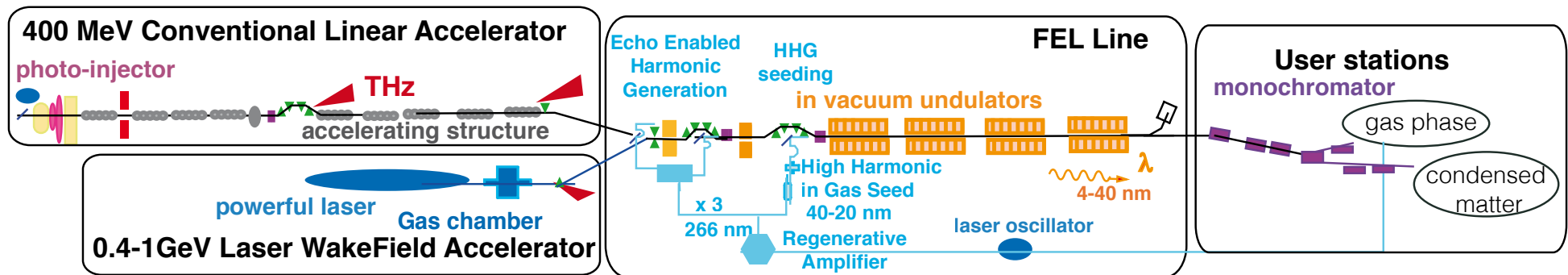
Use of different seeds (short wavelength, tuneable Kagome-fiber) with various frequency up-conversion schemes => short wavelengt FEL

Combination of seeds of different type (echo, TMC...)

Complementary of the seeding source and the FEL on the same site

# Combination of different seeds : ex of LUNEX5

free electron Laser Using a New accelerator for the Exploitation of X-ray radiation of 5<sup>th</sup> generation



40-4 nm, 20 fs and shorter

Beyond **third generation** light source (undulator spontaneous emission, partial transverse coherence),

progress towards **advanced fourth generation (4G+)** light sources (coherent emission, temporal and transverse coherence, femtoseconde pulses, high brilliance) via the **latest free electron laser seeding schemes and electron photon interaction**, to be validated by **pilot user experiments**,

=> *Demonstration of echo at short wavelength*

=> *FEL physics*

=> *Advanced design of FEL source for improved performances, associated with cost and size reduction*

and towards **fifth generation (5G)** (Conventional Linac replaced by a LWFA), FEL being viewed as a qualifying LWFA application : *evaluation of the LWFA performances in «operation-like» conditions (cf EuRRONAc objectives)*

**Complementarity CLA / LWFA :**

CLA high repetition rate, high reliability, LWFA : ultra-short electron bunch, compacity

Seeding and self-seeding at New Fel sources, ICTP, Trieste, Italy, 10-12 Dec. 2012